

Near-Infrared Spectroscopy of Galaxies During Reionization: Measuring C III] in a Galaxy at z = 7.5

Taylor A. Hutchison,^{1,2,*} Casey Papovich,^{1,2} Steven L. Finkelstein,³ Mark Dickinson,⁴ Intae Jung,³ Adi Zitrin,⁵ Richard Ellis,⁶ Sangeeta Malhotra,^{7,8} James Rhoads,^{7,8} Guido Roberts-Borsani,⁶ Mimi Song,^{9,†} and Vithal Tilvi⁸

EoRでの銀河からの電離光子の寄与の評価が重要

でも、このころの銀河の物理的性質はよくわかっていない。暗いrest-UVしか分光できない => LaAくらい

LyAはIGMの吸収/resonant scatterをうける

=> CIII]1907, [CIII]1909に注目

- LyAに次いで強い
- z~2ではlower-Zで等価幅大
- zが正確にわかる

CANDELS z7_GND_42912 (Z=7.5056 LAE)のMOSFIRE H-band分光

- H=25.38AB
- 0.7" slit
- ~10hr total exposure in 2014, 2016 and 2017
- CIII]1907 or 1909が検出
- SiIII]1883, 1892は検出されず

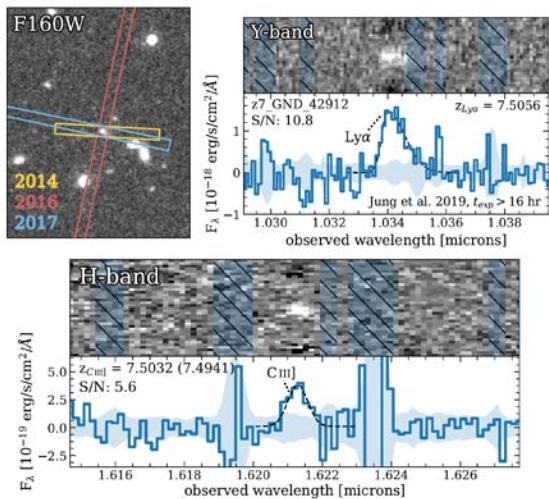


Figure 2. The left: A zoom in of the CANDELS/GOODS-N field in F160W band, centered on z7_GND_42912. The different-colored slits

受かったのはCIII]1907 or 1909?
=> Z>6でΔV_LyA>350km/sなのは M_UV>22magのものだけ
=> おそらく[CIII]1907だろう (ΔV_LyA=88km/s)

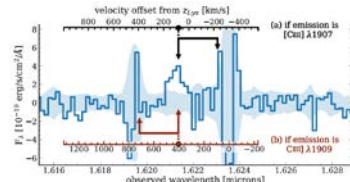


Figure 4. Systemic redshift determination for z7_GND_42912 for the case that the detected emission line is (a) [C III] λ1907 or (b) [C III] λ1909

- [CIII]1907=1.5 CIII]1909を仮定
 - [CIII]1907+CIII]1909/LyA が大きな値になっている。
 - IRAC[3.6]-[4.5]color
 - Strong [OIII]5007
 - 等価幅600A
 - CLOUDYシミュレーション
 - BPASS
 - Starburst99
 - AGNモデル
- => BPASSでないと再現できない(特にIRACcolor)
=> ionization parameter logU>-2

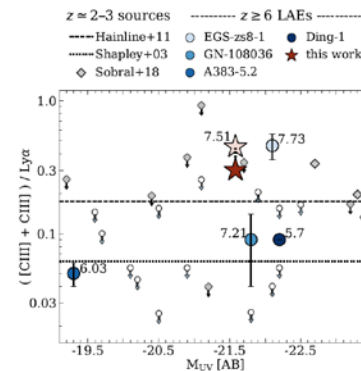


Figure 6. All measured z >= 6 galaxies with spectroscopic measure-

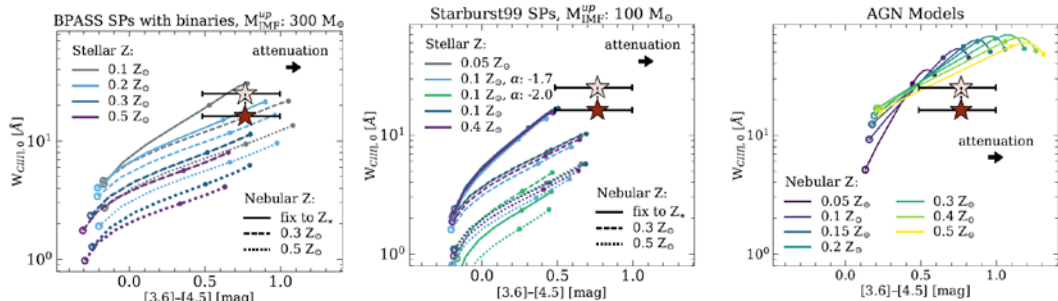


Figure 7. Restframe equivalent with of the C III] doublet versus IRAC color. The large stars represent z7_GND_42912 where the dark red (light

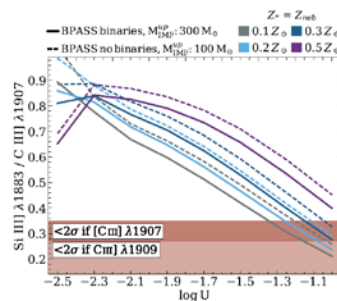


Figure 8. The ratio of Si III] λ1883 / [C III] λ1907 versus ionization

- SiIII]のnon-detection
- Ionization parameterはさらに大きくなる: logU>-1.5