# B4 Seminar #1

Nao Sakai

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### Improving galaxy morphologies for SDSS with Deep Learning

#### H. Domínguez Sánchez,<sup>1,2</sup>\* M. Huertas-Company,<sup>1,2,3</sup> M. Bernardi<sup>1</sup>, D. Tuccillo<sup>2,4</sup>

#### and J. L. Fischer<sup>1</sup>

<sup>1</sup>Department of Physics and Astronomy, University of Pennsylvania, 209 South 33rd Street, Philadelphia, PA 19104, USA <sup>2</sup>LERMA, Observatoire de Paris, PSL Research University, CNRS, Sorbonne Universités, UPMC Univ. Paris 06, F-75014 Paris, France <sup>3</sup>University of Paris Denis Diderot, University of Paris Sorbonne Cité (PSC), 75205 Paris Cedex 13, France <sup>4</sup>MINES Paristech, PSL Research University, Centre for Mathematical Morphology, Fontainebleau, France

#### ABSTRACT

We present a morphological catalogue for ~ 670,000 galaxies in the Sloan Digital Sky Survey in two flavours: T-Type, related to the Hubble sequence, and Galaxy Zoo 2 (GZ2 hereafter) classification scheme. By combining accurate existing visual classification catalogues with machine learning, we provide the largest and most accurate morphological catalogue up to date. The classifications are obtained with Deep Learning algorithms using Convolutional Neural Networks (CNNs).

We use two visual classification catalogues, GZ2 and Nair & Abraham (2010), for training CNNs with colour images in order to obtain T-Types and a series of GZ2 type questions (disk/features, edge-on galaxies, bar signature, bulge prominence, roundness and mergers). We also provide an additional probability enabling a separation between pure elliptical (E) from S0, where the T-Type model is not so efficient. For the T-Type, our results show smaller offset and scatter than previous models trained with support vector machines. For the GZ2 type questions, our models have large accuracy (> 97%), precision and recall values (> 90%) when applied to a test sample with the same characteristics as the one used for training. The catalogue is publicly released with the paper.

Key words: Galaxies - Morphology - Machine learning

## O.Abstruct This Work

a morphological catalogue for ~670,000 galaxies in SDSS

the largest and most accurate morphological catalogue up to date

- better T-Type classification
- separation between pure E from SO

## Method

Deep Learning - Convolutional Neural Network(CNNs)

Data Sets for Training

Galaxy Zoo 2 (GZ2) Nair & Abraham 2010 (N10)

Classification

GZ2

T-Type to the Hubble sequence

## 1.Introduction Motivation

morphology is related to the physical property of the galaxy

accurate morphological classification for large samples

time consuming

not obvious
 i.g. Galaxy Zoo

Deep Learning

Dieleman et al. 2015 (D15)

SDSS DR7 Main Galaxy Sample reproduce the GZ2 **Problem** 

galaxies with uncertain classifications used for training

## This Time

~ improved version ~

galaxies with robust GZ2 classification used for training simplify the galaxy decision tree complement the GZ2 classification scheme with a T-Type

2.Data Sets for Training GZ2 T-Type classification N10 (visual classifications) separate pure E from SO for Testing GZ2 N10 (visual classifications) Huertas-Company et al. 2011 T-Type ETGs and spiral galaxies Cheng et al. 2011 Parent Sample of the Catalogue Presented in This Work Meert et al. 2015, 2016 ~670,000 galaxies SDSS DR7



figure.1 network architecture

GZ2: binary classification mode T-Type: regression mode

# 4.GZ2 based models Questions binary classification mode



figure.3 decision tree

# 4.GZ2 based models Certain Galaxies for Training only use certain galaxies for training

Question	Meaning	Nvotes	Ncertain	Npos
Q1	Disk/Features	239728 (99%)	134475 (56%)	28513 (21%)
Q2	Edge-on disk	151560 (63%)	123201 (81%)	17631 (14%)
Q3	Bar sign	11 <b>7262 (48%)</b>	76746 (65%)	6595 (8%)
Q4	Bulge prominence	117245 (49%)	49345 (42%)	27185~(55%)
$\mathbf{Q5}$	Cigar shape	180223 (75%)	124610 (70%)	28230~(23%)
$Q_6$	Merger signature	239669 (99%)	110079 (46%)	1399 (1%)

table.1 questions in each tier

#### P > 0.8 or P < 0.2

# 4.GZ2 based models Test

Question Q1	Meaning Disk/Features	$P_{thr} \\ 0.2 \\ 0.5 \\ 0.8$	TPR 0.97 0.95 0.90	Prec. 0.91 0.96 0.99	Acc. 0.98
Q2	Edge-on	0.2 0.5 0.8	1.00 0.99 0.92	0.67 0.83 0.95	0.97
Q3	Bar sign	0.2 0.5 0.8	0.93 0.79 0.58	0.48 0.80 0.92	0.97
Q6	Merger signature	0.2 0.5 0.8	0.98 0.96 0.90	0.54 0.82 0.97	0.97

$$Acc = \frac{TP + TN}{(P + N)}$$

$$Prec = \frac{TP}{TP + FP}; \ R = \frac{TP}{TP + FN} = TPR$$

TP: true positive FP: false positive FN: false negative

Pthr: threshold

user can optimize this value

#### table.2 precision and TPR value for different $P_{thr}$



figure.10 probability distribution obtained by applying the model to a sample well classified

# 5.N10 based models

N10 very detailed morphological catalogue

## T-Type regression mode

- T-Type < 0 ETGs
- T-Type = 0 S0
- T-Type > 0 spiral galaxies
- T-Type = 10 irregular galaxies



figure.13 T-Type questions scheme

# 5.N10 based models E versus S0

apply 681 N10 galaxies not used for training



training N = 4000 positive -3 <= T-Type <= 0 negative T-Tye = -5

figure.14 probability distribution of being SO rather than E

better performance than Cheng et al. 2011

## 5.N10 based models **Barred Galaxies** alternative model to the GZ2

training N = 7000

the model trained with GZ2 is worst for barred galaxies

only exist (strong) or not



P bar

apply 1595 unbarred 314 barred galaxies not used for training

strong, intermediate, week

figure.15 probability distribution of having bar signature

P bar

z

# 6.Details Content

Col.	Name	Meaning	Train sample
1	dr7objid	SDSS ID	
<b>2</b>	$\operatorname{galcount}$	Meert15 ID	
3	$P_{disk}$	Prob. features/disk	$\mathbf{GZ2}$
4	P <sub>edge-on</sub>	Prob. edge on	$\mathbf{GZ2}$
5	$P_{bar-GZ2}$	Prob. bar signature	GZ2
6	$P_{bar-N10}$	Prob. bar signature	N10
7	P <sub>merg</sub>	Prob. merger	GZ2
8	P <sub>bulge</sub>	Prob. bulge prominence	$\mathbf{GZ2}$
9	$P_{cigar}$	Prob. cigar shaped	$\mathbf{GZ2}$
10	T-Type	T-Type	N10
11	$P_{S0}$	Prob. S0 vs $E$	N10

## 6.Details Unambiguous Classification

#### successful

Pdominant + Pobvious



#### different approach

figure.17 probability distribution

## 6.Details Correlation with Other Morphological Parameters



figure.18 mean probability values

# 6.Details Visual Inspection



figure.20 edge on galaxies certain in this work though uncertain in GZ2

## no "true reference" catalogue

# 7.Summaries Improve D15

independently train each question from the GZ2 scheme use only certain galaxies for training binary classification mode

670,722 galaxies large accuracy unambiguous classification (disk/features)

## Complement the GZ2 Type Classification

T-Type ~50 times larger than N10 large accuracy separation between E from S0 large accuracy bar classification large accuracy

## Forthcoming Work

apply the models to other SDSS samples to complement the morphological classification catalogue