





TMT/IRISのマイクロレンズアレイ及びスライサーIFU





TMT/IRIS







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IRIS hung under NFIRAOS





IRIS Capabilities





- First Light Imager and Spectrograph working in parallel at the diffraction limit of the Thirty Meter Telescope.
 - NGSAO and LGS MCAO with NFIRAOS 0
 - Wavelength Range 0.84-2.4 microns 0
 - RMS Wavefront Error < 40 nm in fine scales 0
 - High Order Atmospheric Dispersion Correction 0
- On-Instrument wavefront sensors (OIWFS).
 - Three sensors to measure tip/tilt, focus and distortion across field. 0
 - Near infrared sensors to gain from NFIRAOS AO correction. 0
- "Wide-Field" Imager

- 撮像系
- 34 arcsec field of view (2x2 grid of H4RG-10 Teledyne Detectors)
- 4 mas plate scale (Nyquist @ 1.15 μm)
- Integral Field Spectrograph (H4RG-15 Teledyne Detector)
 - IFS with Four Plate Scales (4, 9, 25 and 50 mas per sample)
 - Up to 14,378 individual, simultaneous spectra.
 - Spectral Resolutions of 4000, 8000 and few exotic modes





Integral Field Spectrograph Optomechanical

















Three OIWFS Arms patrol 2 arcminute NFIRAOS output field.













SLICER SPECTROGRAPH





- Coarse platescales/ Wide FOV
- Slicer has 88 units rearranged into 2 pseudoslits, each with 2x22 "brick wall" pattern
 - 1.125"x2.2" @ 0.025" scale
 - 2.25"x4.4" @ 0.050" scale
 - Data Product: 45 x 88 x 1600(λ) elements



Figure 60 - CAD model of the IRIS image slicer shown (left) in isometric and (right) side view. All 88 channels are shown. The trace begins at the slicer stack and ends at the slicer long slit plane. Note that the IFU is shown here



30

20

10

0

-10

-20

-30

-30

-20

Detector Y (mm)





LENSLET ARRAY





- 112 x 128 elements each with 500 spectral elements
 - 0.45"x0.51" @ 0.004" scale
 - 1.01"x1.15" @ 0.009" scale
- Spot sizes set primarily by diffraction





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Simulated Spectra

Thirty Meter Telescope



Stretched vertically to show individual spectra (only lower 512 rows)



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IRIS Imager/IFS capabilities

Table 2: A top-level summary of IRIS capabilities with an R=4,000 spectral resolution requirement and a higher spectral resolution R=8,000 and 10,000 as an instrument goal. Both the grating and filter selection have been prioritized by the science team.

Capability mode	Spatial sampling (mas)	Field of View (arcsec)	Resolution (λ/dλ)	Min/Max wavelength (µm)	Bandpass ³
Imager	4 mas	34 x 34	Set by filter	0.84-2.4	60 filters BB ¹ and NB ²
Slicer IFS					
88x45	50 mas	4.4 x 2.25	4,000, 8,000	0.84-2.4	20%,5%
Spaxels	25 mas	2.2 x 1.125	4,000, 8,000	0.84-2.4	20%,5%
Slicer IFS					
44 x 45	50 mas	2.2 x 2.25	4,000, 8,000	0.84-2.4	20%,5%, H+K
Spaxels	25 mas	1.1 x 1.125	4,000, 8,000	0.84-2.4	20%,5%, H+K
Lenslet IFS	J				
112x128	9 mas	1.01 x 1.15	4,000	0.84-2.4	5%
Spaxels	4 mas	0.45 x 0.51	4,000	0.84-2.4	5%
Lenslet IFS	1	and the second	A A		
16x128	9 mas	0.144 x 1.15	4,000, 8,000	0.84-2.4	20%, H+K
Spaxels	4 mas	0.064 x 0.51		0.84-2.4	20%, H+K



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IRIS gratings (2017/3/15)





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Grating Rankings (2018/5/3)

							-
		ISDT		IRIS Science Team		IRIS Science Team + ISDT	
		RANKING		RANKING		RANKING	
		(Weighted Ranks)		(Weighted Ranks)		(Weighted Ranks)	
	H4000	6.82	K4000	9.72	K4000	8.33	
	K4000	6.59	H4000	8.61	H4000	8.18	
	J4000	6.38	J4000	8.33	J4000	7.68	
	H8000	5.97	HK4000	6.83	HK4000	6.46	
	J8000	5.74	Z4000	6.33	Z4000	6.04	
14 gratings	K10000	5.67	Y4000	6.06	H8000	5.89	
	HK4000	5.54	K8000 (Kn1-Kn3)	5.28	Y4000	5.75	
	Z4000	5.36	K8000 (Kn4-Kn5)	5.22	K8000 (Kn1-Kn3)	5.74	
	H10000	5.36	H8000	4.33	J8000	5.63	
	K8000 (Kn1-Kn3)	5.26	K8000 (Kbb)	4.28	K8000 (Kn4-Kn5)	5.30	
	Y4000	5.13	J8000	4.00	K10000	5.11	
	K8000 (Kbb)	4.82	Z8000	3.78	K8000 (Kbb)	4.95	
	K8000 (Kn4-Kn5)	4.64	H10000	3.67	H10000	4.95	
	J10000	4.46	K10000	3.61	Z8000	4.37	
	Z8000	4.36	Y8000	3.11	J10000	4.07	
	Y8000	4.28	J10000	2.83	Y8000	3.93	
	Z10000	4.13	Z10000	2.67	Z10000	3.47	
	Y10000	3.69	Y10000	2.17	Y10000	2.88	



IFS Saturation— 1 second integration





	4 mas Vega (mag)	9 mas Vega (mag)	25 mas Vega (mag)	50 mas Vega (mag)
Zbb	< 8.7	< 10.0	< 11.2	< 11.0
Ybb	< 8.5	< 9.9	< 10.9	< 10.9
Jpp	< 8.1	< 9.6	< 10.5	< 10.6
Hbb	< 7.6	< 9.1	< 10.2	< 10.4
Kbb	< 6.3	< 7.9	< 9.2	< 9.5

Assuming 90% non-linearity at 90,000 e-



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Field setup on gravitationally-lensed field with OIWFS, Imager, & IFS



Issues of saturation: IRIS imager 60 second integrations

Thirty Meter Telescope







All green highlighted sources will saturate given exposure time



IRIS IFS field of view









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IRIS Exposure Time Calculator

https://www.tmt.org/etc/iris



TMT INTERNATIONAL OBSERVATORY

IRIS Exposure Time Calculator

The IRIS Exposure Time Calculator (ETC) is a tool to assist with the development of science cases that involve observations with IRIS (TMT's first light diffration limited near-IR (1-2.4 μm) imager and Integral Field Spectrograph). The IRIS ETC carries out a full decomposition with the expected PSF and may take several seconds to run.

IRIS Exposure Time Calculator User Guide

Enquiries and questions to: instruments@tmt.org

USEFUL LINKS

IRIS Exposure Time Calculator User Guide

IRIS Instrument Description

IRIS team webpage

開発途中です。User Guideも含め、バグや 不具合、不明な点な どのフィードバックを歓 迎します。