

東大Tomo-e Gozenと京大MUレーダーによる 超微光流星の同時観測

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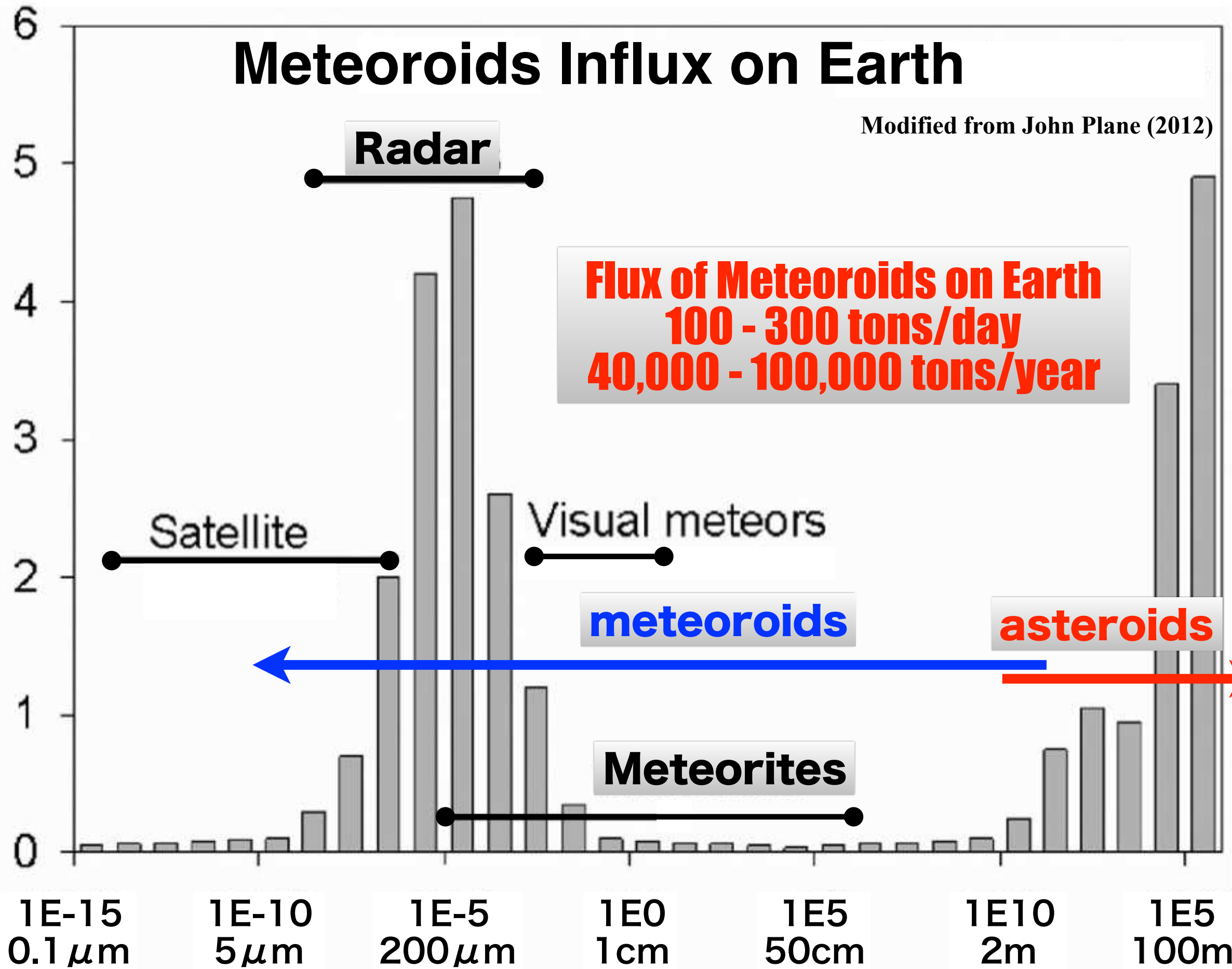
(4)National Institute of Polar Research, Tokyo, Japan

(5)National Astronomical Observatory of Japan, Mitaka, Tokyo, Japan

Meteoroids Influx on Earth

Modified from John Plane (2012)

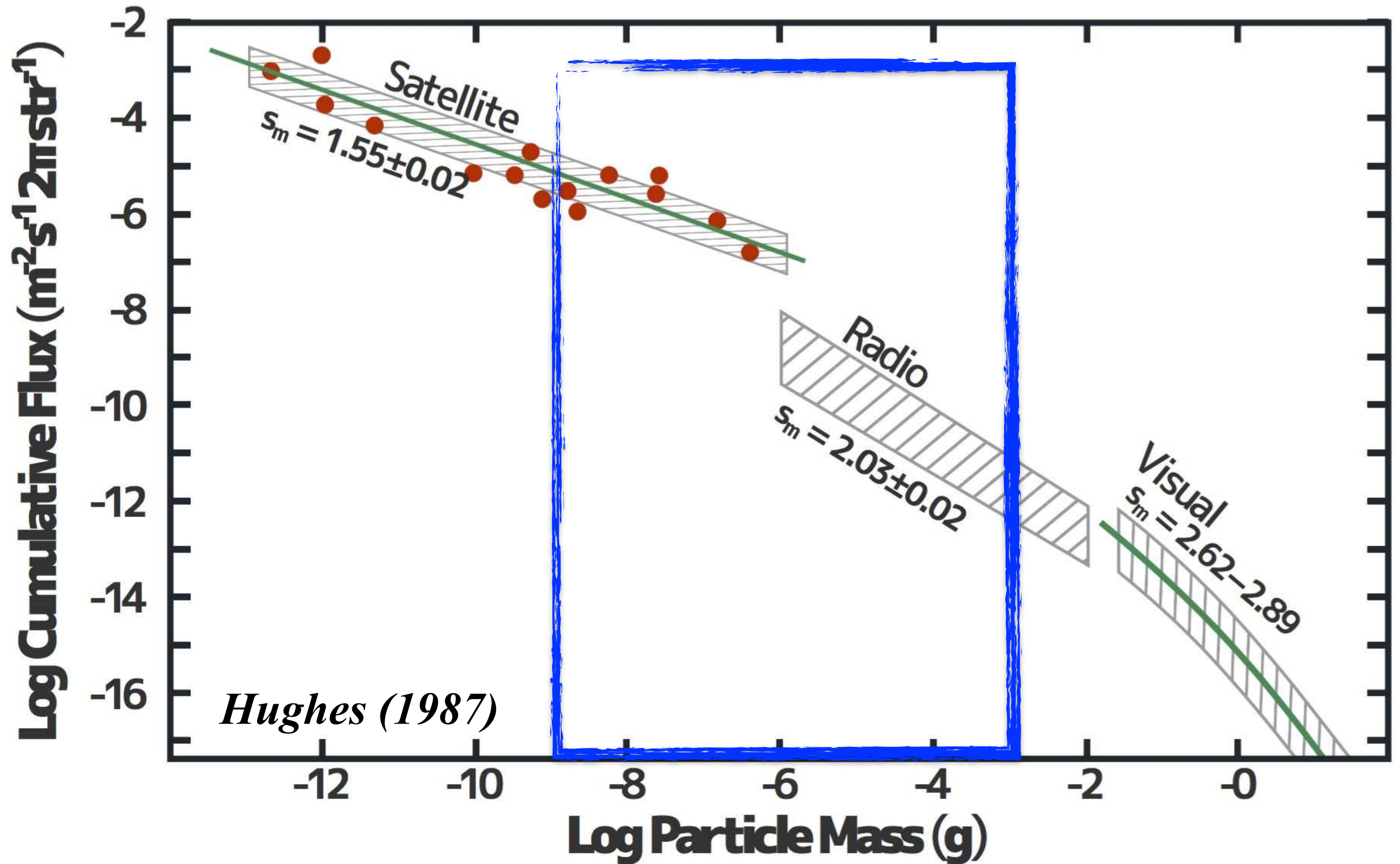
Mass Influx [10^6 kg/year]



Mass[g] and Diameter of particle

assuming $\rho=2.0$ g/cc

Sensitivity of MU Radar Meter Head-echo Observations



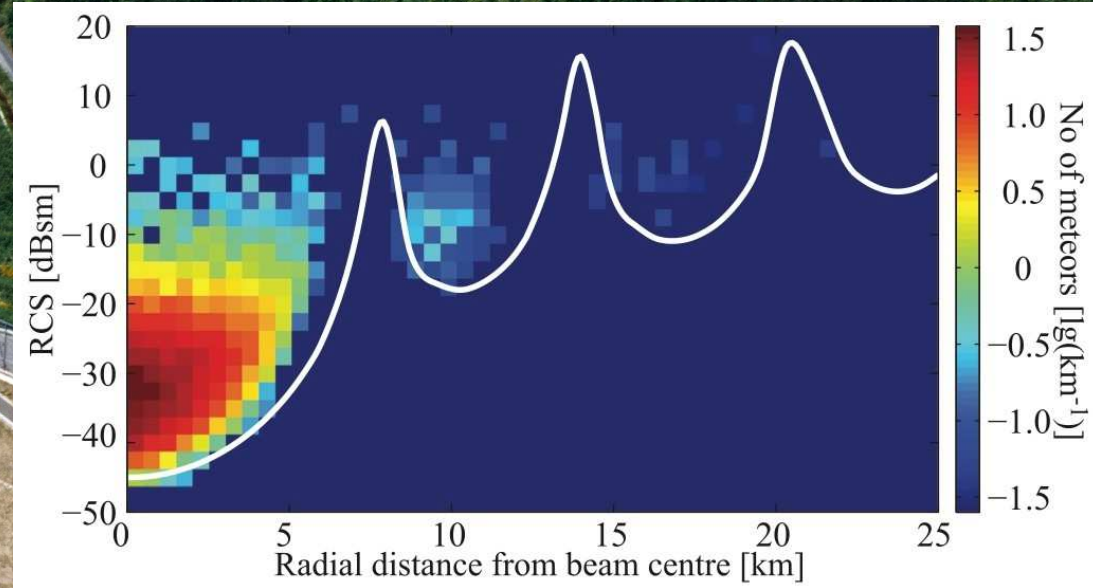
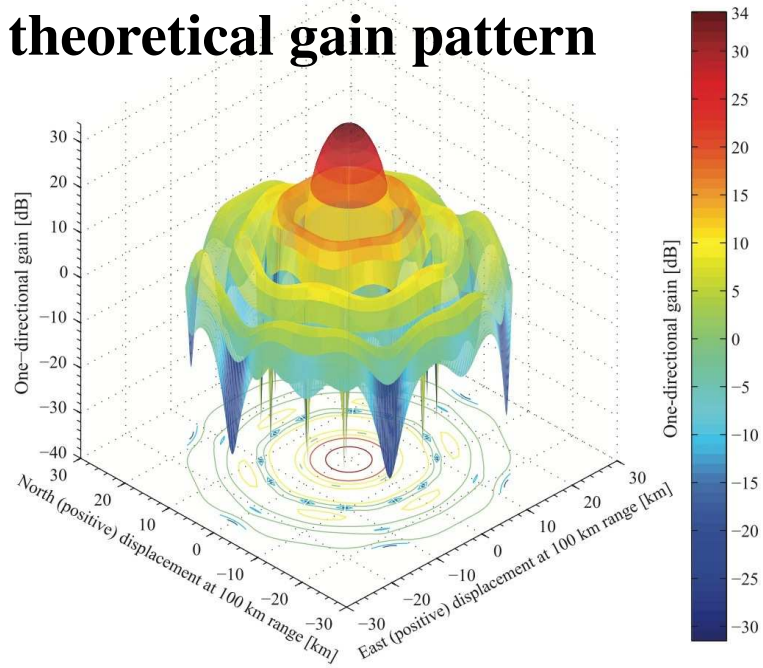
Kyoto University RISH MU Radar Middle and Upper Atmosphere Radar

Monostatic coherent pulse Doppler radar

VHF (46.5 MHz), 1MW peak power, 475 crossed Yagi antennas

Pulse length: 1-500 μ s, Antenna aperture: 8330m² (D=103m)

theoretical gain pattern



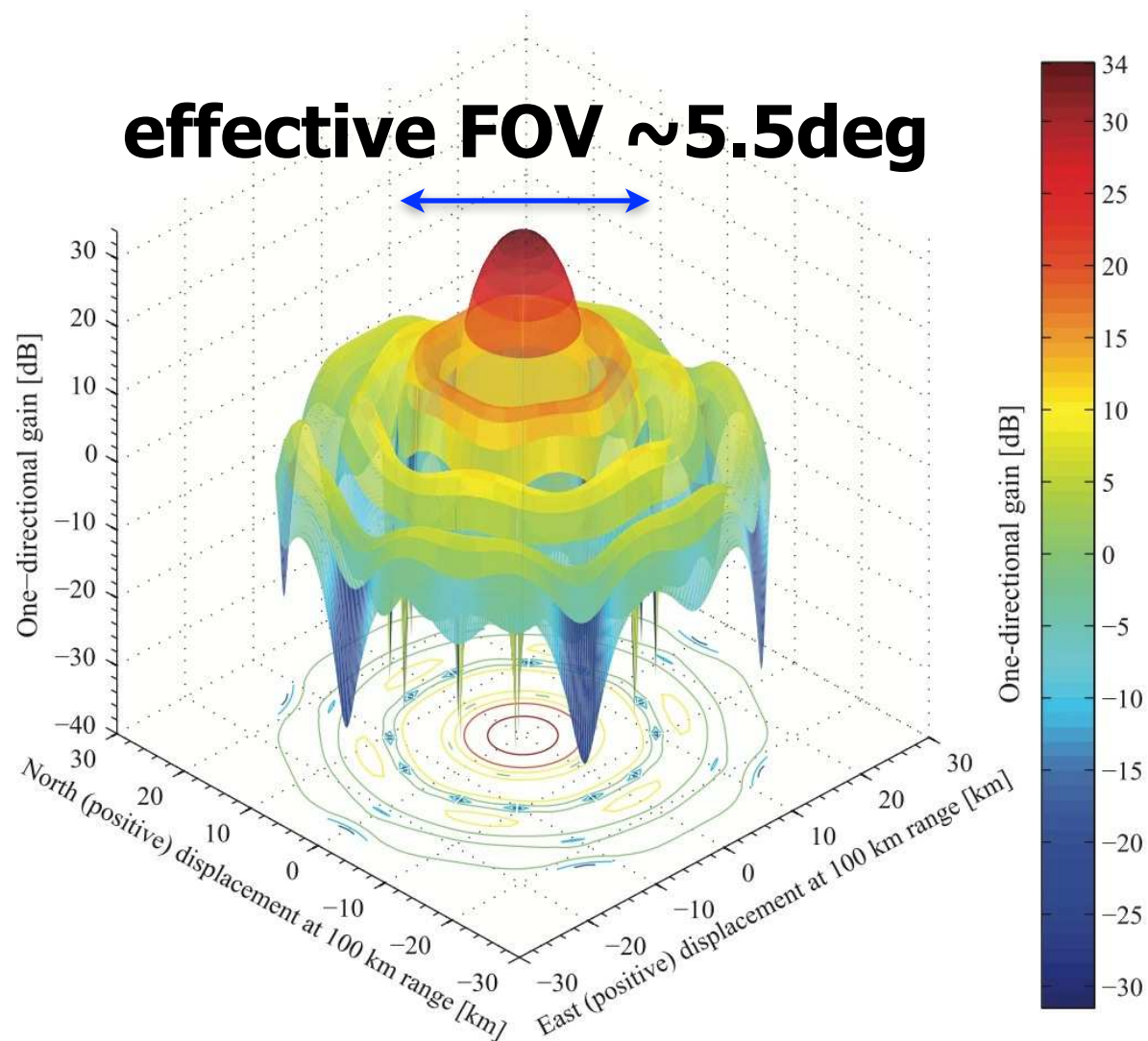
Observed number of meteors, normalized by beam area, versus RCS (Radar Cross Section) and radial distance from beam centre.

>180k meteoroids were detected during 2009-2016

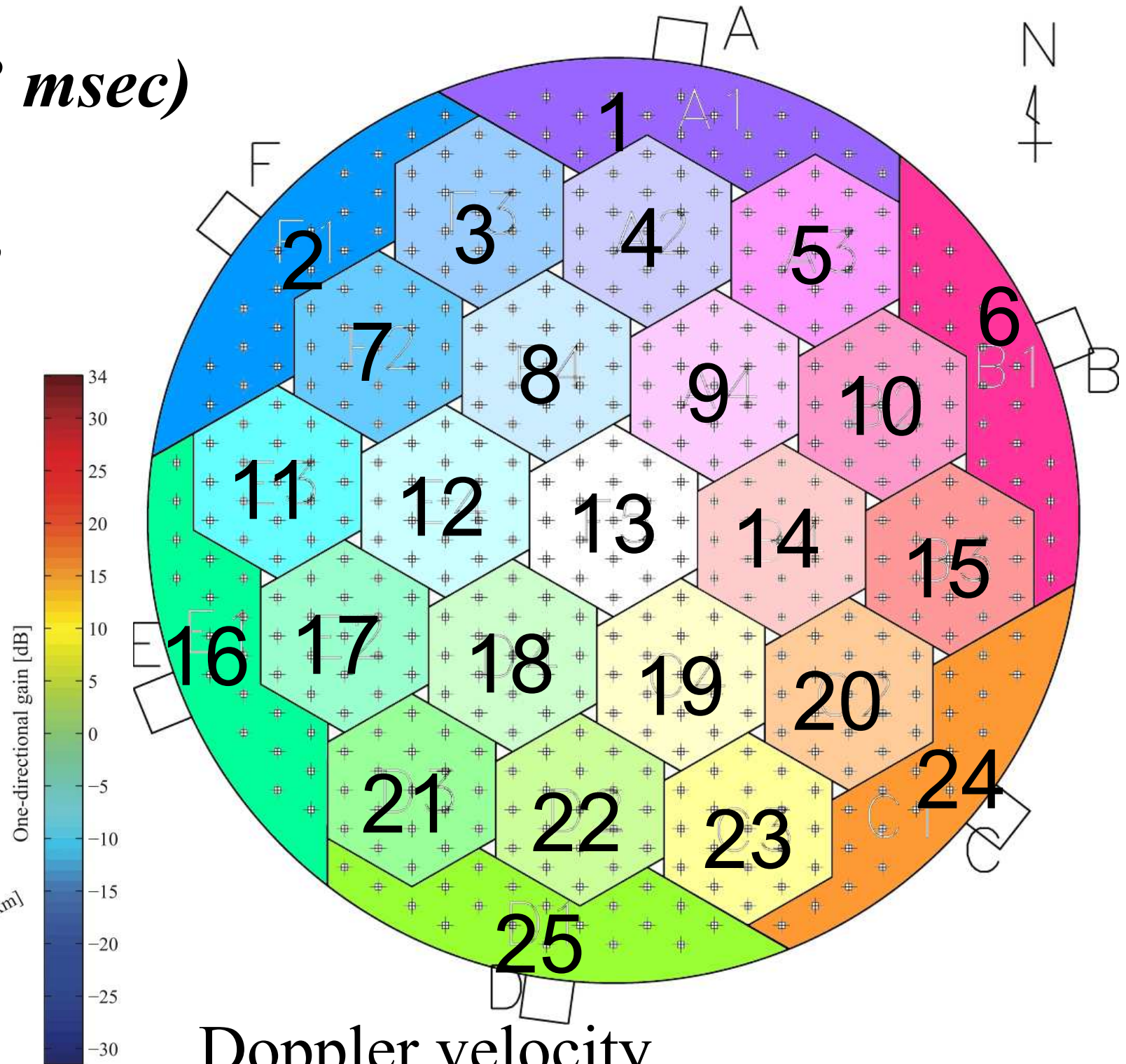
average velocity error = 0.25 km/s
average perihelion distance = 0.003 AU
3,000 - 4,000 meteor head echoes / day

Kyoto University RISH MU Radar Middle and Upper Atmosphere Radar

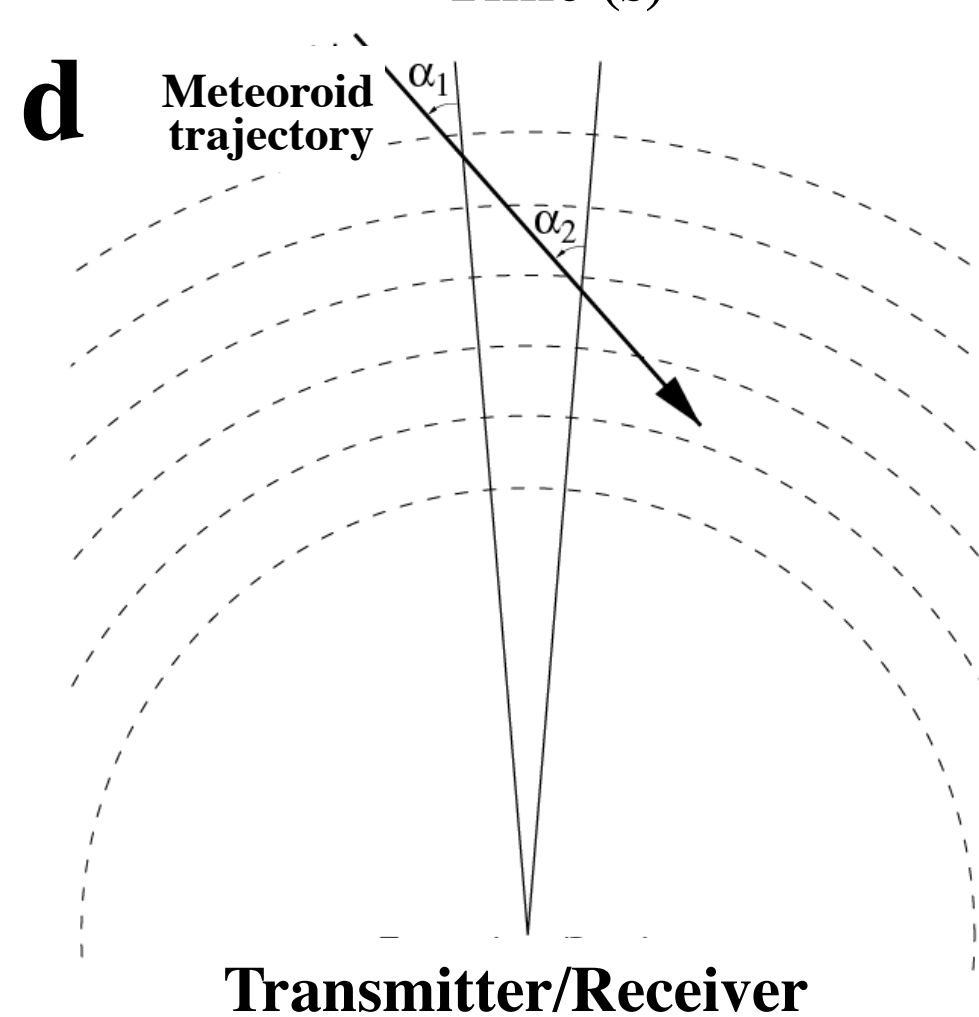
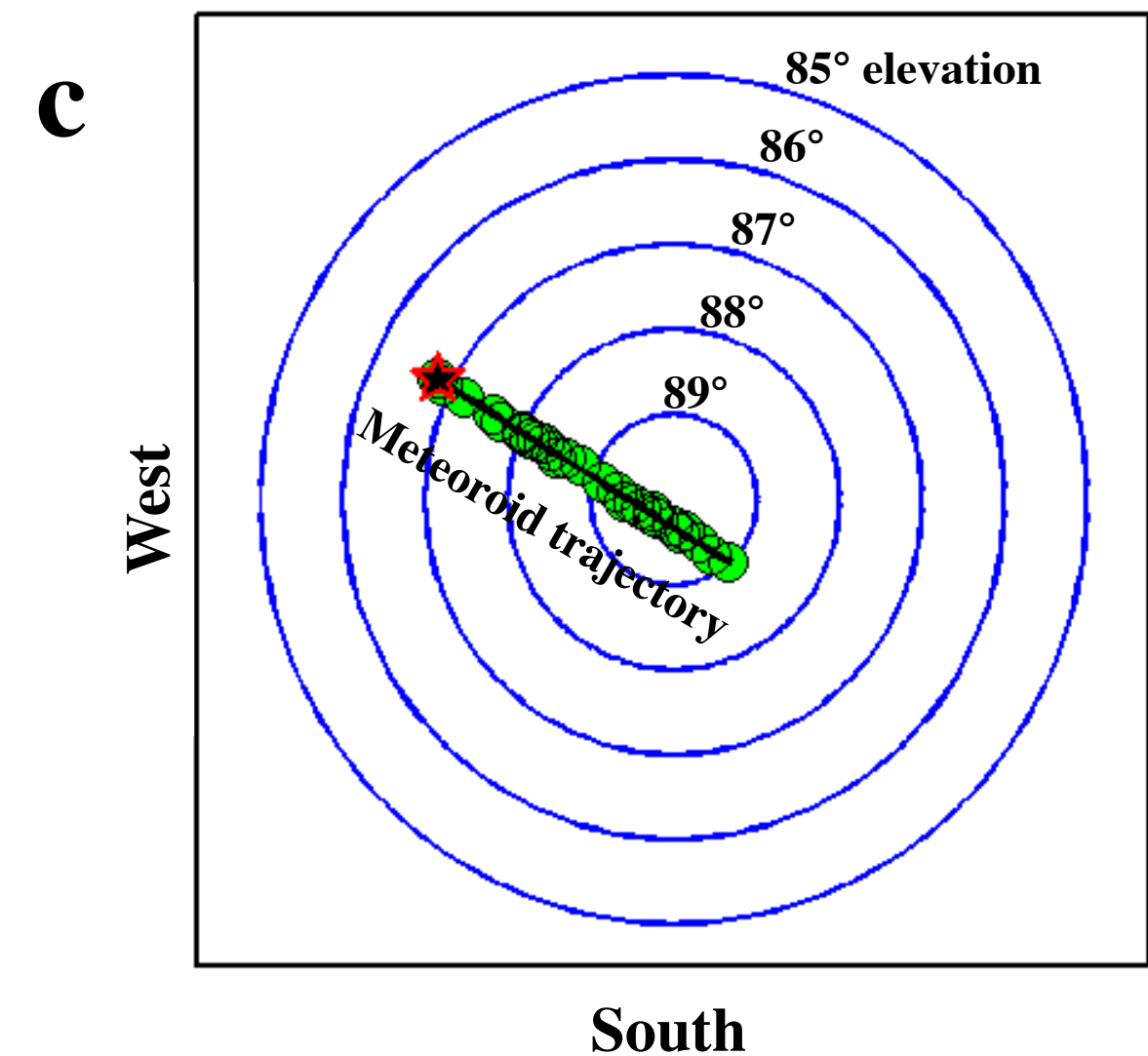
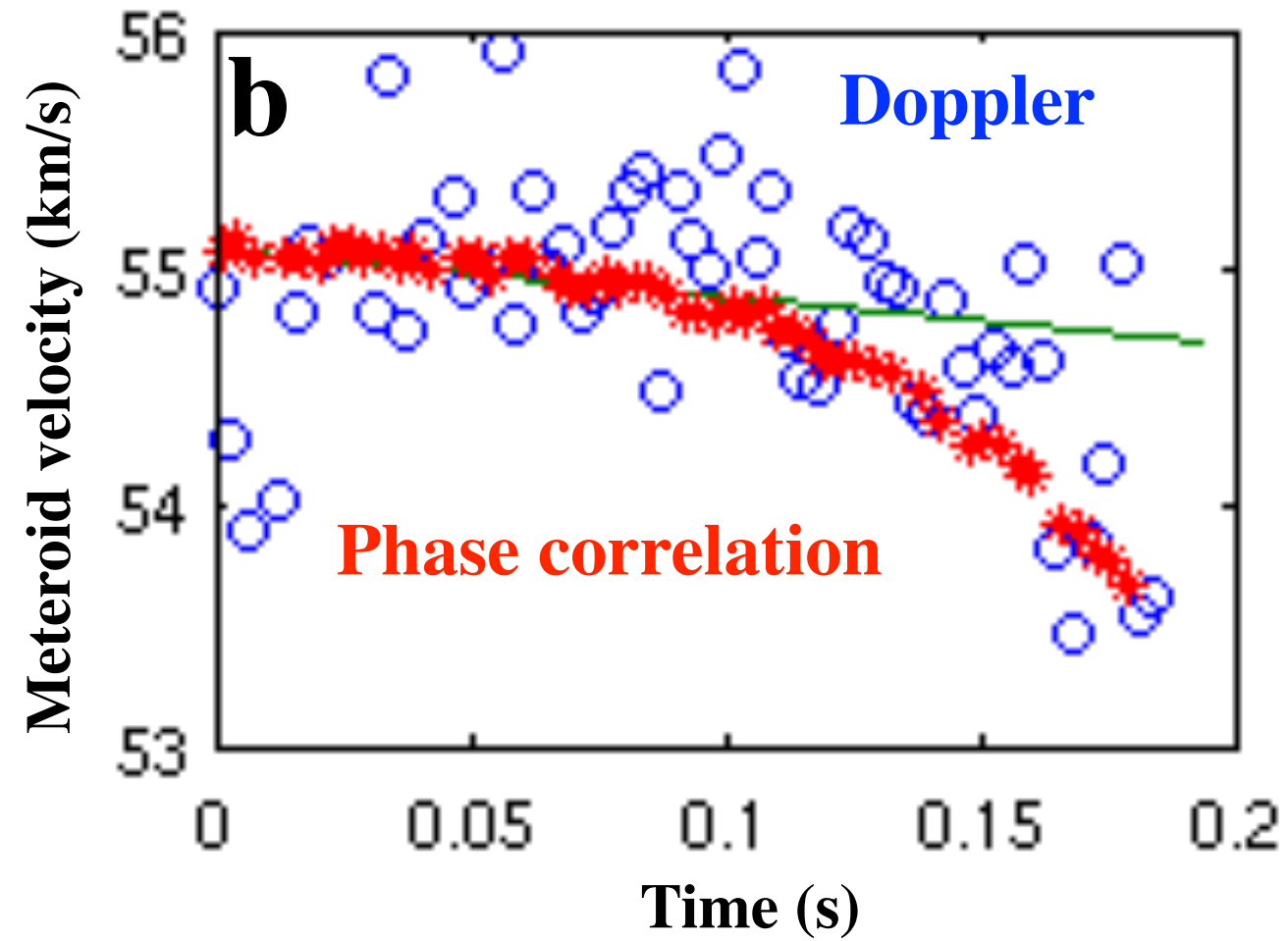
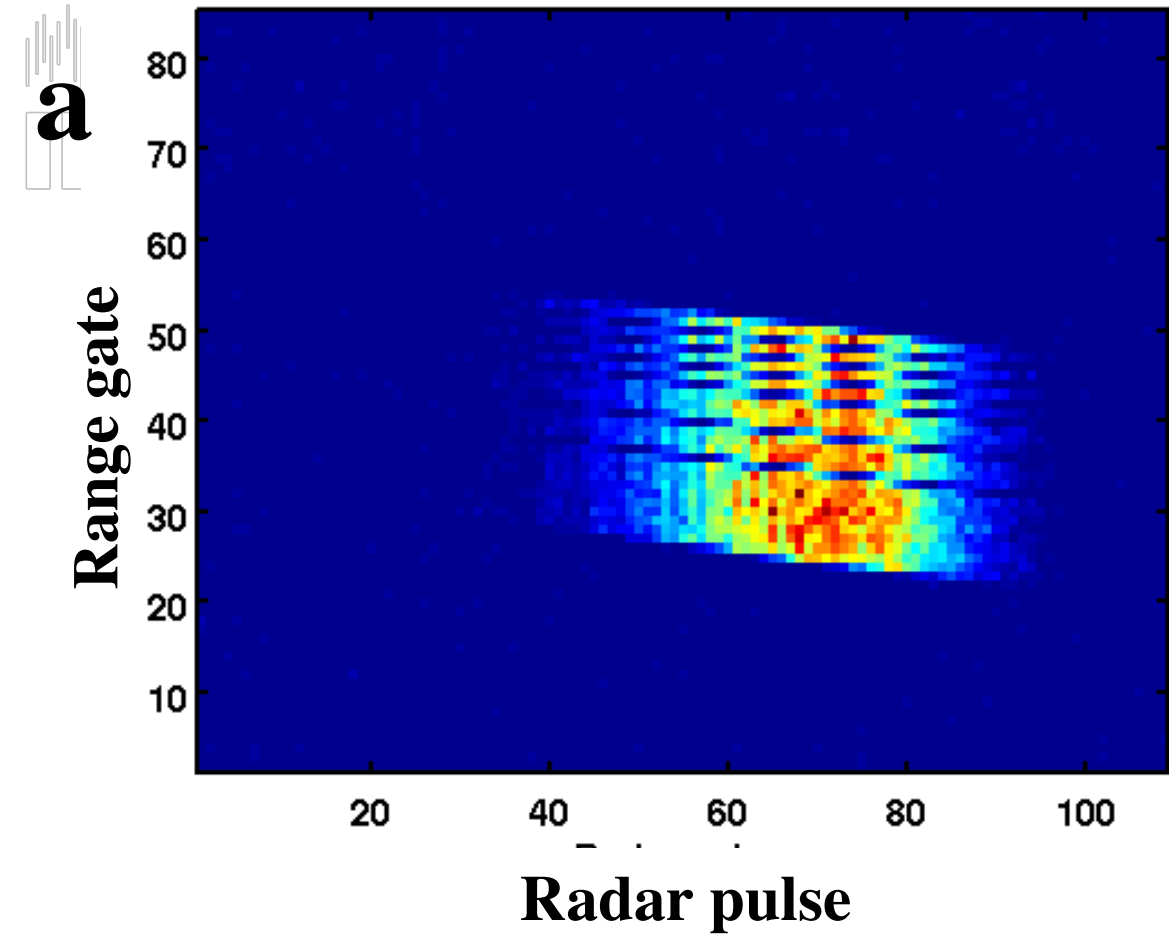
- *Digital 25 channels*
- *332 times per second (3 msec)*
- *85 ranges every 3 msec*
- *Data rate ~20 GB/hour*



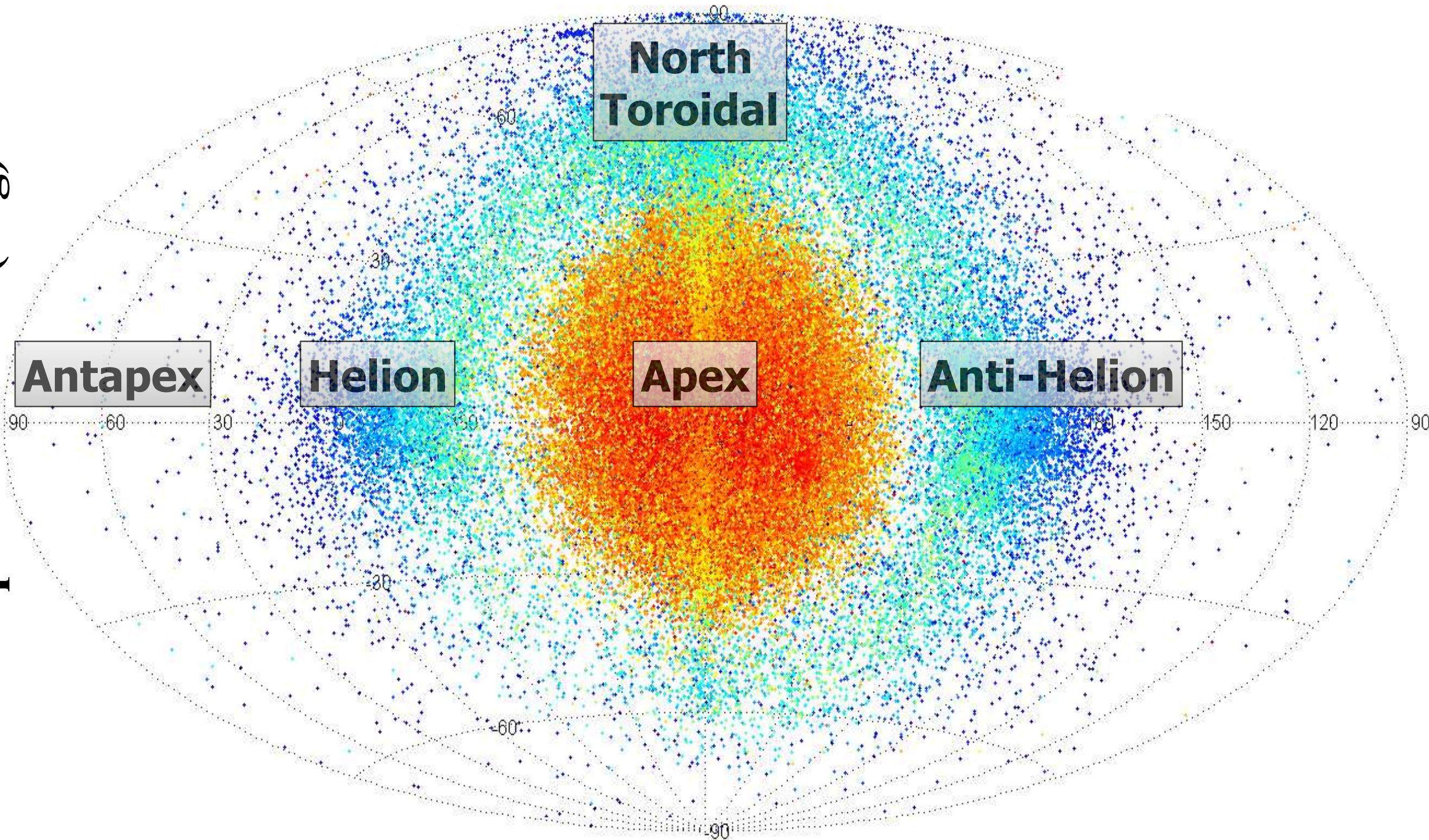
theoretical gain pattern



Doppler velocity
19 Interferometers



Ecliptic Latitude (deg)



Antapex

Helion

Apex

Anti-Helion



Geocentric velocity (km/s)

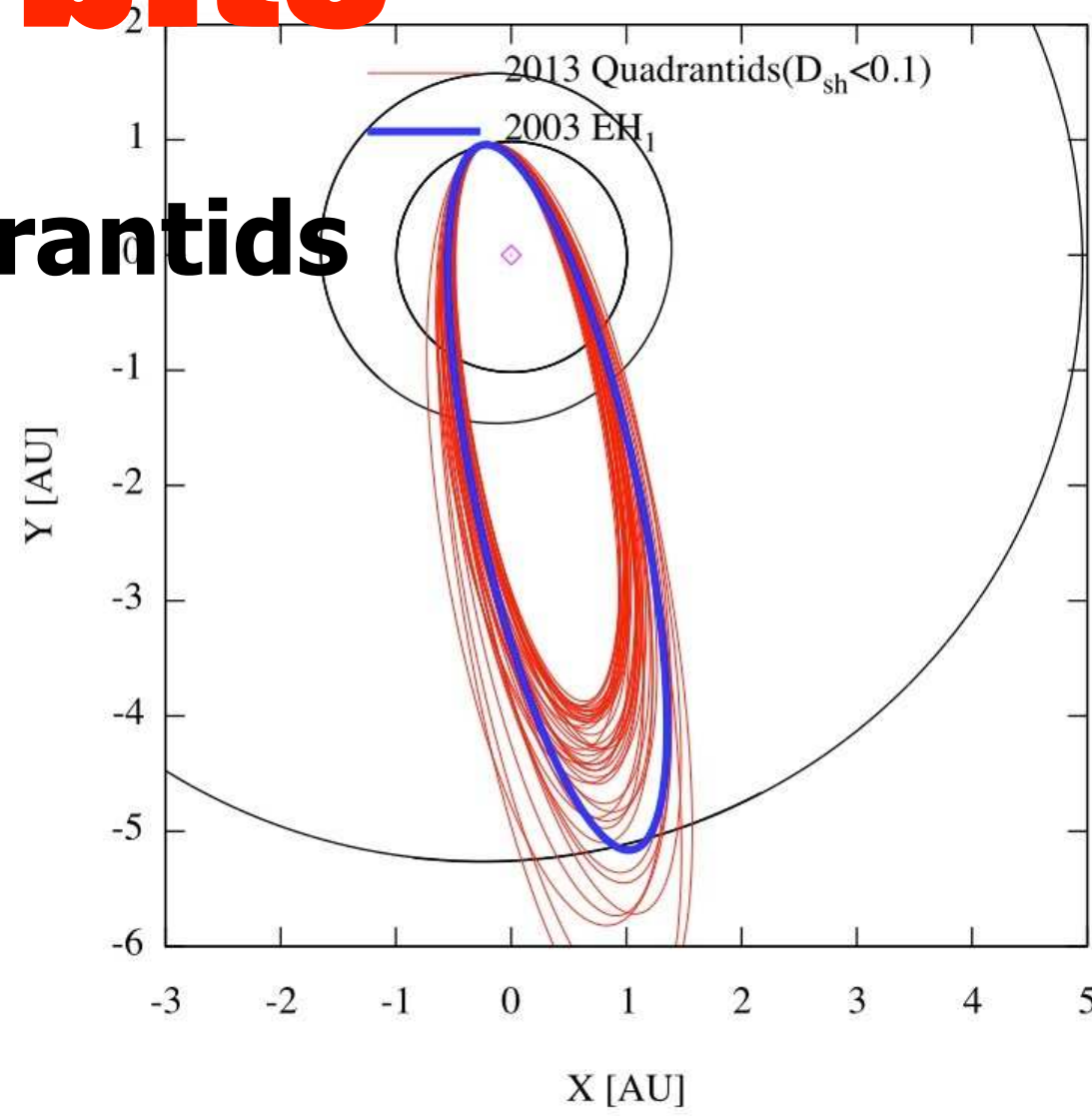
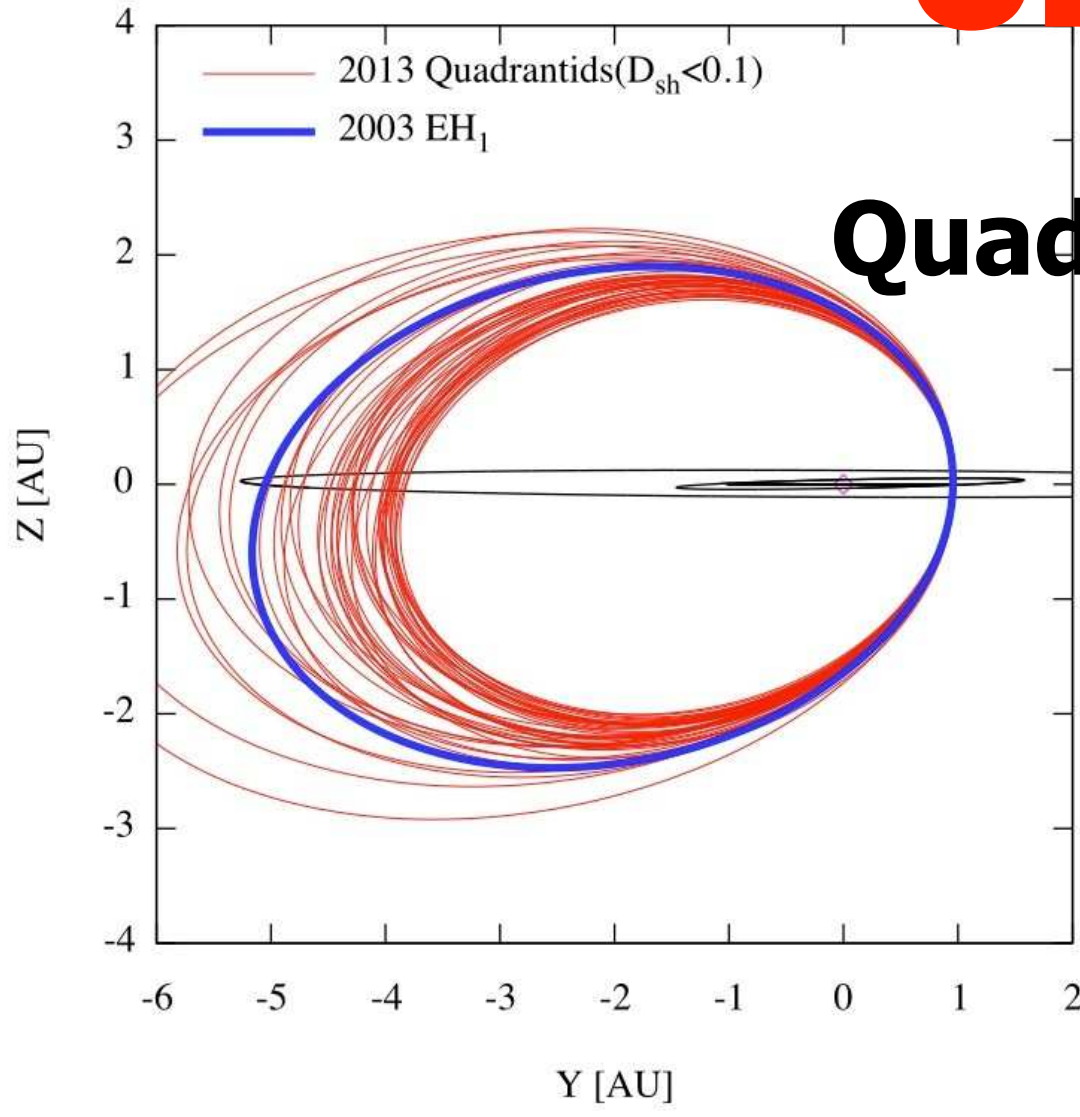
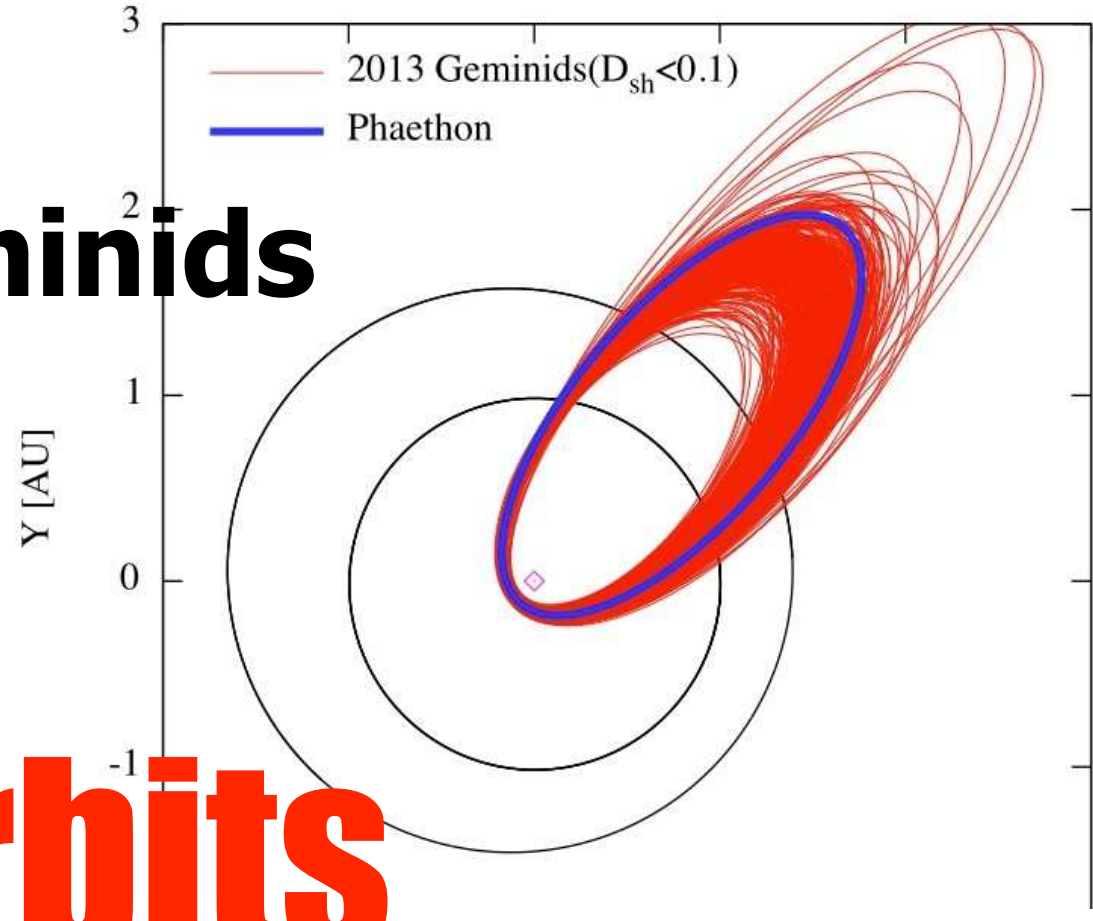
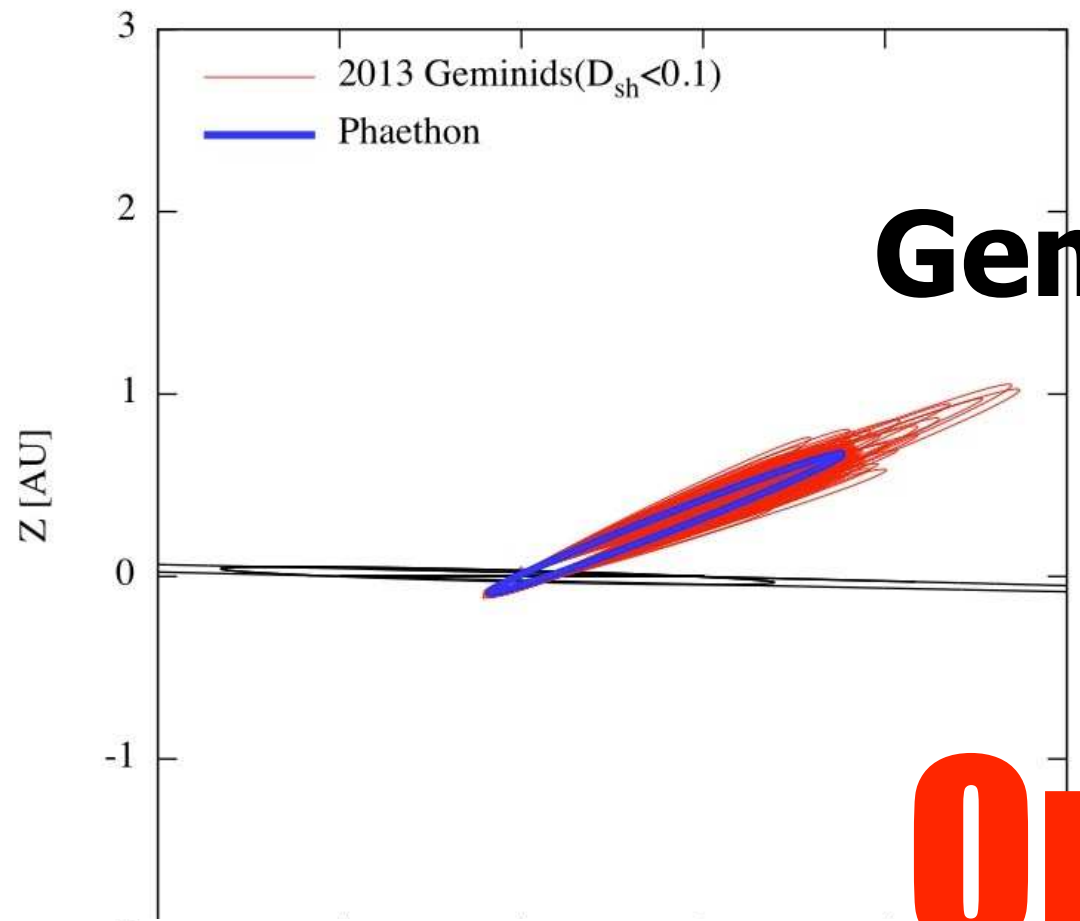
Kero+ (2010)
Abe, Kero, Nakamura+, in prep

Ecliptic Longitude from Earth apex (deg)

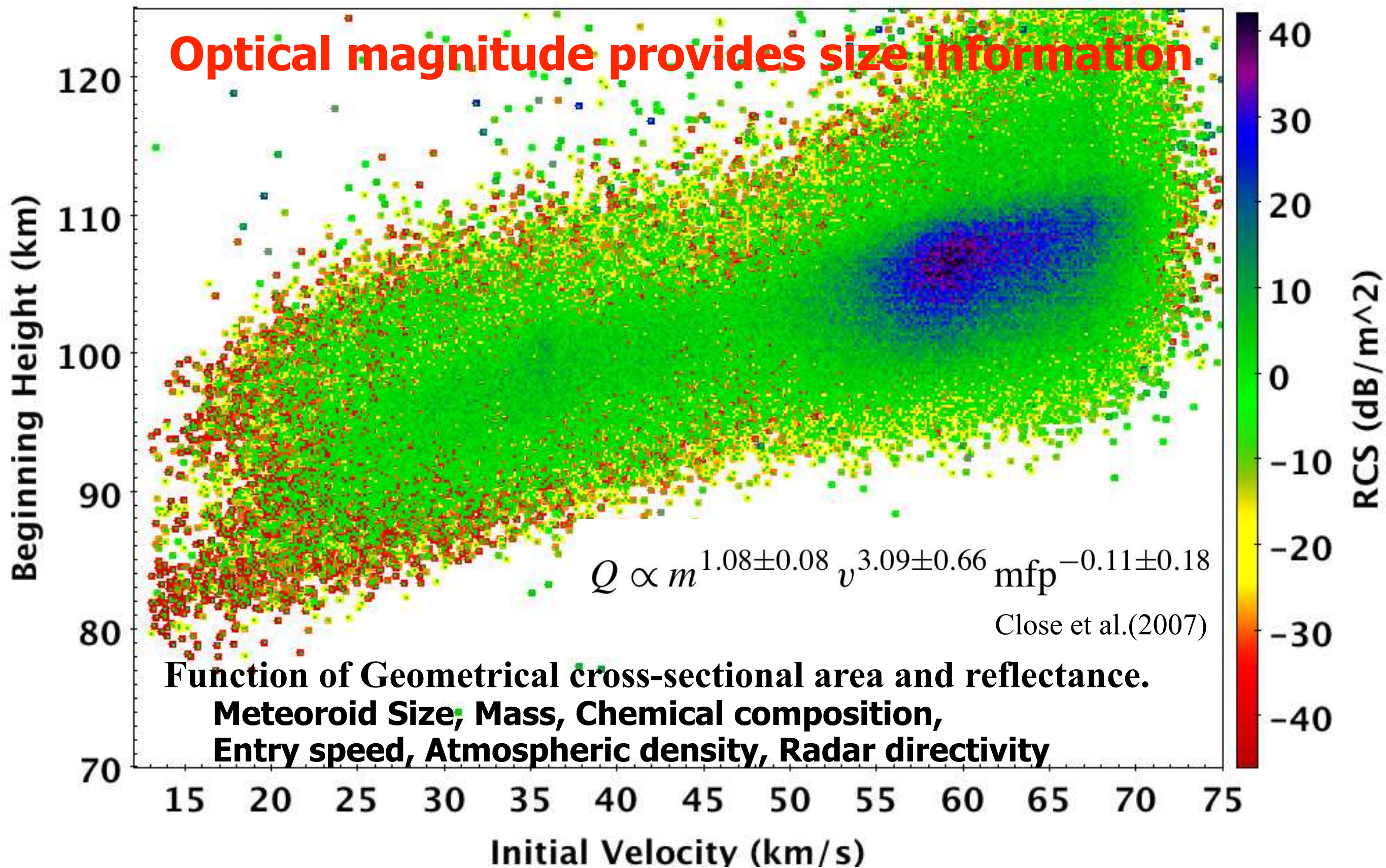
Comparison of Orbits between MU Radar and Optical Observations

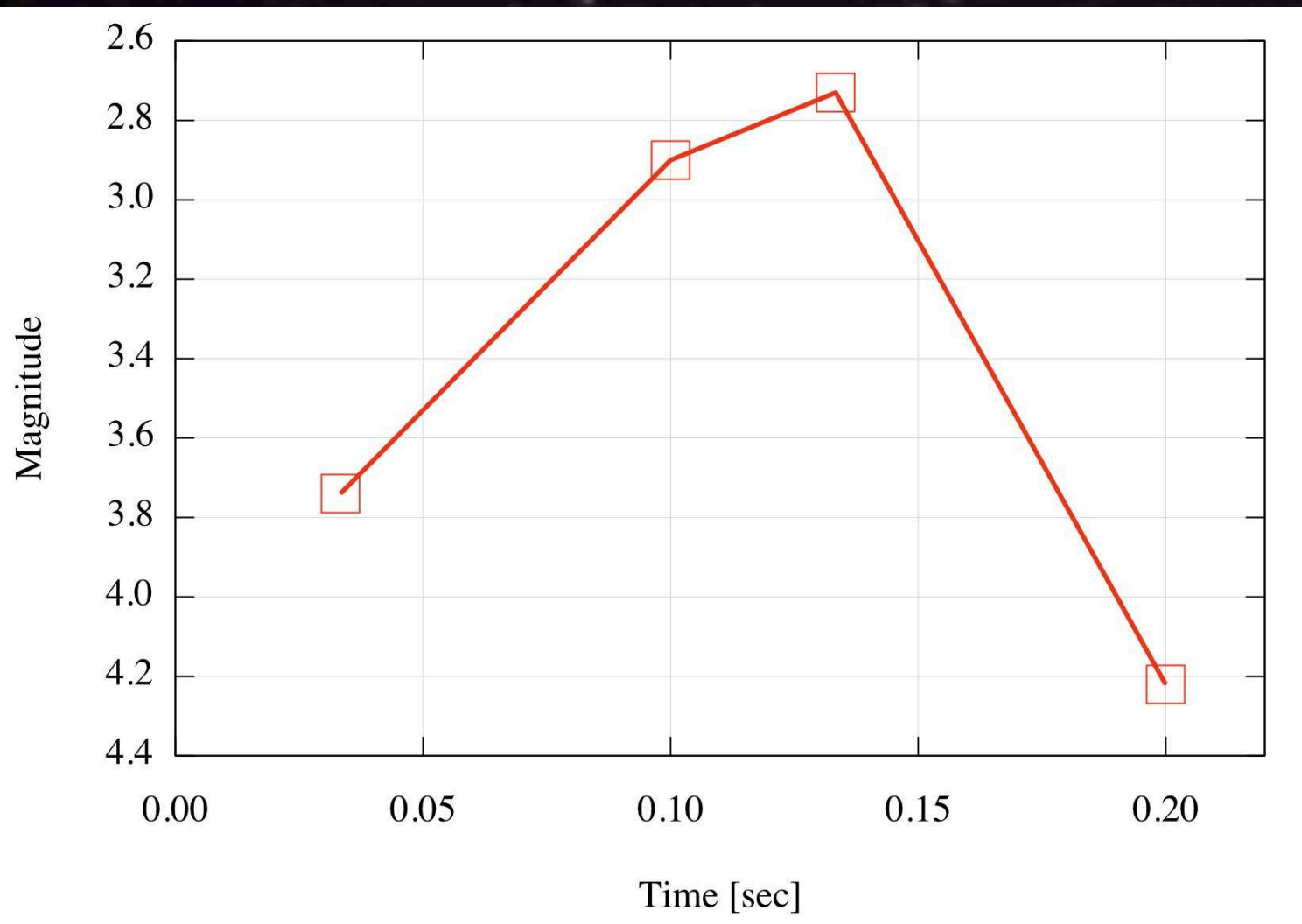
Object	<i>Date</i>	<i>a</i>	<i>e</i>	<i>i</i>	ω	Ω	D_{sh}
	<i>UT</i>	au	—	°	°	°	—
Phaethon	-	1.27	0.89	22.2	322.1	265.2	-
1-radar	Dec/14	1.27	0.89	23.6	325.1	262.6	
1-opt	15:29	1.22	0.88	23.5	325.1	262.6	0.013
2-radar	Dec/13	1.20	0.89	24.1	325.8	261.7	
2-opt	18:49	1.39	0.91	23.2	325.8	261.7	0.030
3-radar	Dec/13	1.21	0.89	22.5	324.5	261.6	
3-opt	16:14	1.26	0.88	22.7	324.5	261.6	0.037
Geminids	2010	1.30	0.899	25.0	326.1	262.3	-

Orbital determination by Meteor Head-echo and optical observation is comparable.



RCS (Radar Cross Section) controversy





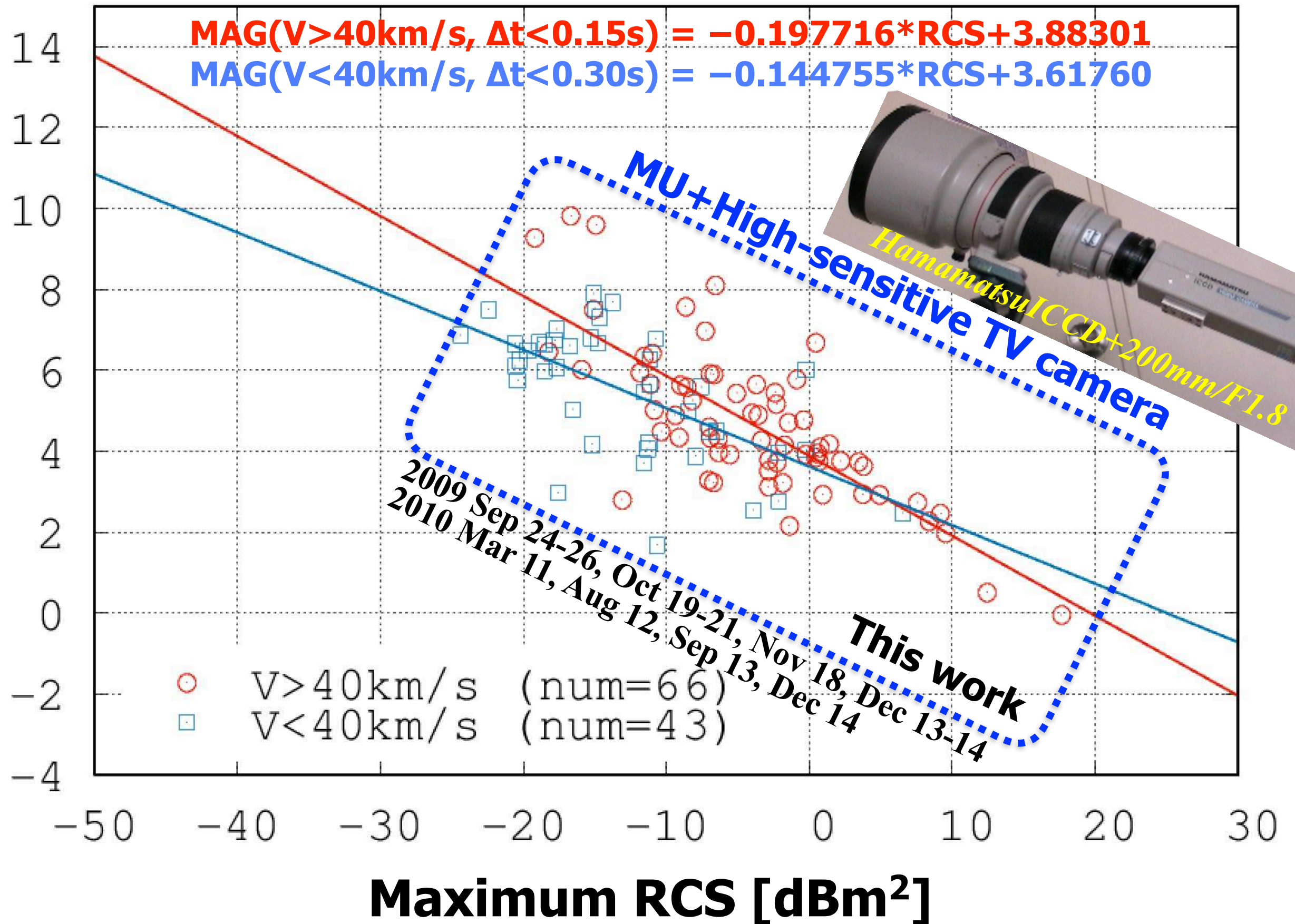
24:17:20 29

*2009/12/15 00:17:21.263(LT) 0014 00032 V00429+102 UFOCaptureV2 NF720D

Simultaneous observation with MU Head-echo and TV

Visual magnitude as functions of RCS

Maximum visual magnitude



Faint Meteors Imaging ~ 13 th magnitude

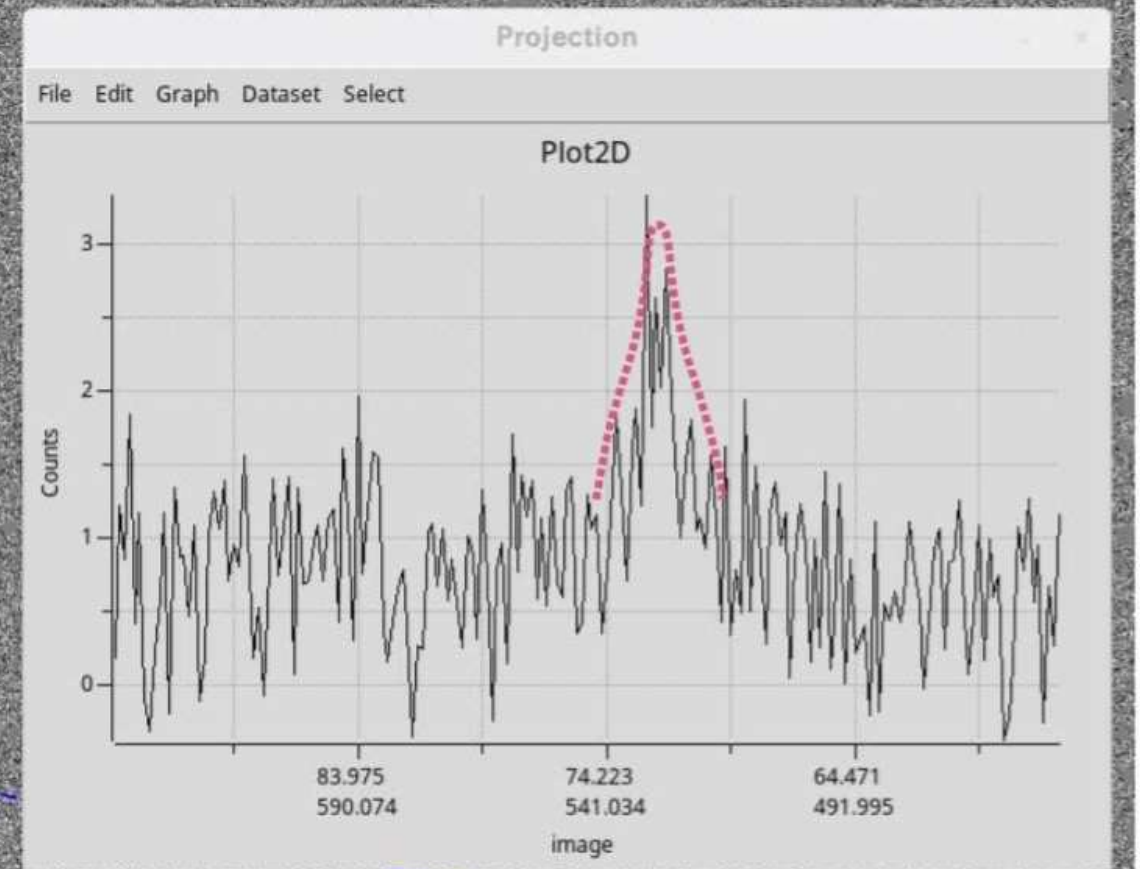
An Image containing **a faint meteor** (stellar sources are masked)

April 11, 2016

1514 events / 5 hours

15 meteors / 180 sec

V Magnitude = 4.5 - 12.5

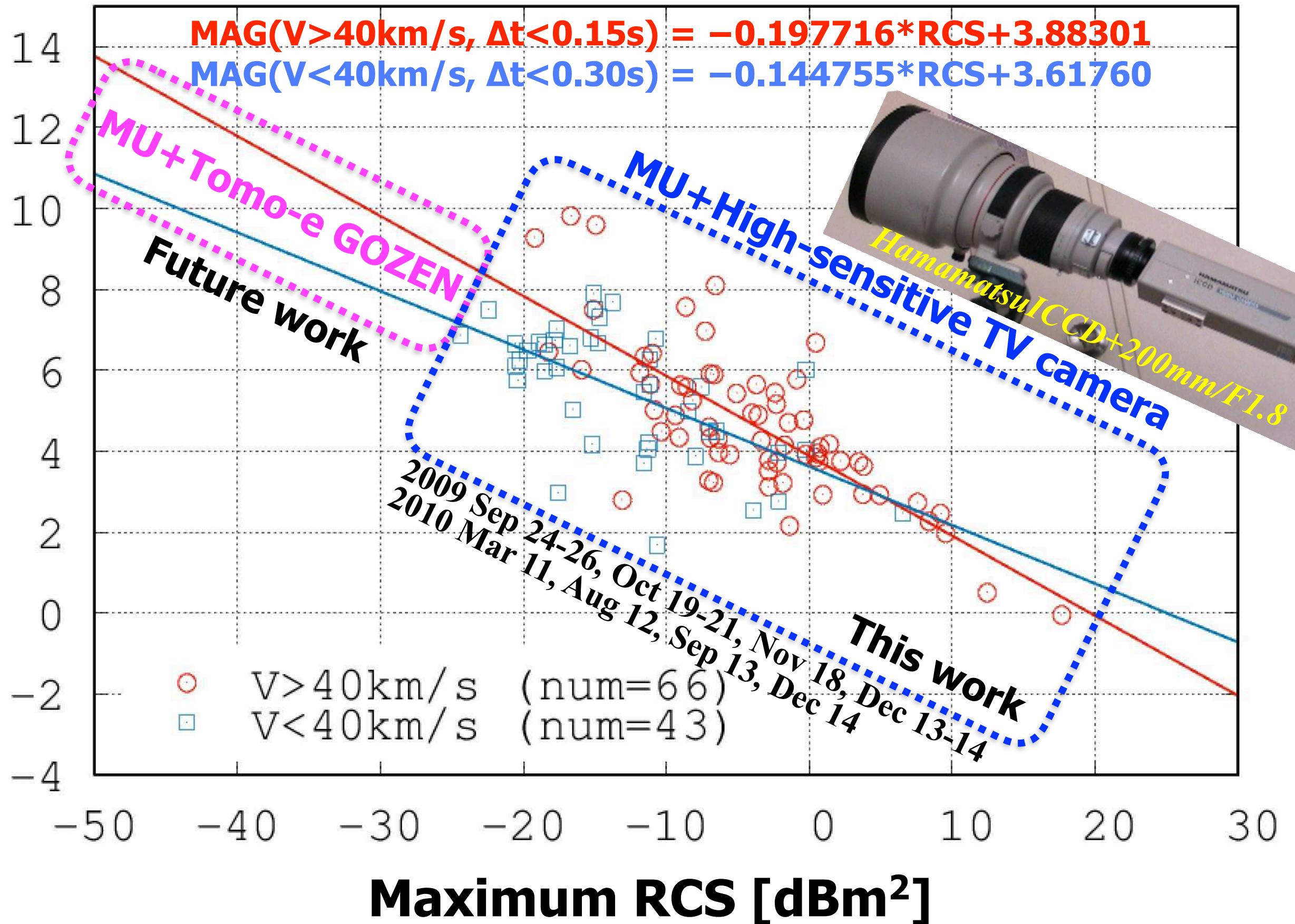


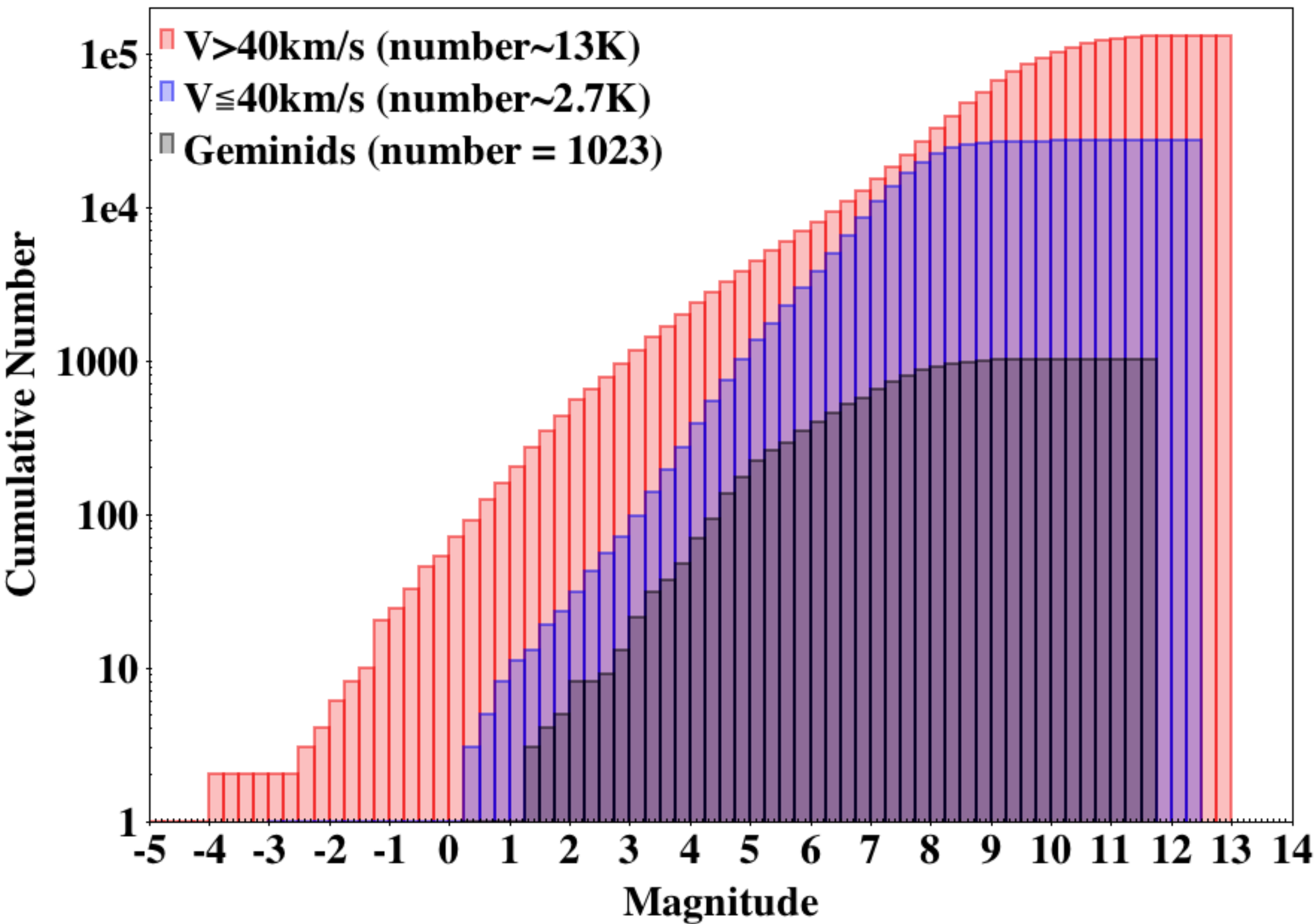
Detected Faint Meteors by Hough transform algorithm

Osawa, Sako, et al. (Univ. Tokyo)

Visual magnitude as functions of RCS

Maximum visual magnitude

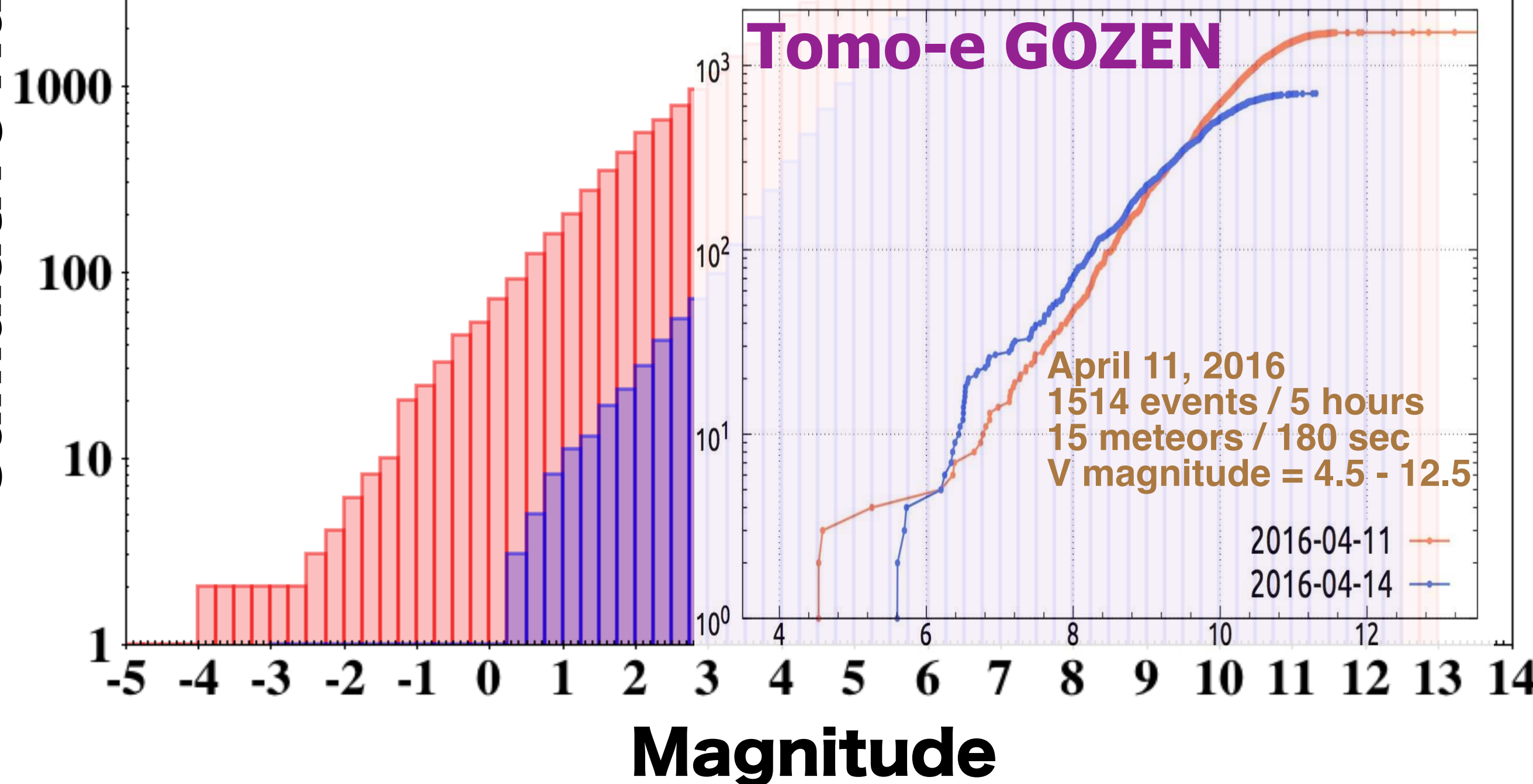




Cumulative Number

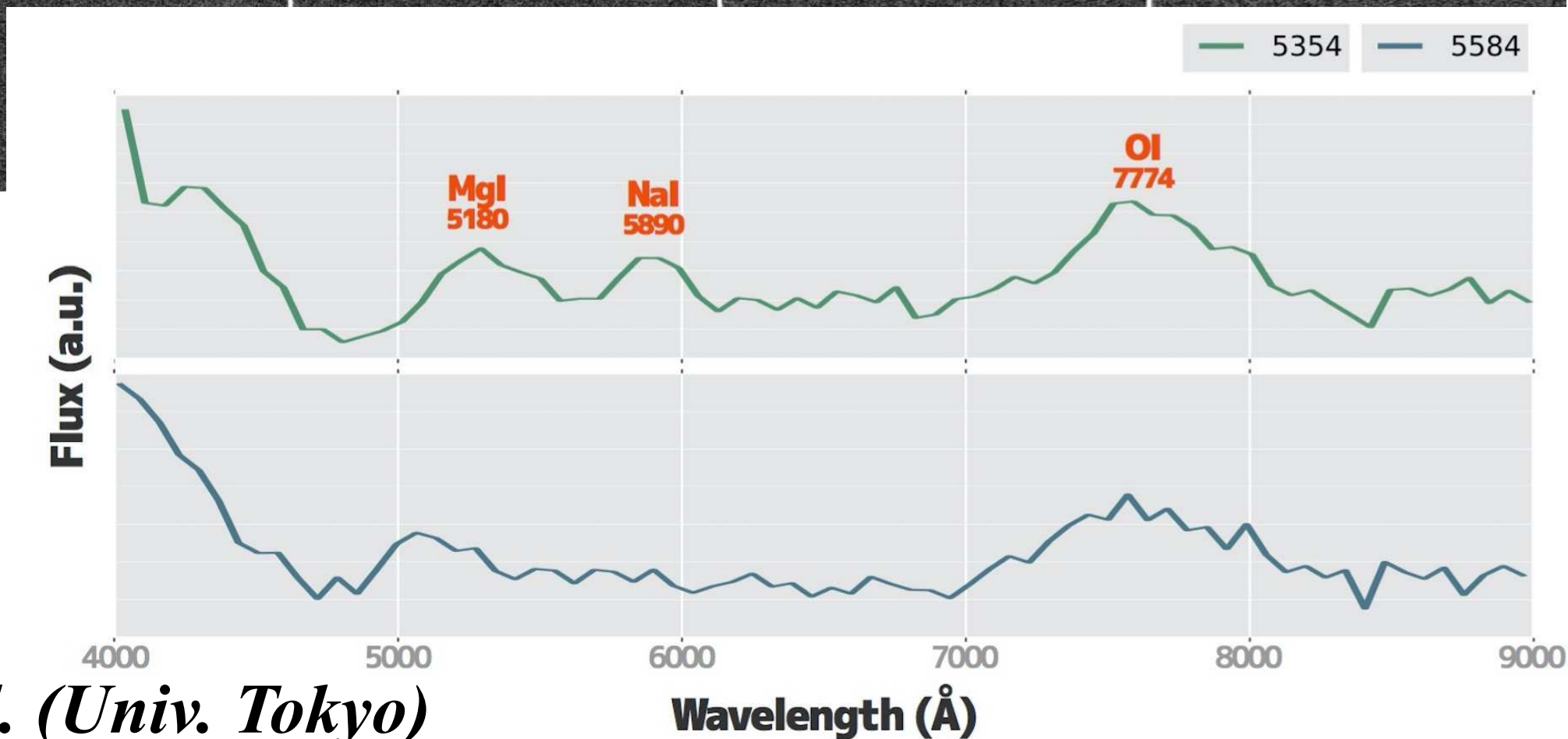
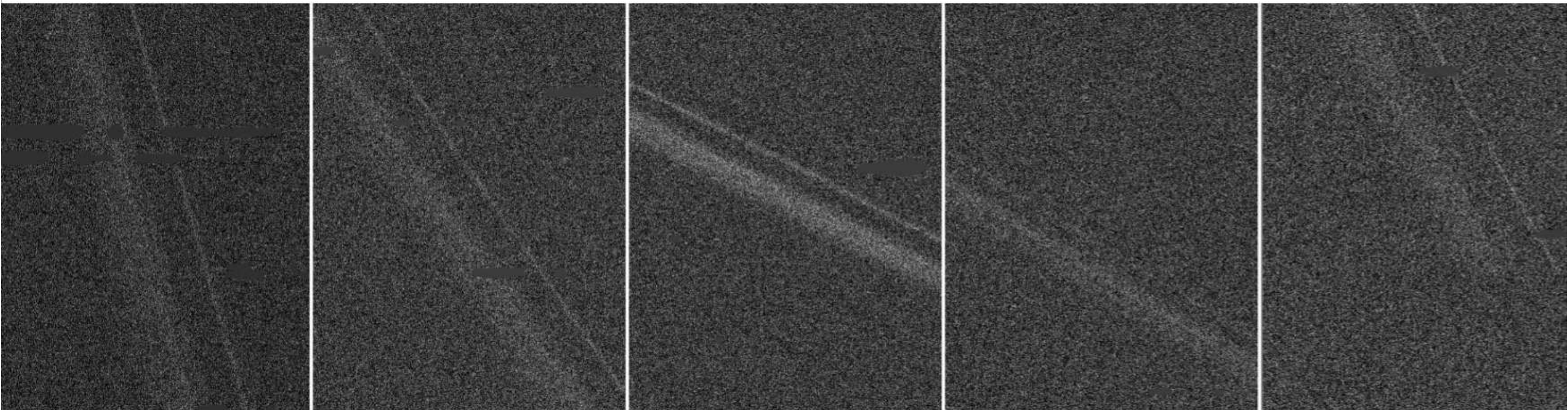
■ $V > 40 \text{ km/s}$ (number $\sim 13 \text{ K}$)
■ $V \leq 40 \text{ km/s}$ (number $\sim 2.7 \text{ K}$)

**Individual observations
by
Radar and Optical methods**



Faint Meteors Spectroscopy

~8-9th magnitude with R=10

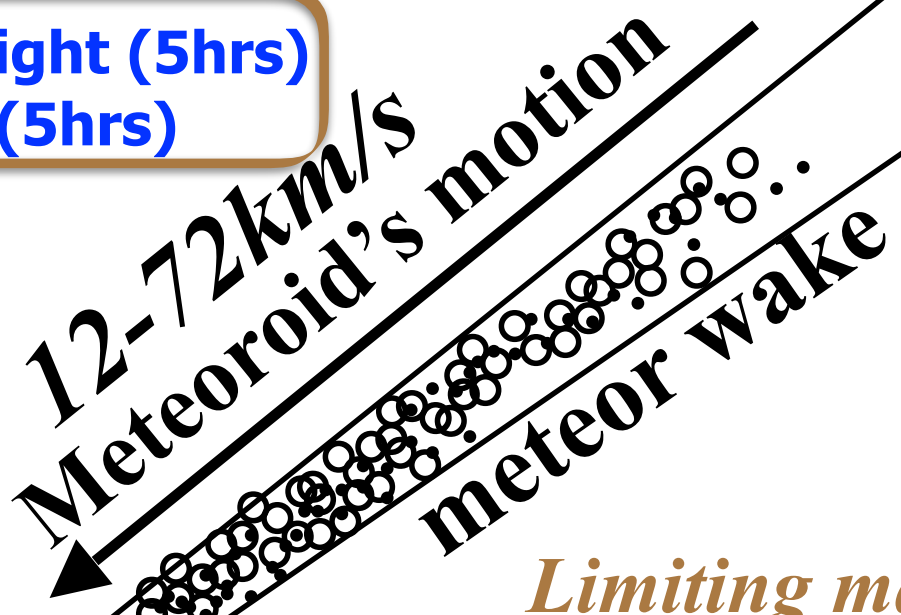


~1000 orbits and size distribution per night (5hrs)
~100 spectroscopy with orbit per night (5hrs)

Future Plan

Meteoroid

Meteor head
Dense plasma



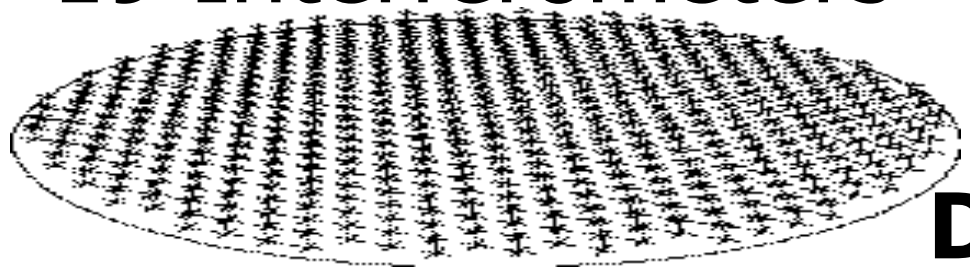
*Limiting magnitude for stars
~19 mag. (2Hz at V-band)*

332Hz
D=103m
FOV ~ Φ 4deg
46.5MHz, 1MW

faint meteors ~13th mag.

2Hz
D=1.05m
FOV ~ Φ 9deg
400-700nm

19 Interferometers



Height; 70-130km

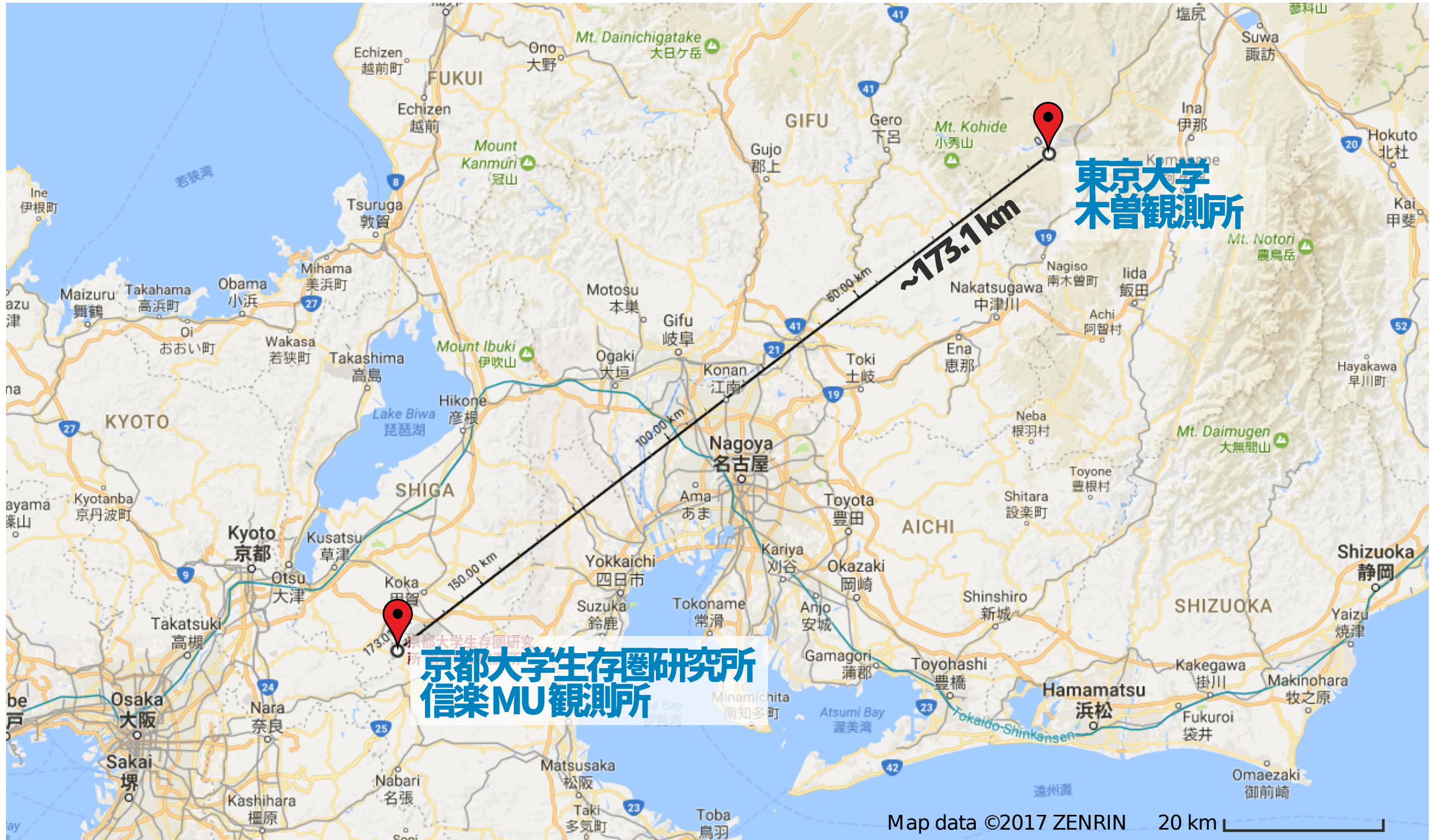
Distance ~173km



MU Radar

Schmidt telescope

木曾観測所と京都大学MUレーダの位置関係



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