



INSTITUTO DE ASTROFÍSICA
FACULTAD DE FÍSICA



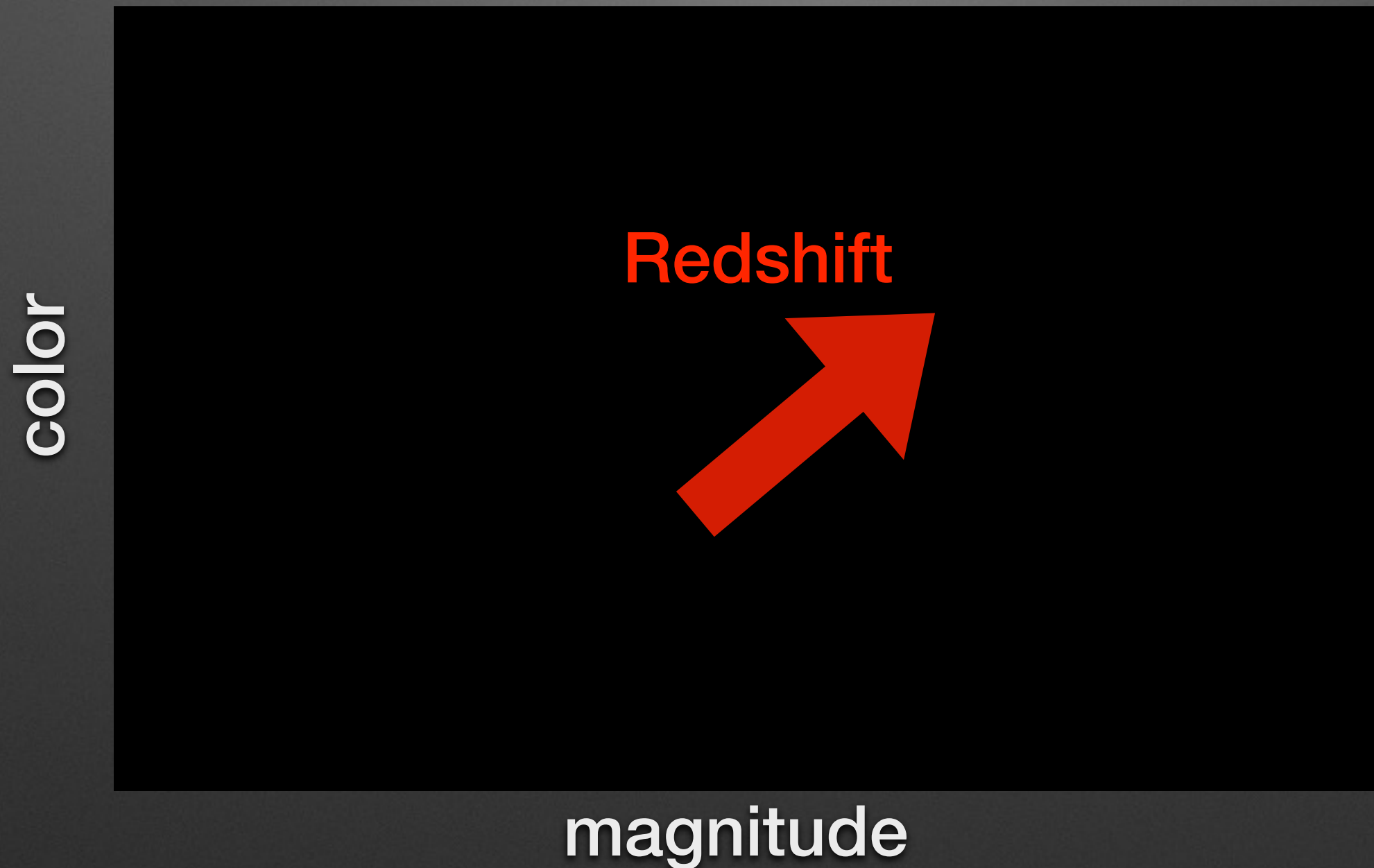
Gas distribution in lensed “normal” galaxy at $z \sim 2$

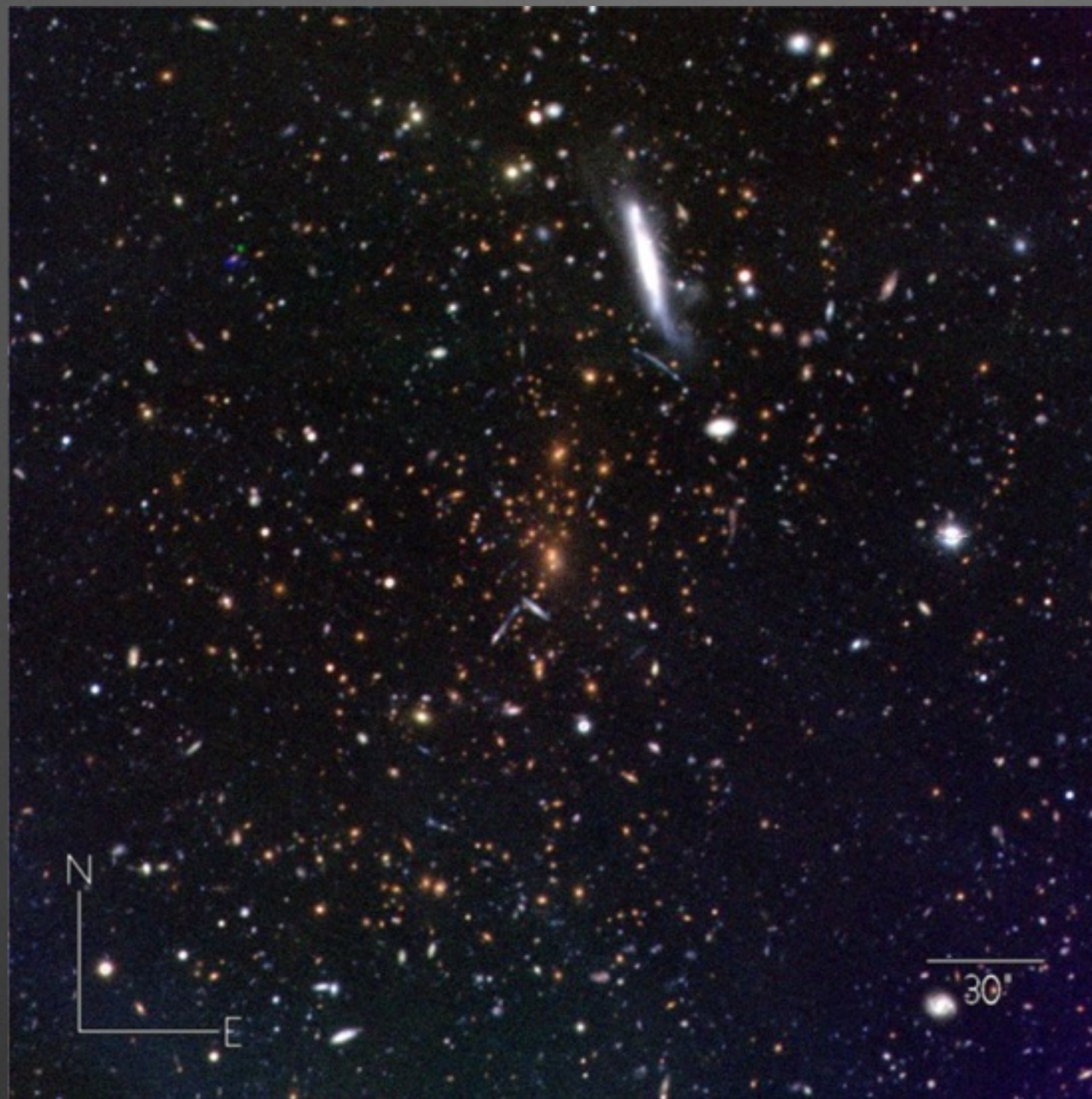
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Matt Bayliss
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- We have conducted several imaging surveys to identify clusters of galaxies for cosmological applications
- Red-sequence Cluster Survey 1 (RCS1)
 - 100 sq deg
 - R and z bands
- Red-sequence Cluster Survey 2 (RCS2)
 - 1000 sq deg
 - gr(i)z bands

We have used the homogeneity of the early-type population as a tracer of massive clusters





RCS2 J232727.6-020437
 $M \sim 3 \times 10^{15} M_{\odot}$, $z = 0.7$

Sharon et al. 2015





One of the
brightest arcs
known

RCS2 J032727-132627

$M \sim 10^{15} M_{\odot}$, $z_{CL} = 0.56$

$z_{arc} = 1.7$, $r_{AB} = 19.1$

38" length in 4 images

$u \sim \langle 30x \rangle$ from 4x to 100x

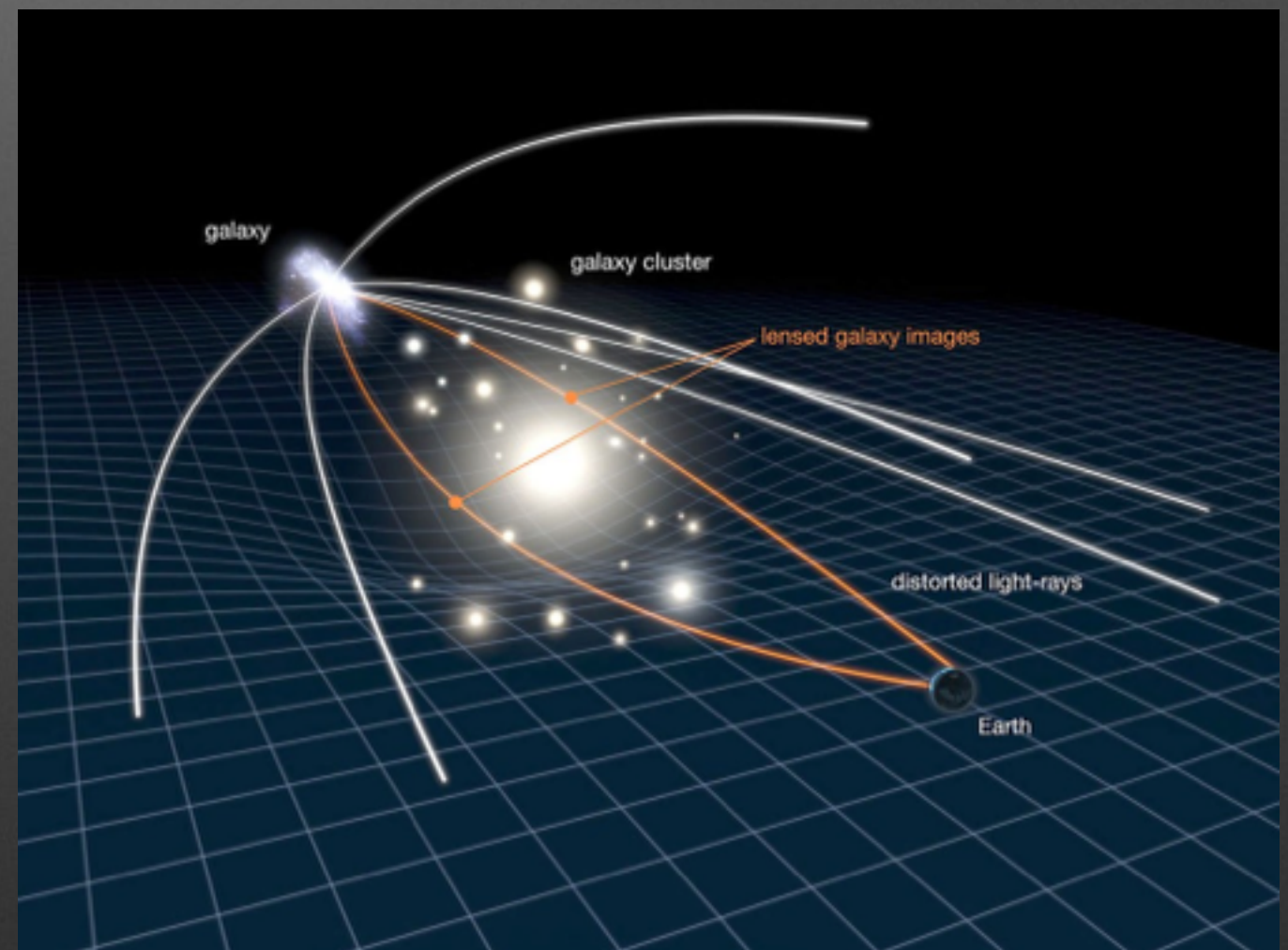
Wuyts et al. 2010

Strong lensing depends on three elements

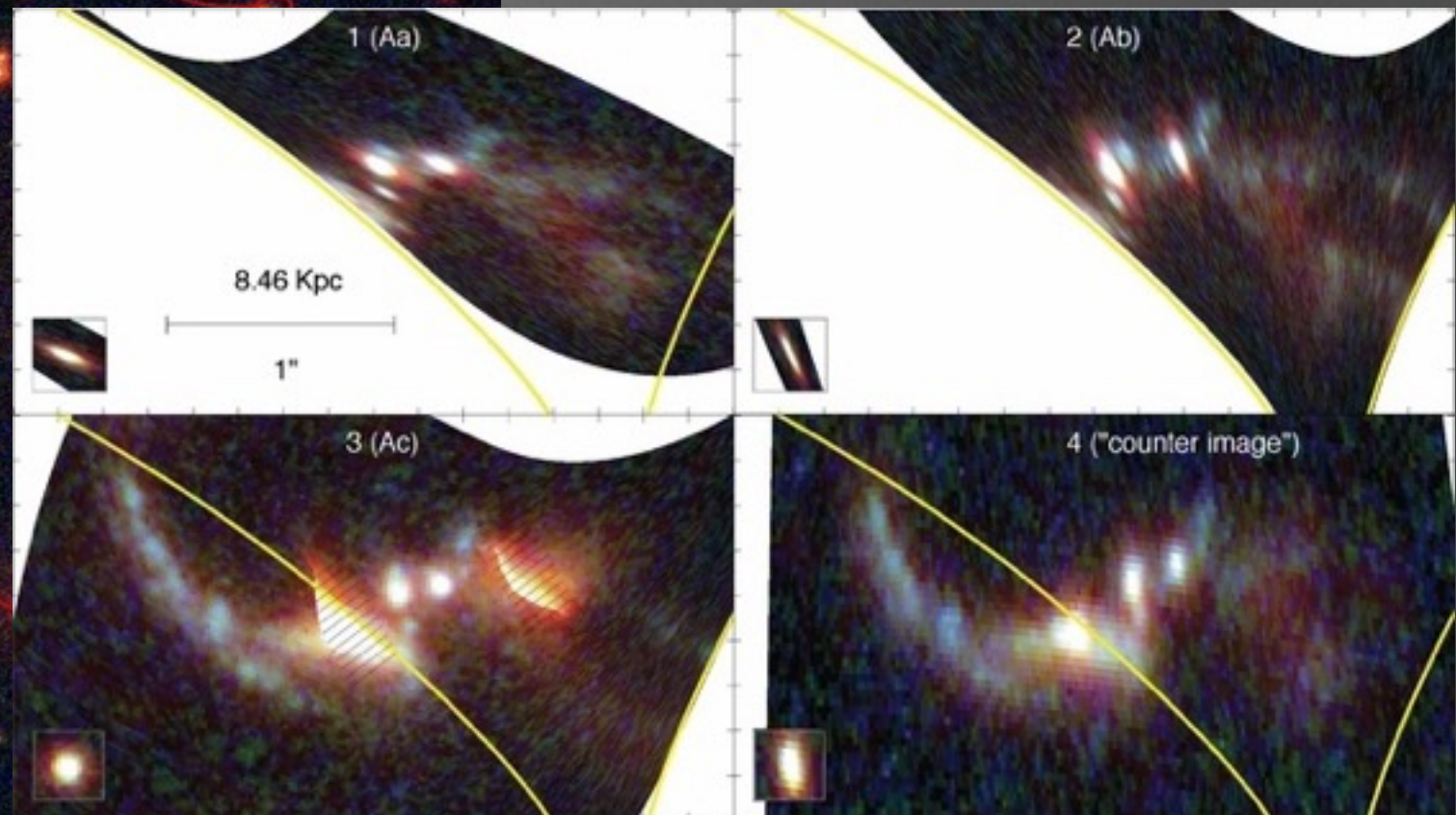
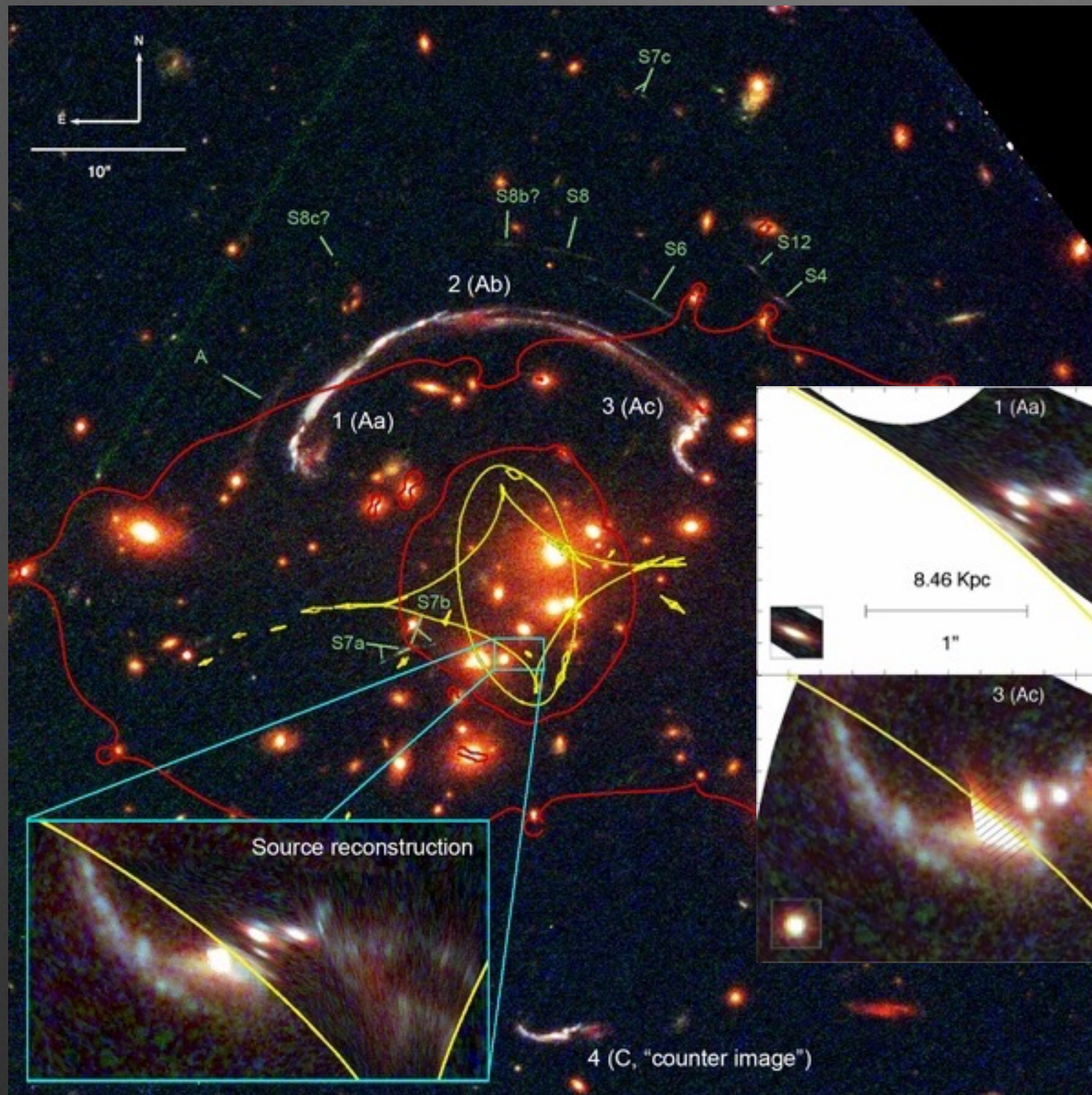
1. The lens - here the cluster of galaxies

2. The source - the distant galaxy

3. The distances - cosmology

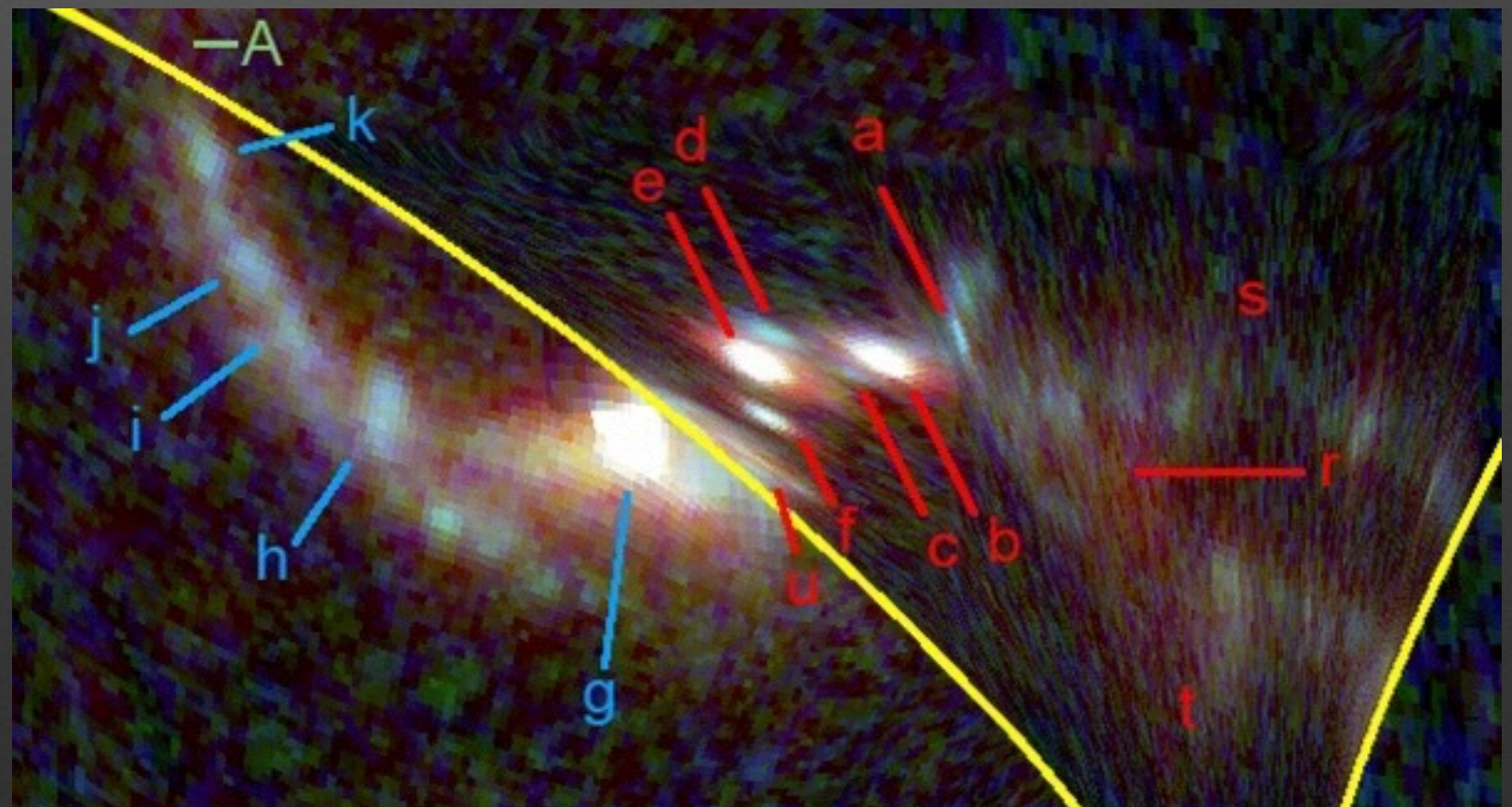
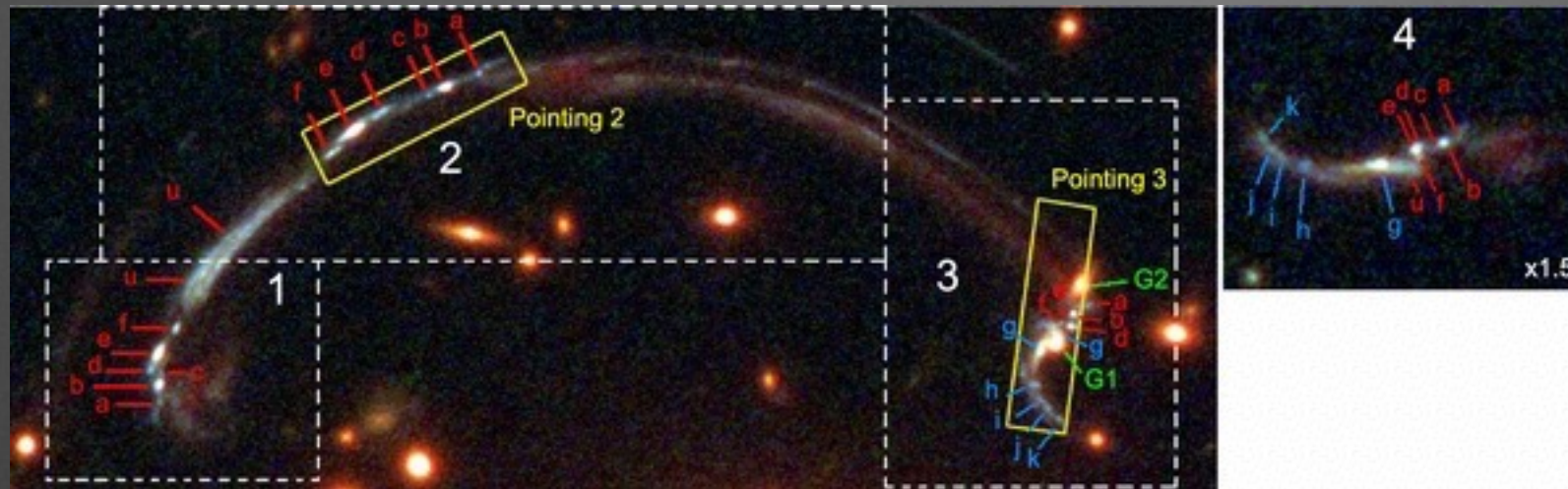


HST imaging Lens model

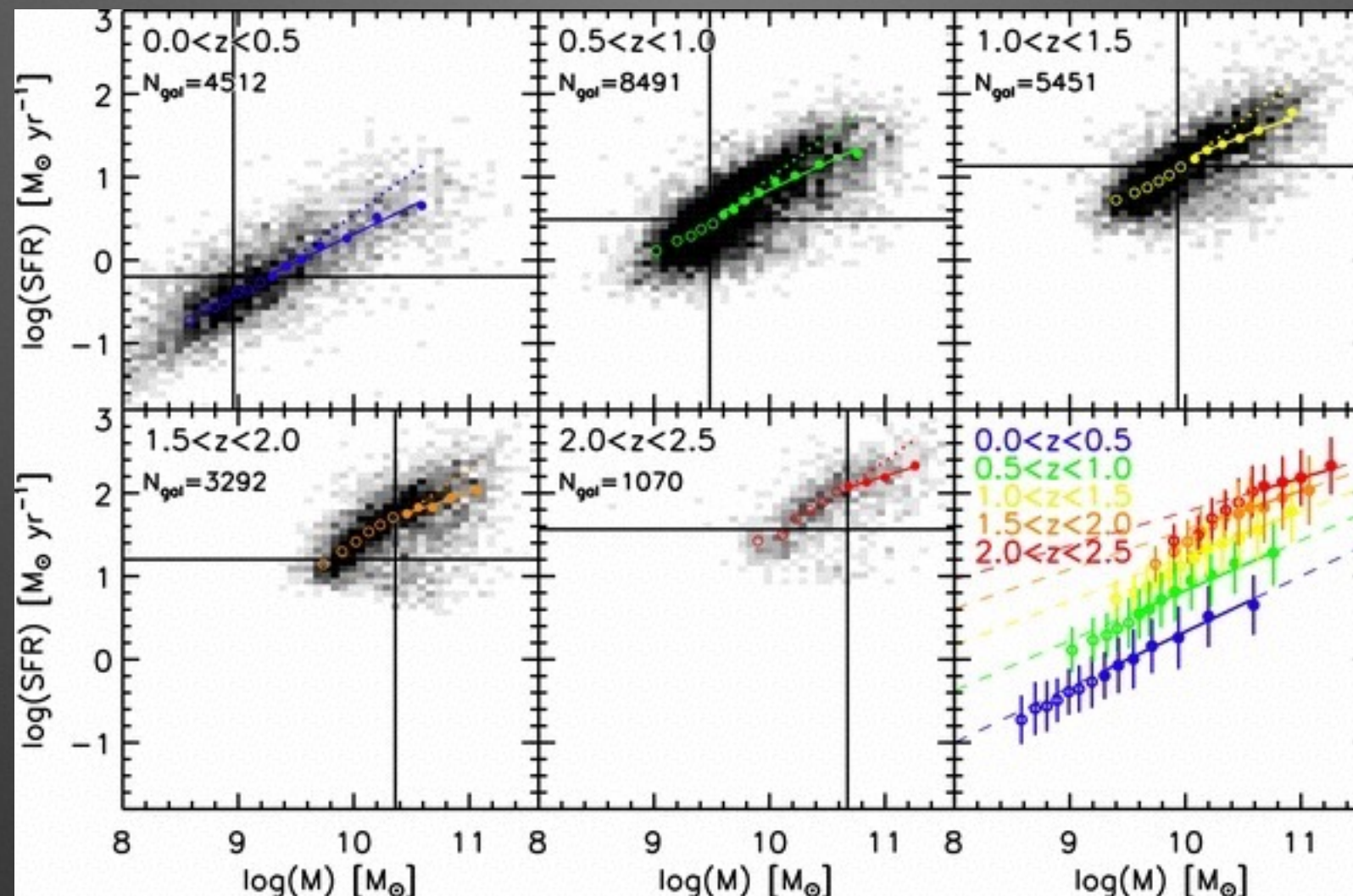


Sharon et al 2012

An accurate model allows us to study this galaxy in a x100 pc scale



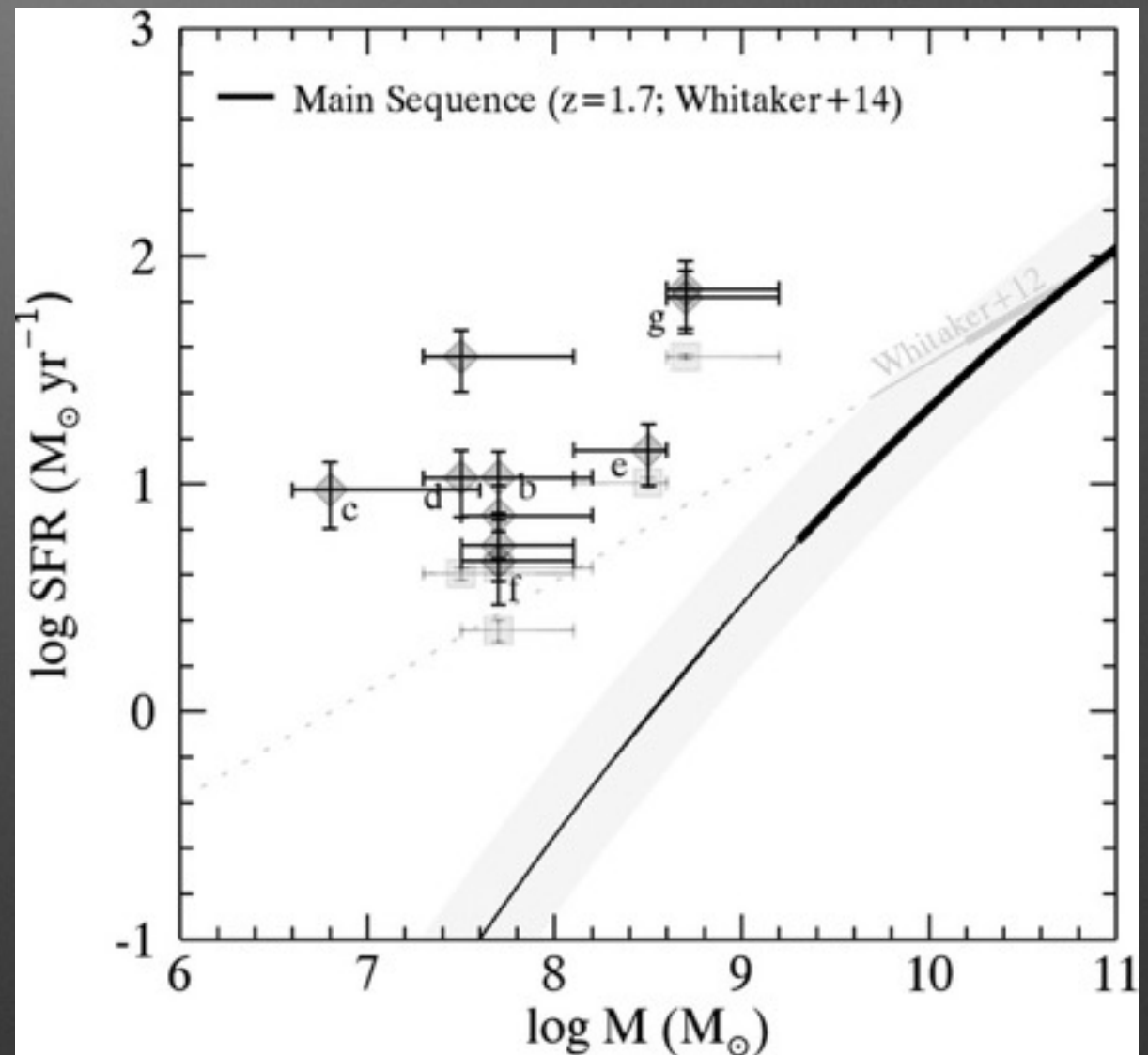
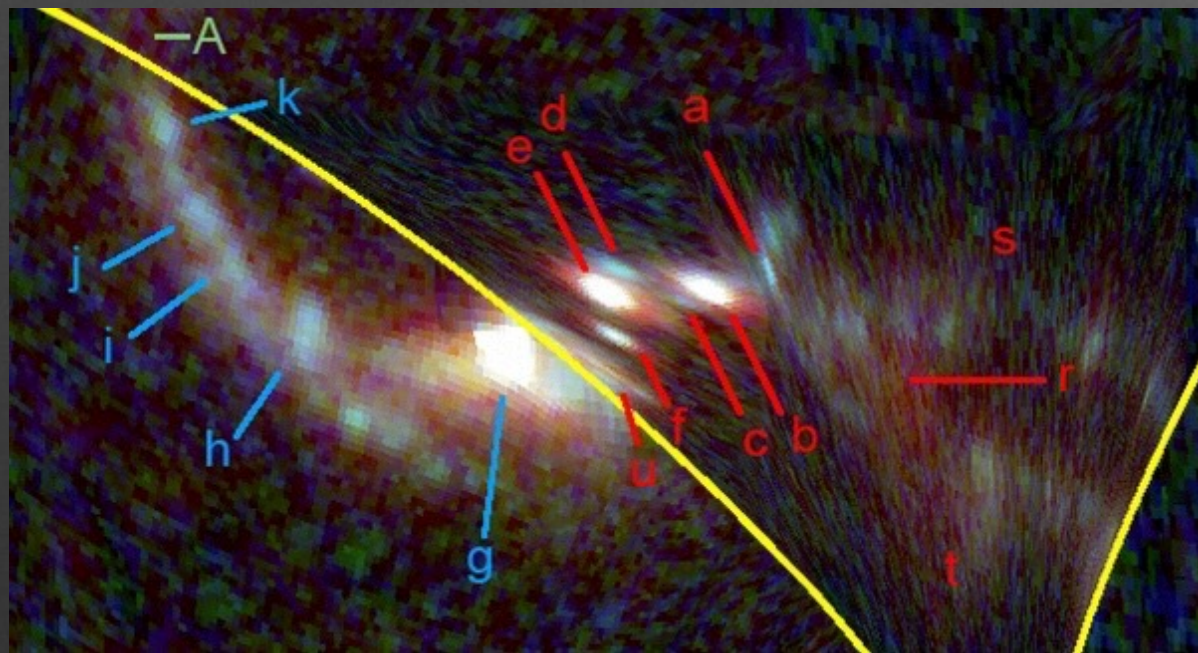
SFR mass sequence for star forming galaxies (integrated quantities)



Whitaker et al 2012

RCS0327 provides an opportunity for probing low mass galaxies at high redshifts

*individual clumps



Whitaker et al 2014