

SDSS-IV MaNGA: Spatial Evolution of Gas-phase Metallicity Changes Induced by Galaxy Interactions

H.-A. Pan et al. (2025), <https://iopscience.iop.org/article/10.3847/1538-4357/adbbd2>

Introduction

- The gas-phase metallicity of galaxies is strongly related to star formation, feedback mechanisms, and gas inflow and outflow.
- Galaxy mergers can cause gas inflow and metal redistribution.
- Previous studies have shown that interacting galaxies have a flatter radial metallicity gradient than isolated galaxies.
- However, observational evidence for stage-dependent spatial changes is still limited.
- To investigate how gas-phase metallicity evolves spatially across different merger stages using MaNGA data.

Data and Analysis

- Sample:
 - **205** interacting galaxies and **1348** control galaxies from SDSS-IV MaNGA
- Merger stages are visually classified into 4 categories (S1–S4):
 - **S1:** Well-separated pairs that do not show any morphology distortion
 - **S2:** Close pairs showing strong signs of interaction, such as tails and bridges
 - **S3:** Well-separated pairs, but showing weak morphology distortion
 - **S4:** Merging or post-merger (1 or 2 nuclei)
- The gas-phase metallicity is determined using the O3N2 calibration.
- $\Delta O/H$ and $\Delta \log(sSFR)$ are calculated by comparing with control galaxies matched in redshift, stellar mass, and radius.

Results and Discussion

- Galaxies in p/m generally have lower metallicity than isolated control galaxies (Fig3-a)
- Metallicity gradient evolution
 - Flattened metallicity gradients observed in S2 and S4 stages
 - S3 shows a negative metallicity gradient similar to the initial gradient
- Metallicity and sSFR
 - There is an anticorrelation between $\Delta O/H$ and $\Delta \log(sSFR)$: The higher the sSFR, the lower the central oxygen tends to be (Most pronounced in S2 and S4(1)) (Fig3-b, Fig4-b)
 - It suggests a strong relation between the gas inflow triggering star formation and the subsequent metal enrichment
- Metallicity and projected separation
 - The closer galaxy proximities correlate with more significant changes in metallicity, particularly in central regions (Fig5)
 - However, projected separation alone does not fully reflect merger stage

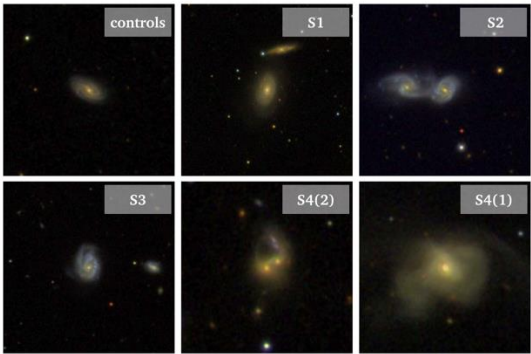


Figure 1

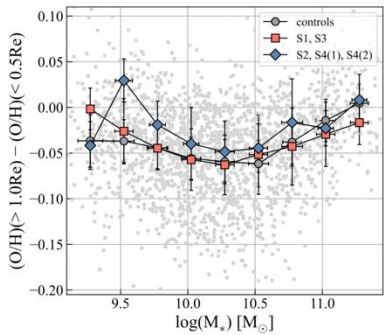


Figure 2

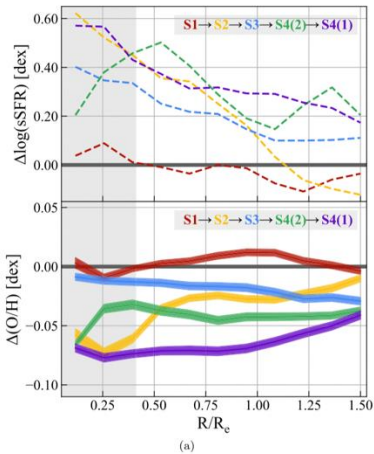


Figure 4

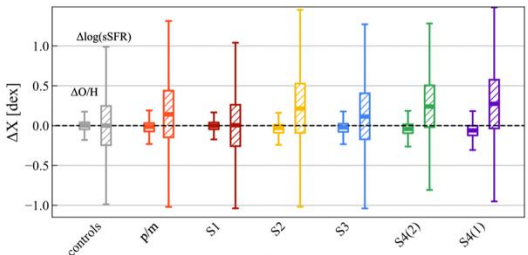
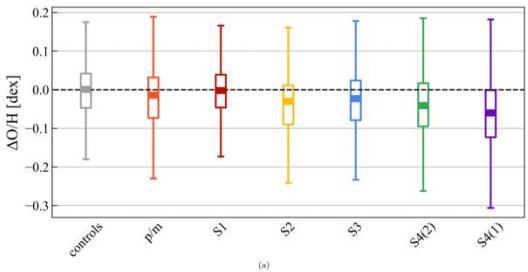
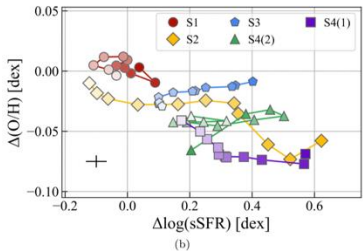


Figure 3

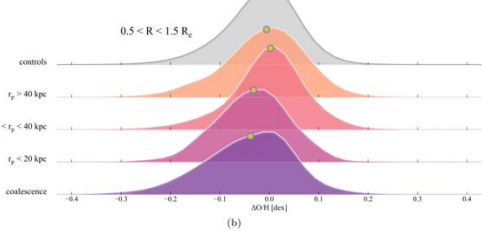
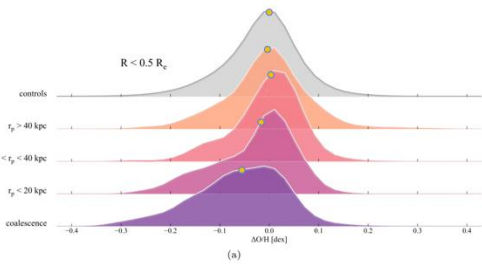


Figure 5