

# Supernova Survey with KWFC

KWFCによる  
超新星サーベイ

**Masaomi Tanaka**  
田中 雅臣  
(IPMU, U Tokyo)

**IPMU** INSTITUTE FOR THE PHYSICS AND  
MATHEMATICS OF THE UNIVERSE

NOAO/AURA/NSF  
Bob Ferguson and Richard Desruisseau/Adam Block



# Contents

- What is Supernova?
- Supernova Survey
  - Past and on-going surveys in the world
- Supernova Survey with KWFC
  - Science cases (see also Morokuma-san's talk)
- Summary

# Type Ia supernova (Thermonuclear SN)

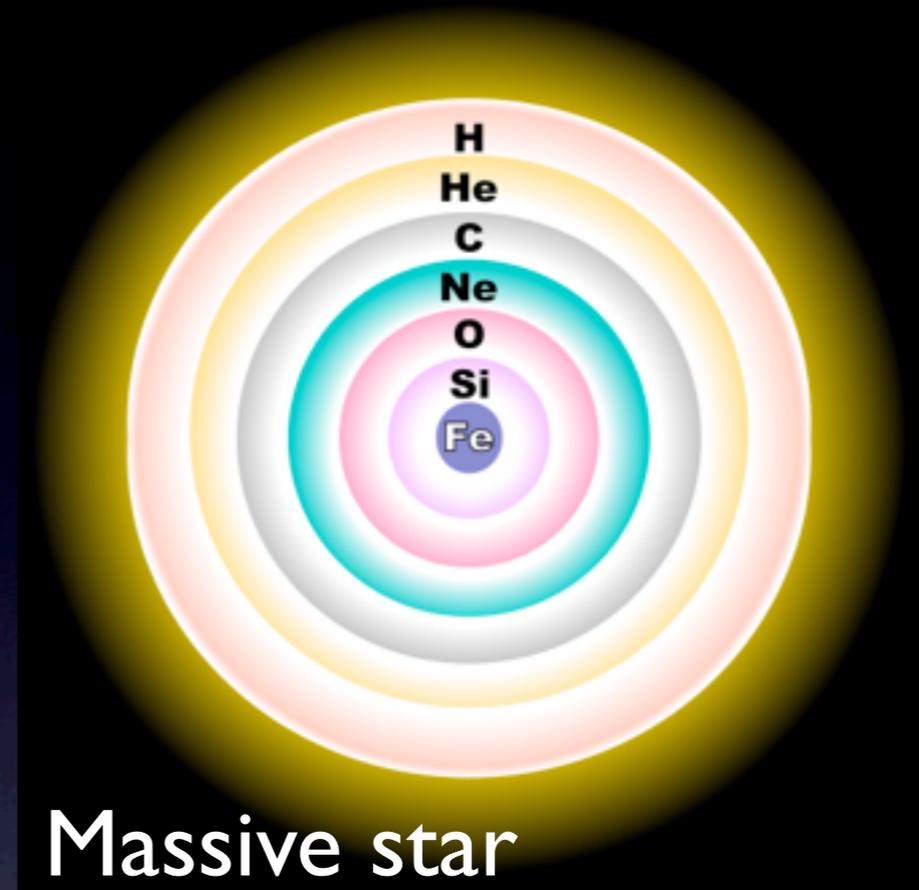


Standard candle

Progenitor system

Nucleosynthesis  
(Fe, Si)

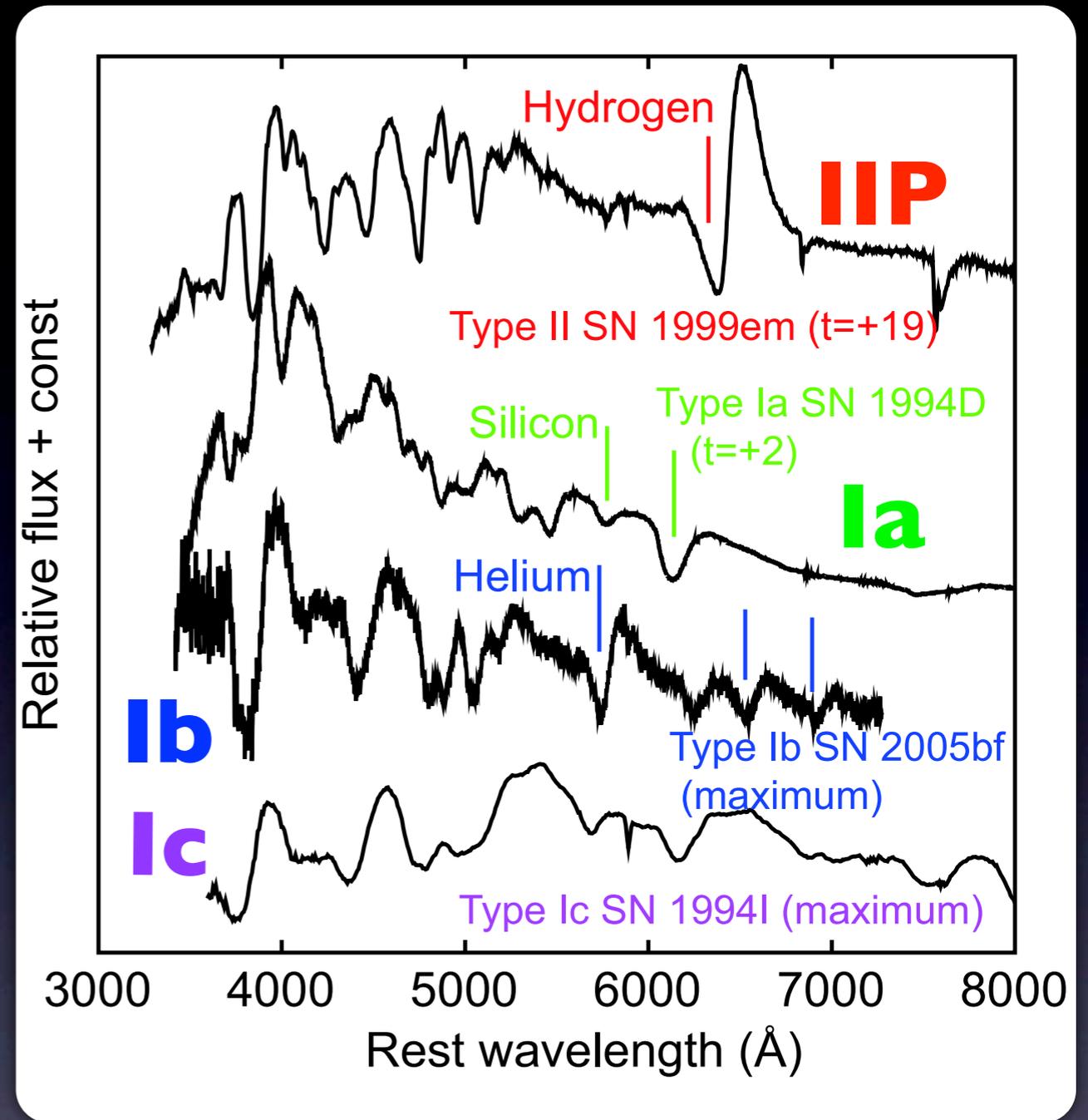
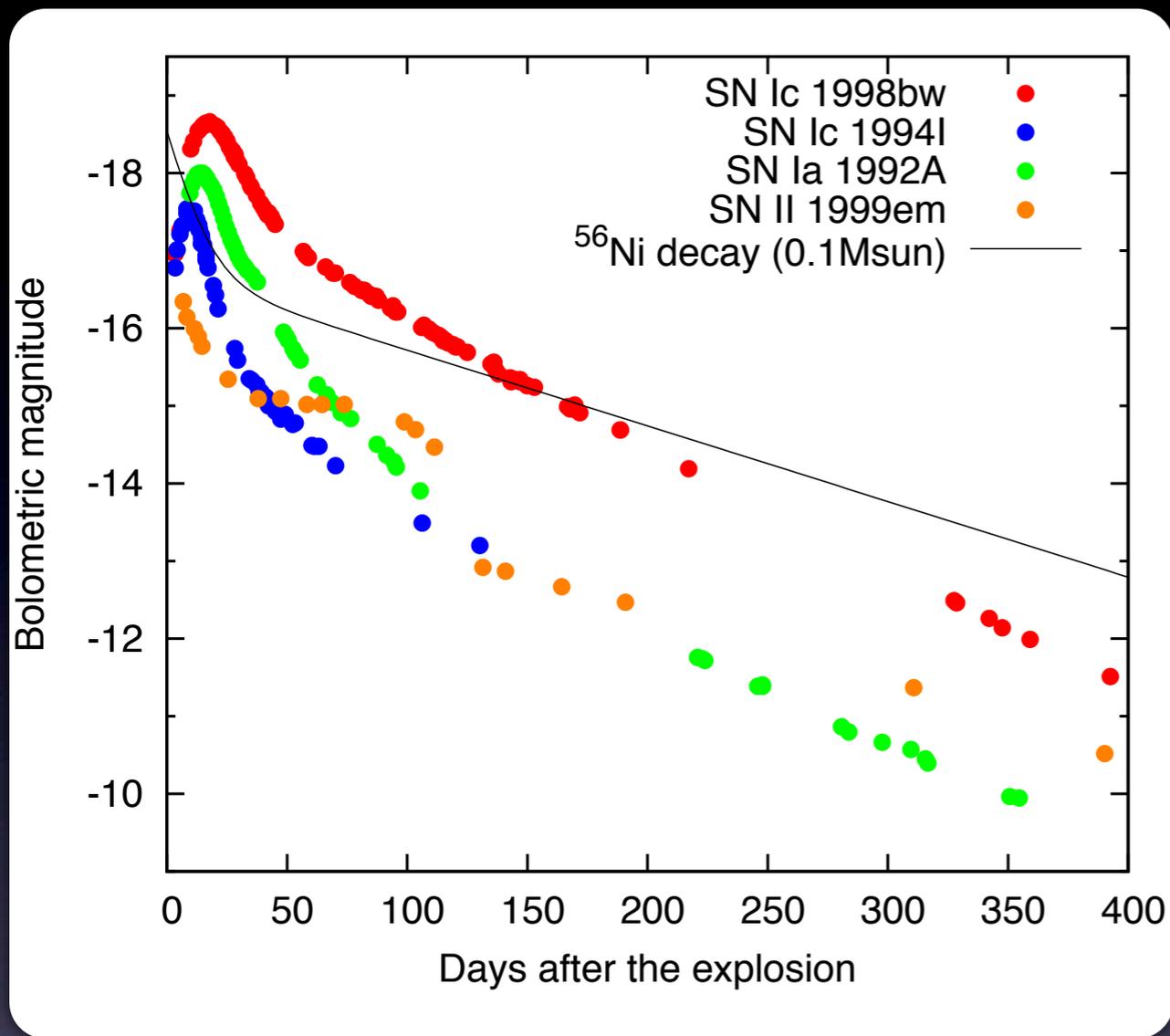
# Core-collapse supernovae



Explosion mechanism

Progenitor star

Nucleosynthesis  
(C, O, Ne, Mg)

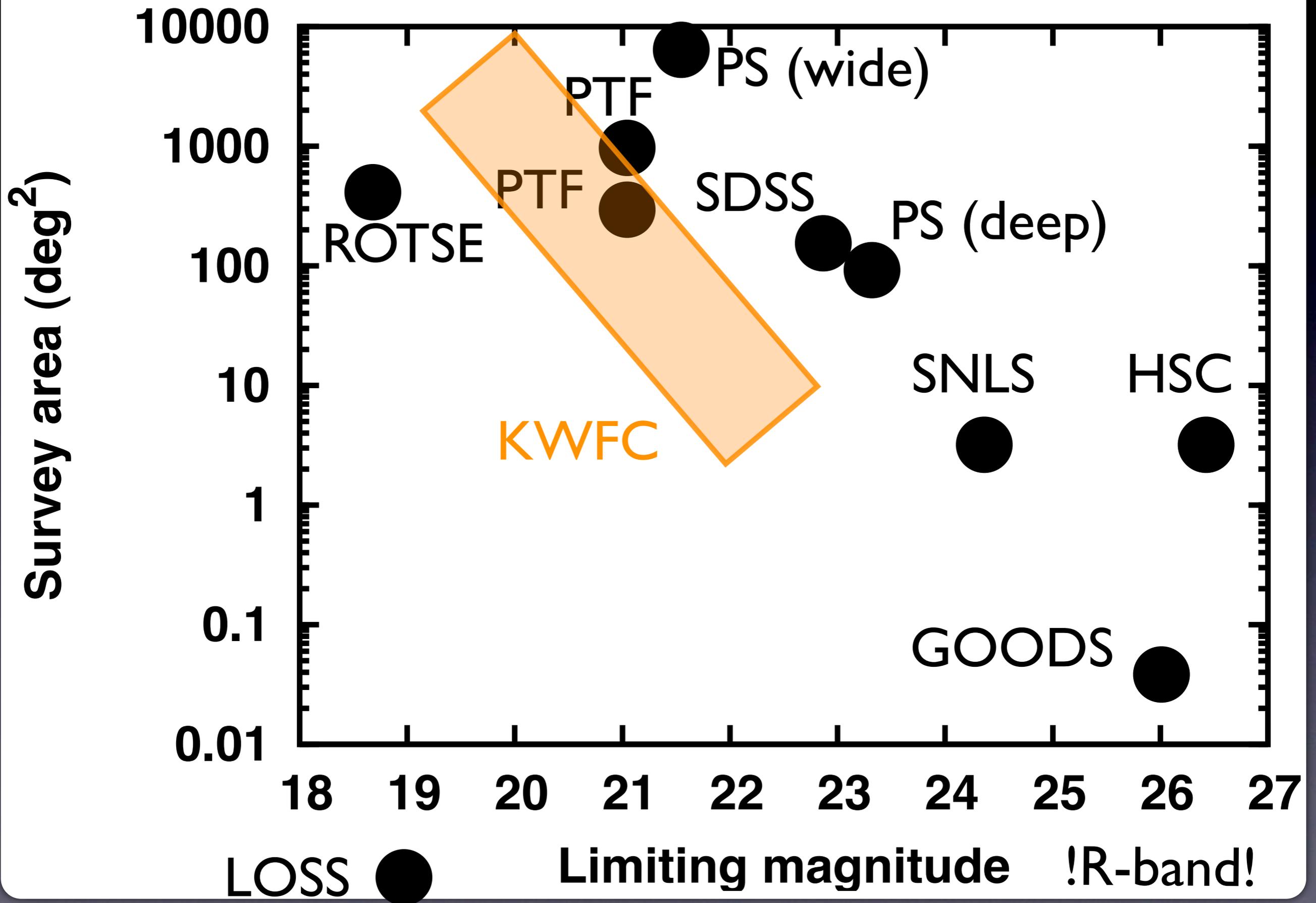


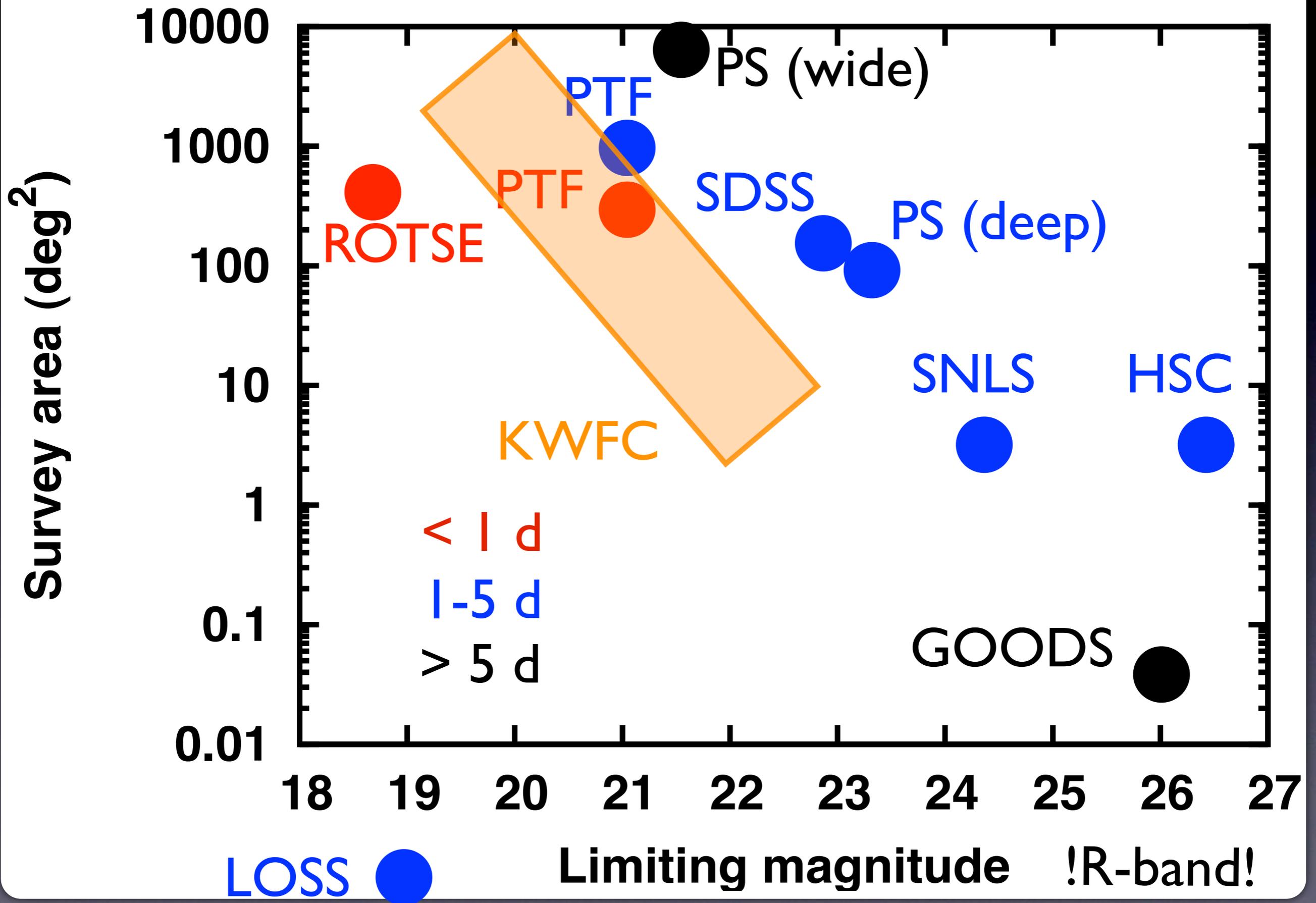
Survey with  $\sim 21$  mag depth  
 Type Ia SNe @  $z \sim 0.2$  ( $\mu \sim 40$ ,  $d \sim 1$  Gpc)  
 Core-collapse SNe @  $z \sim 0.05$  ( $\mu \sim 37$ ,  $d \sim 200$  Mpc)

# Supernova Survey

Survey	Diameter (m)	FOV (deg <sup>2</sup> )	Depth (R mag)	Area/day (deg <sup>2</sup> )
LOSS	0.76	0.01	19	1000 galaxy
ROTSE-III	0.45	3.42	18.5	450
PTF	1.26	7.8	21	1000
Pan-STARRS	1.8	7	21.5	6000
SDSS-II	2.5	1.5	22.6	150
SNLS	3.6	1	24.3	2
GOODS	2.5 (HST)	0.003	26	0.04
HSC	8.2	1.75	26.5	1.75
KWFC	1.08	4		

(partly taken from Rau et al. 2009, PASP, 121, 1334)

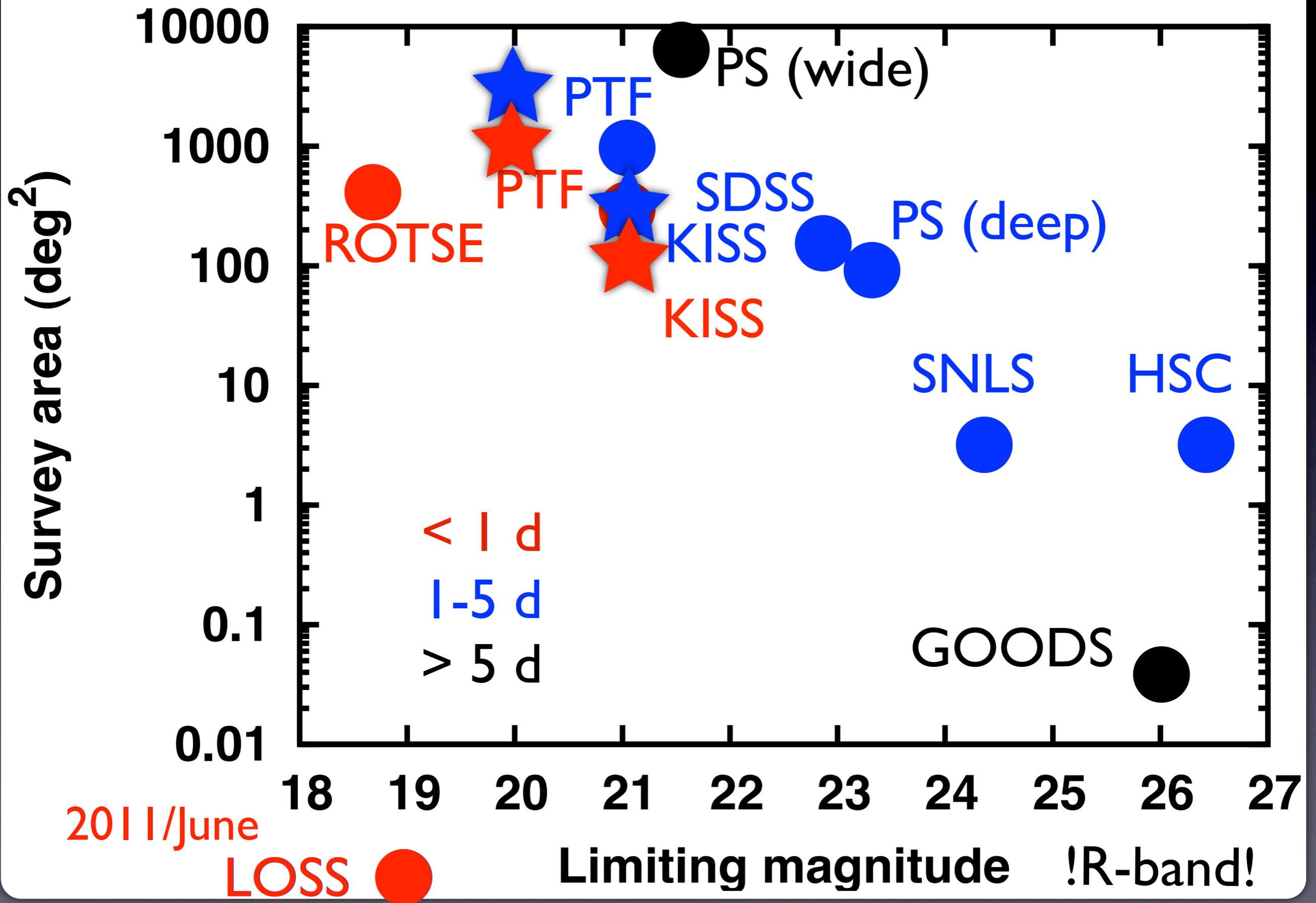




# SN Survey with KWFC

## **KISS**: **K**iso **S**upernova **S**urvey

- Survey in blue band  
(B or g, possibly + U or u)
- 3-day cadence
  - 15 min exp. (B/g~22 mag) => ~300 deg<sup>2</sup>
  - 1.5 min exp. (B/g~21 mag) => ~2700 deg<sup>2</sup>
- 1-day cadence
  - 15 min exp. (B/g~22 mag) => ~100 deg<sup>2</sup>
  - 1.5 min exp. (B/g~21 mag) => ~900 deg<sup>2</sup>



# Unsolved Problem - Type Ia SNe -



David A. Hardy

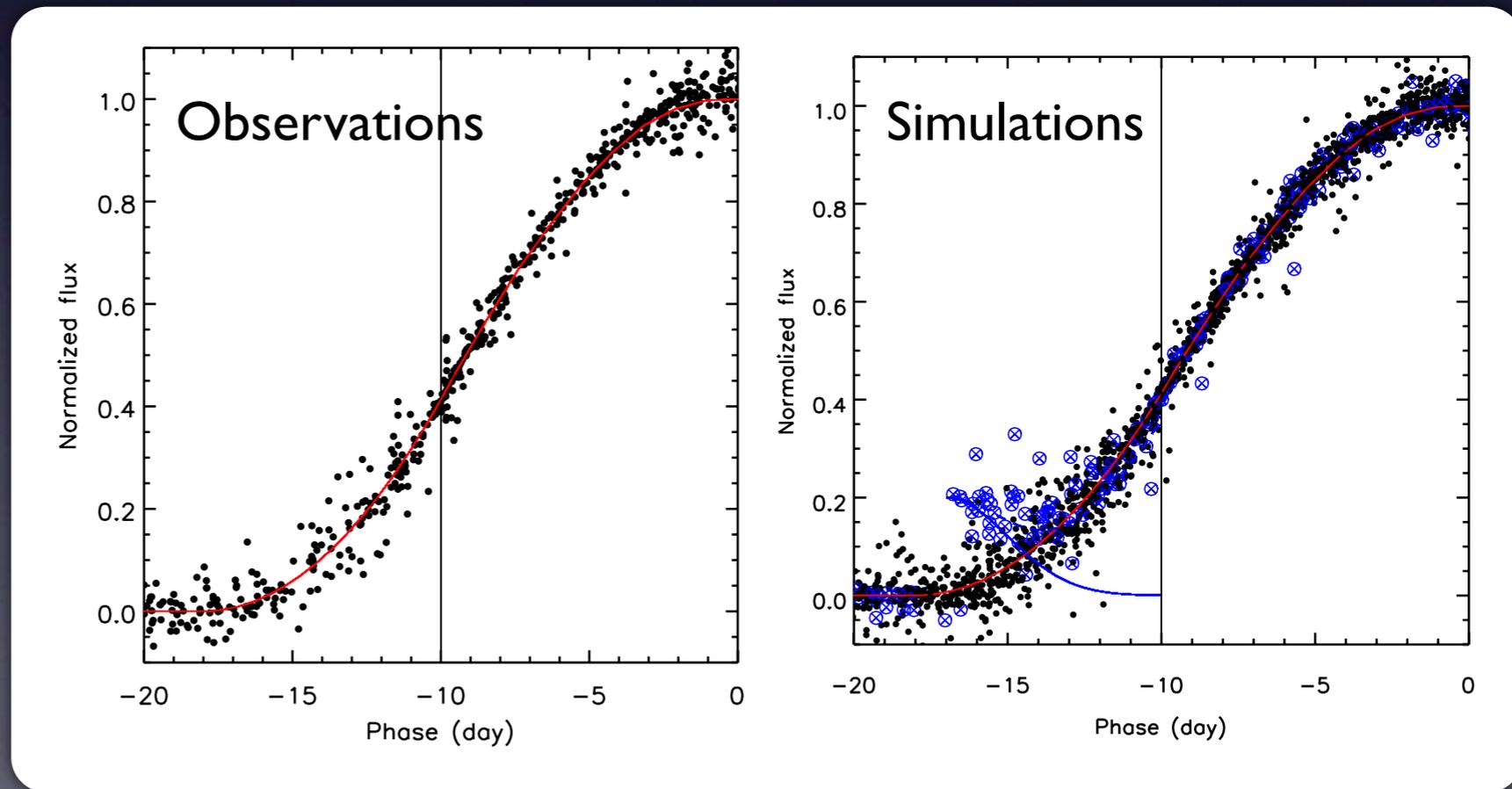
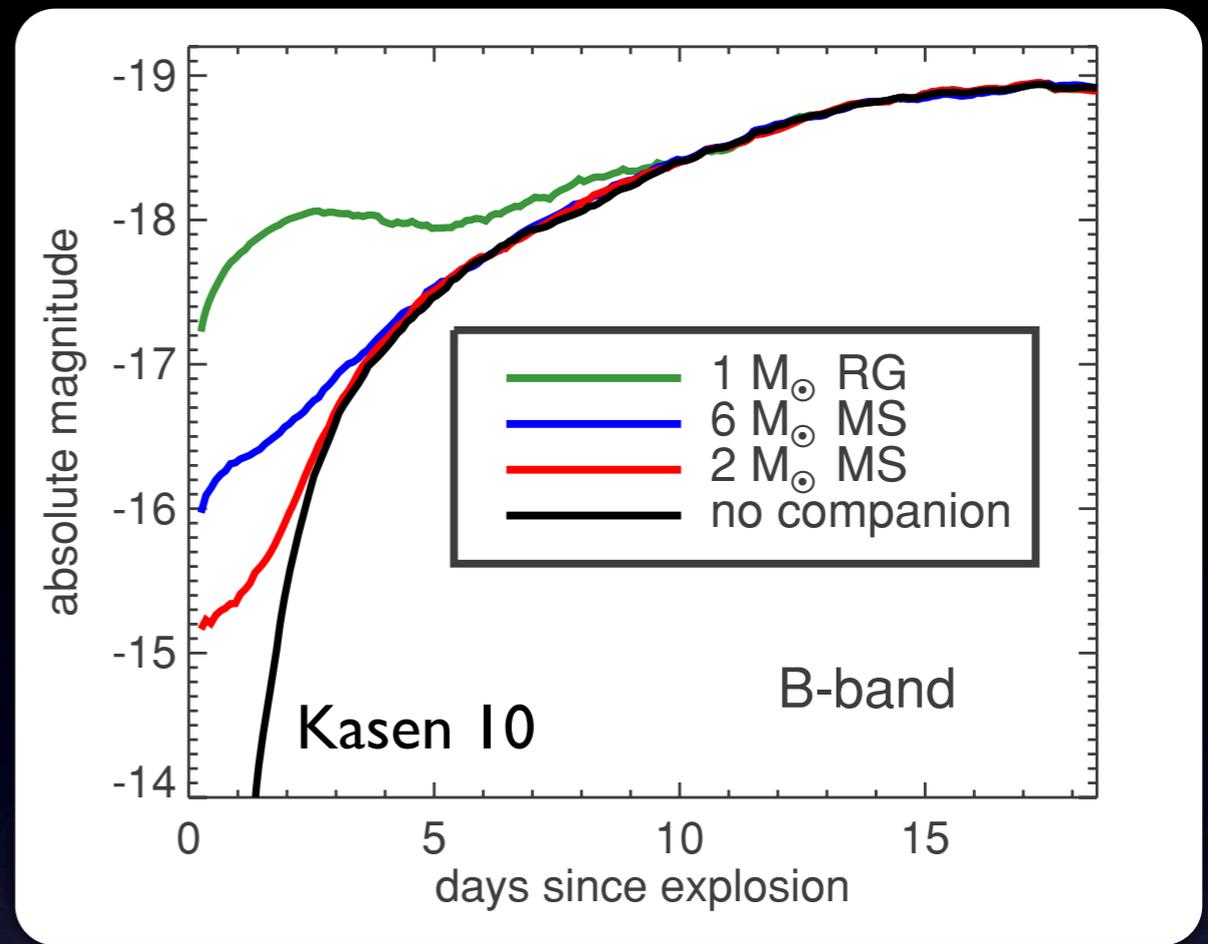
V.S.



GSFC/D. Berry

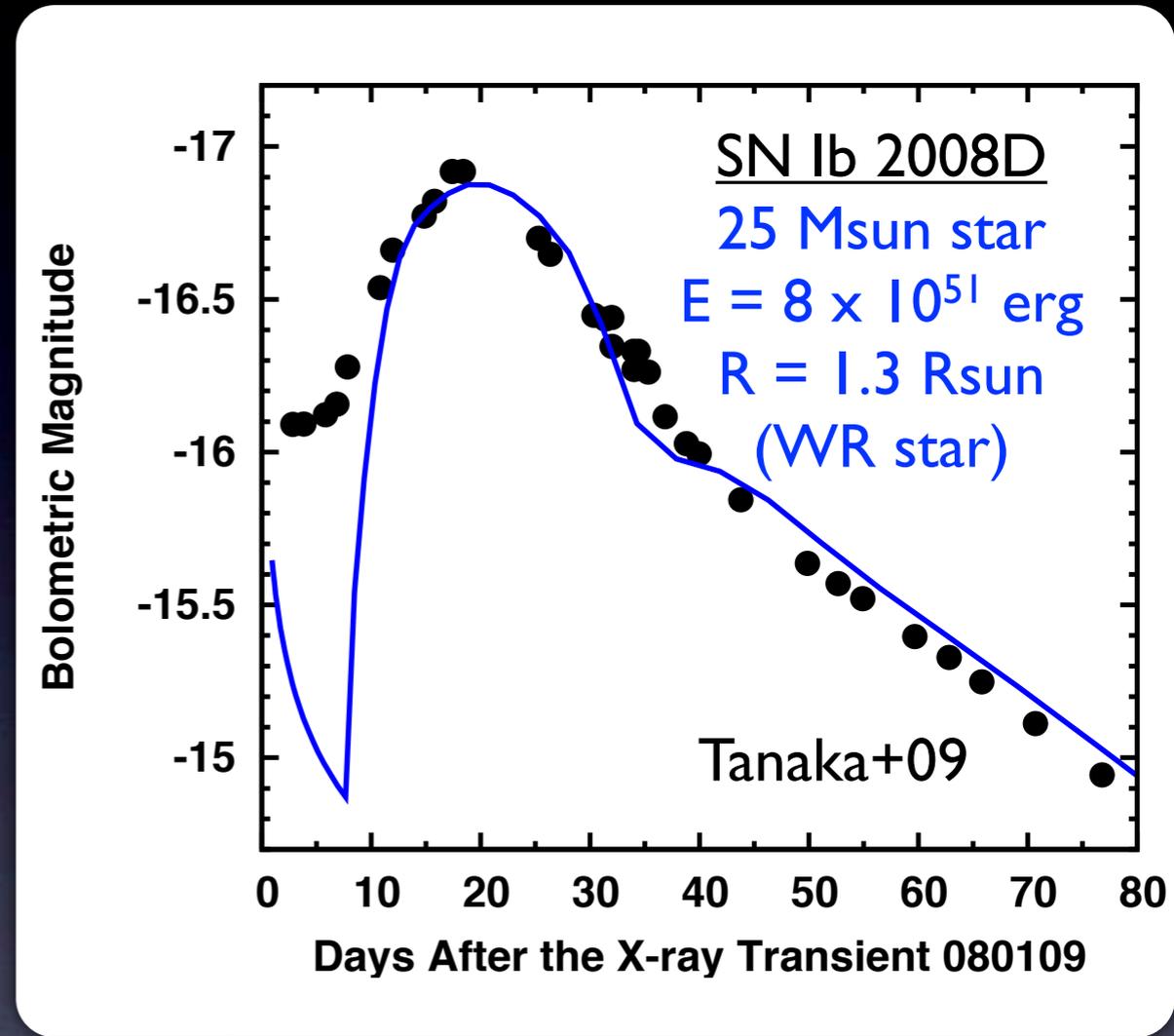
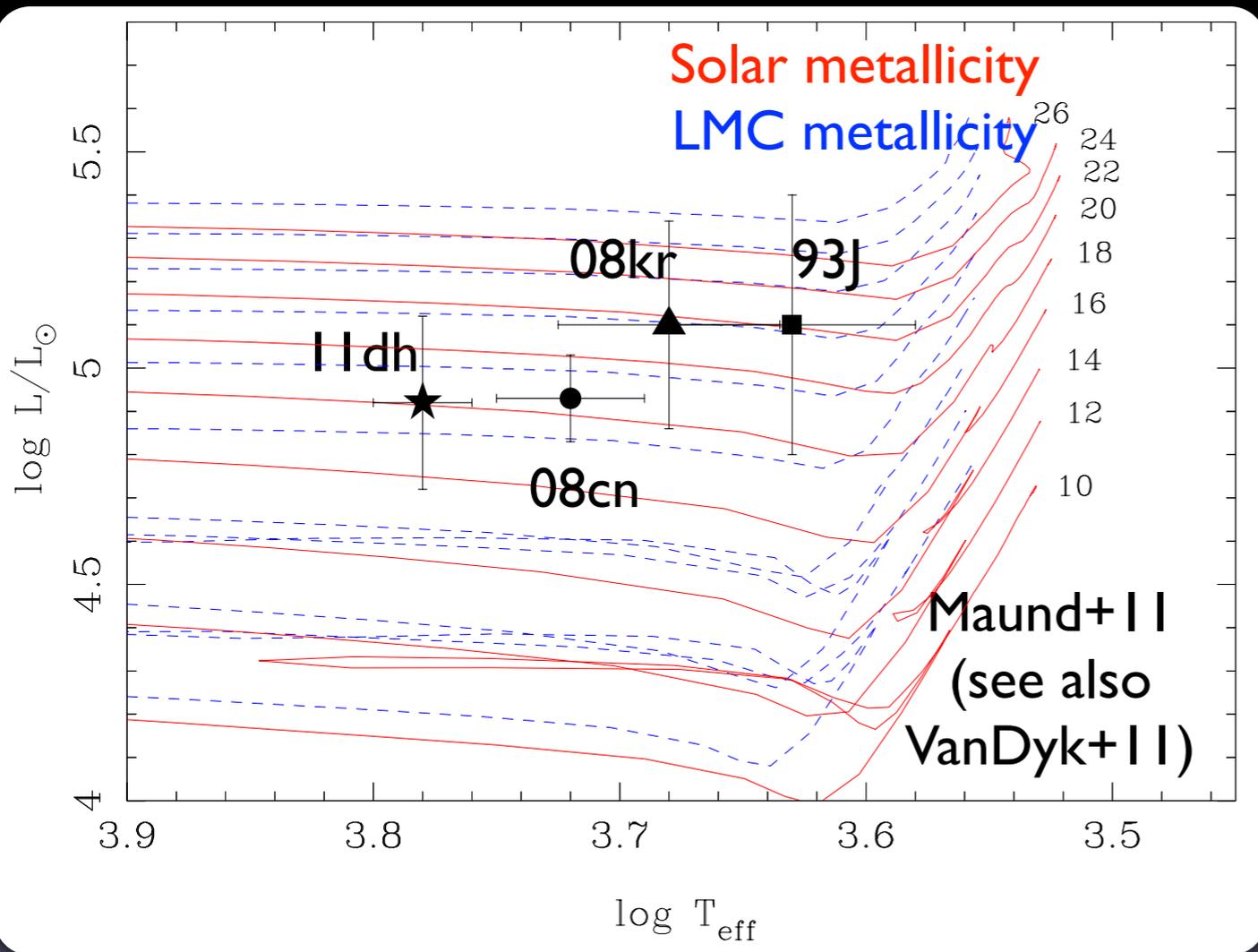
加藤拓也さん講演

(2010年シュミットシンポ)



Hayden+10 (SDSS), see also Bianco+11 (SNLS)

# Unsolved Problem - Core-Collapse SNe -



Progenitor star of  
extremely nearby SNe

**Radius??**

$$L_c = 3.3 \times 10^{42} E_{51}^{0.91} M_0^{-0.74} R_{12} (F_1/1.35)^{-0.17} t_d^{-0.34} \text{ ergs s}^{-1}$$

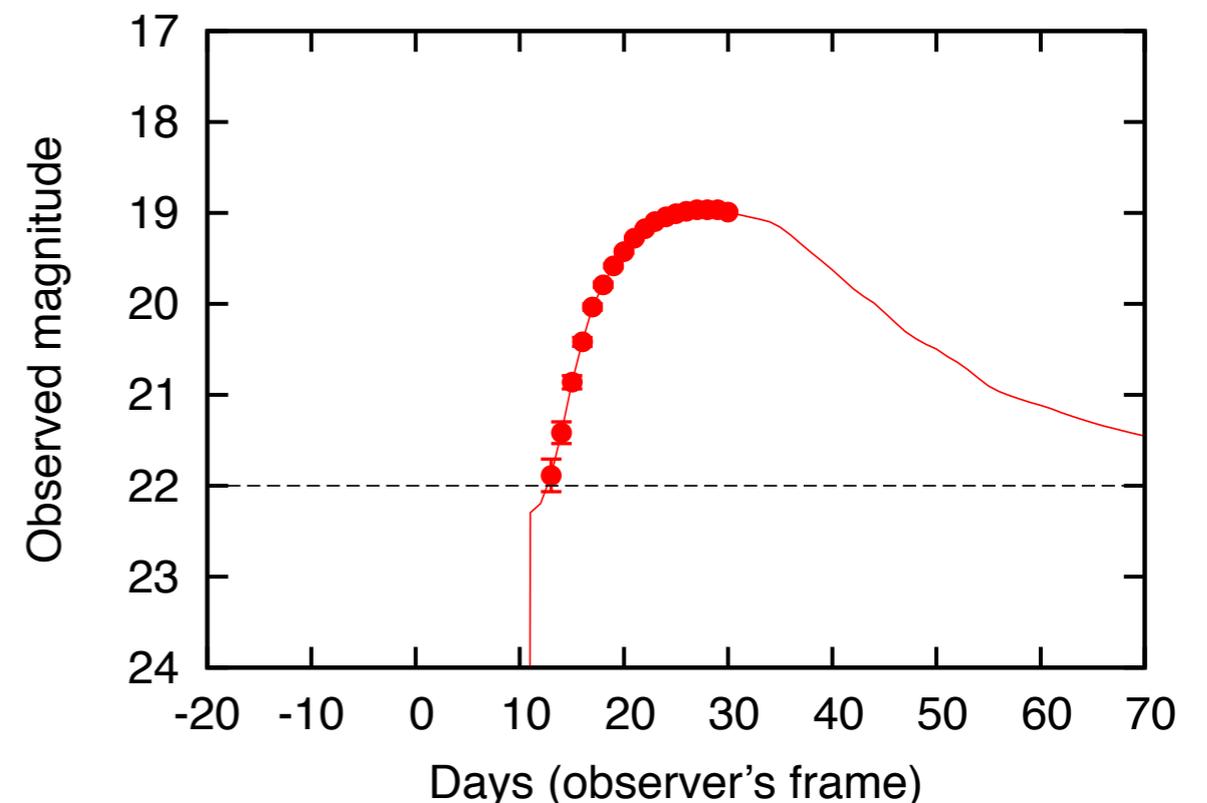
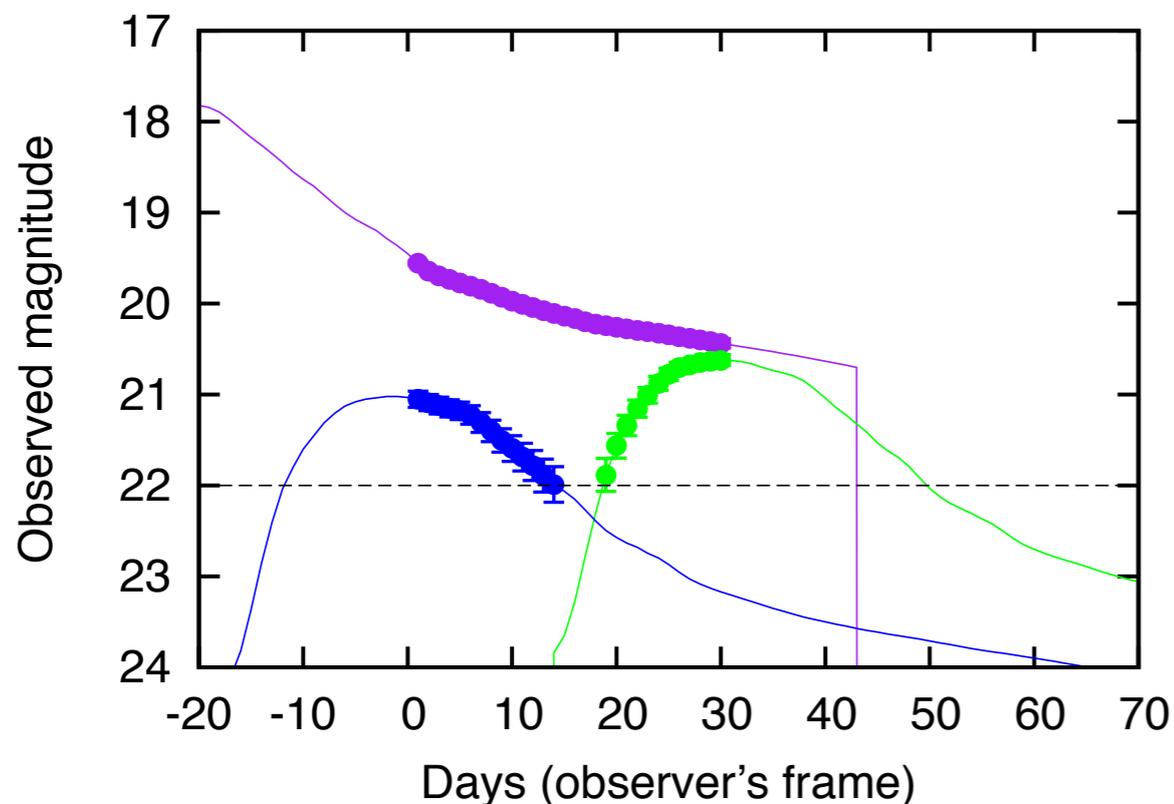
# Expected Number of SN Detection

1-month survey  
(1-day cadence, B-band)

Number in () = Number of discovery  
at < 5 days after the explosion

	Exp. time (min)	Limiting mag (B/g)	Survey area (deg <sup>2</sup> )	# of SN Ia	# of SN IIP	# of SN Ibc
Deep	15	22	100	110 (10)	6 (0.18)	2 (0.04)
Wide	1.5	21	900	300 (18)	15 (0.45)	5 (0.025)

## Example of Type Ibc SNe





**KANATA**  
(Opt. spec)

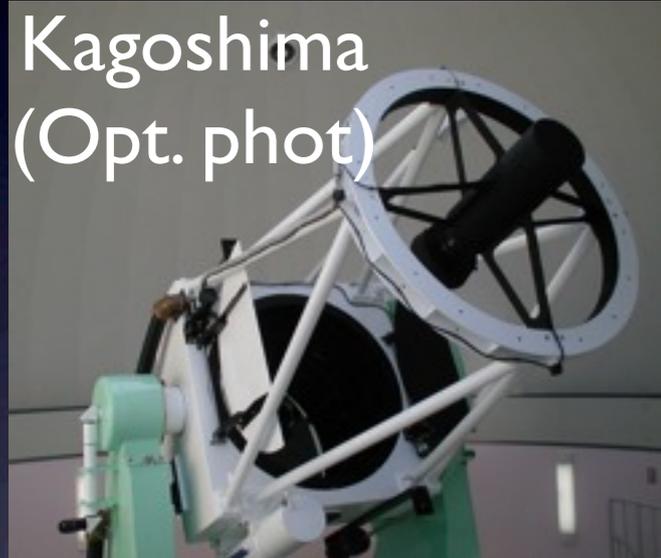
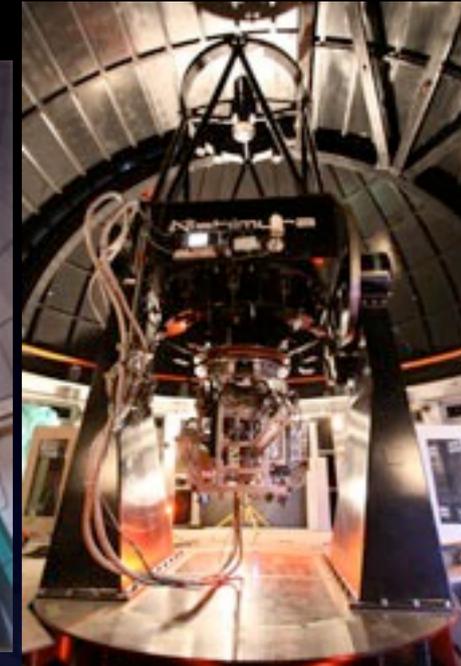


**Pirika**  
(Opt. spec)

**IRSF**  
(NIR phot)



**miniTAO**  
(NIR phot)



**Kagoshima**  
(Opt. phot)

**KISO/KWFC**  
(Opt. survey)



**OA0188**  
(NIR phot/spec)



**Subaru**  
(Any)

**MITSuME**  
(Opt. phot)



# KISS: KIso Supernova Survey

- **High cadence!!**
  - Competition in the “time” domain
- Catch the very moment of SN explosion
  - Radius of the massive stars at the very final stage of the evolution
  - Progenitor of Type Ia SNe
- Expected Results with KISS (# in 1 month)
  - Deep: (=> Morokuma-san’s talk)
    - \* Core-collapse SNe ~ 10 (~0.2 very early detection, 0.04 SNe Ibc)
    - \* Type Ia SNe ~ 100 (~10 very early detection)  
*(!!strongly affected by the bad weather!!)*
  - Wide:
    - \* Core-collapse SNe ~ 20 (~0.5 very early detection, 0.025 SNe Ibc)
    - \* Type Ia SNe ~300 (~20 very early detection)