

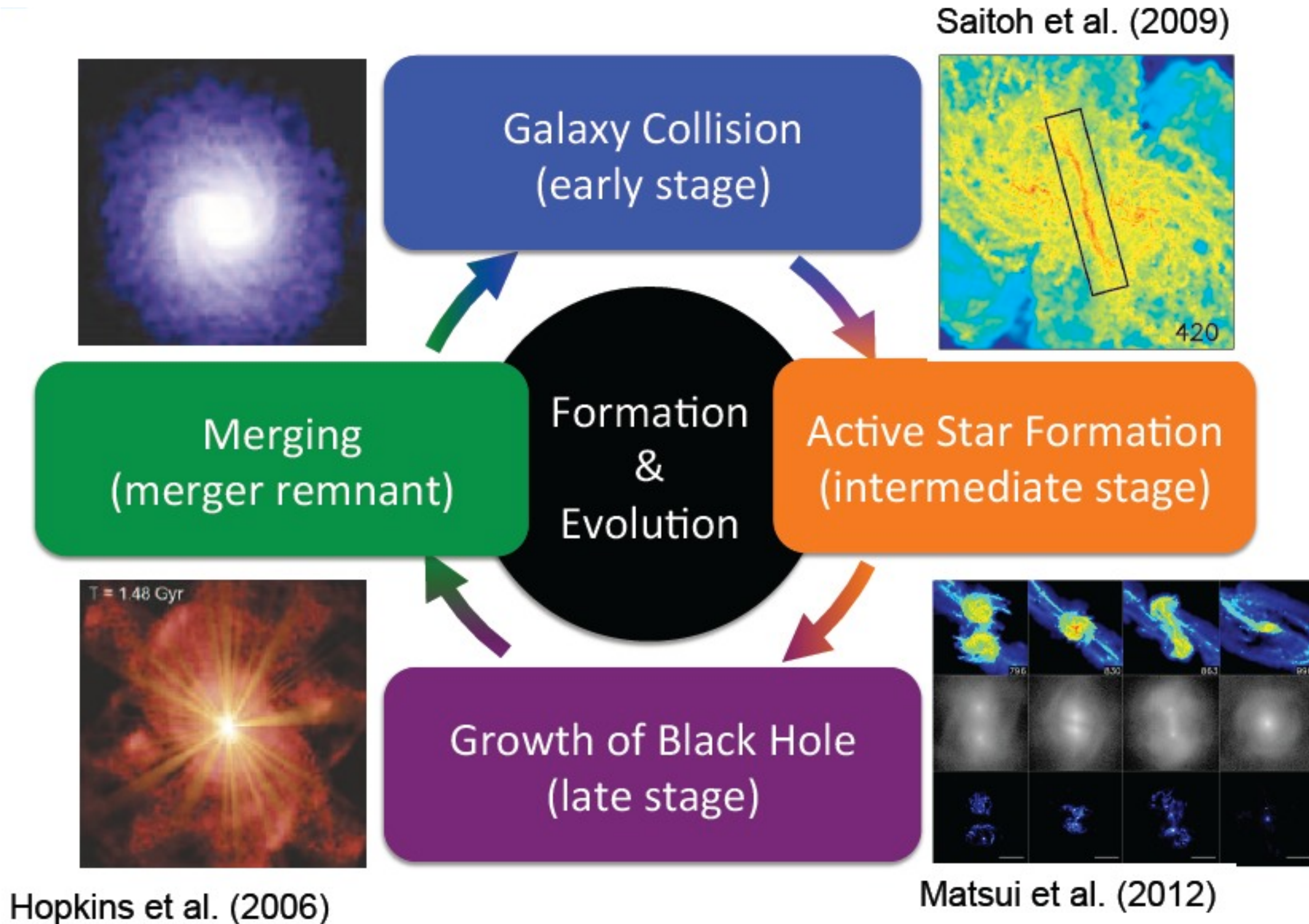
SWIMSとALMAの分光観測 で探る近傍衝突銀河の星形成

Toshiki Saito
University of Tokyo/NAOJ

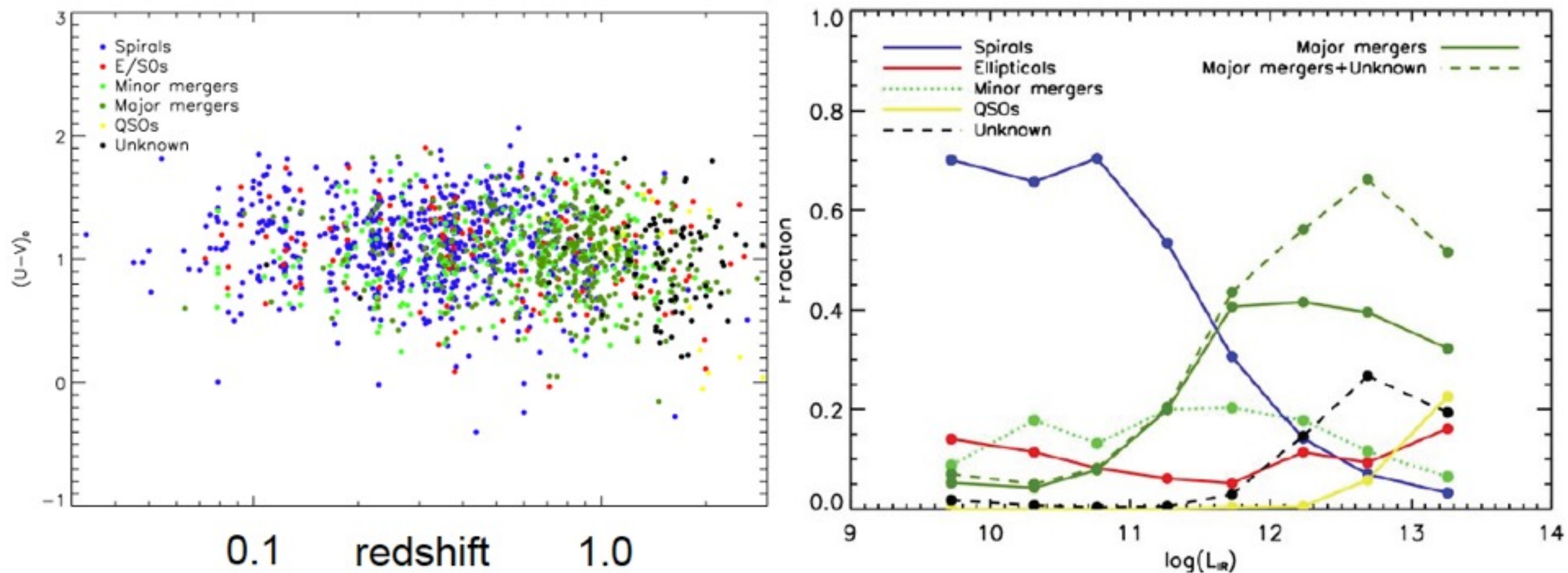


Daisuke Iono, Min S. Yun, Junko Ueda, Daniel Espada, Yoshiaki Hagiwara, Masatoshi Imanishi, Kentaro Motohara, Kouichiro Nakanishi, Hajime Sugai, Ken Tateuchi, and Ryohei Kawabe

Major merger evolution

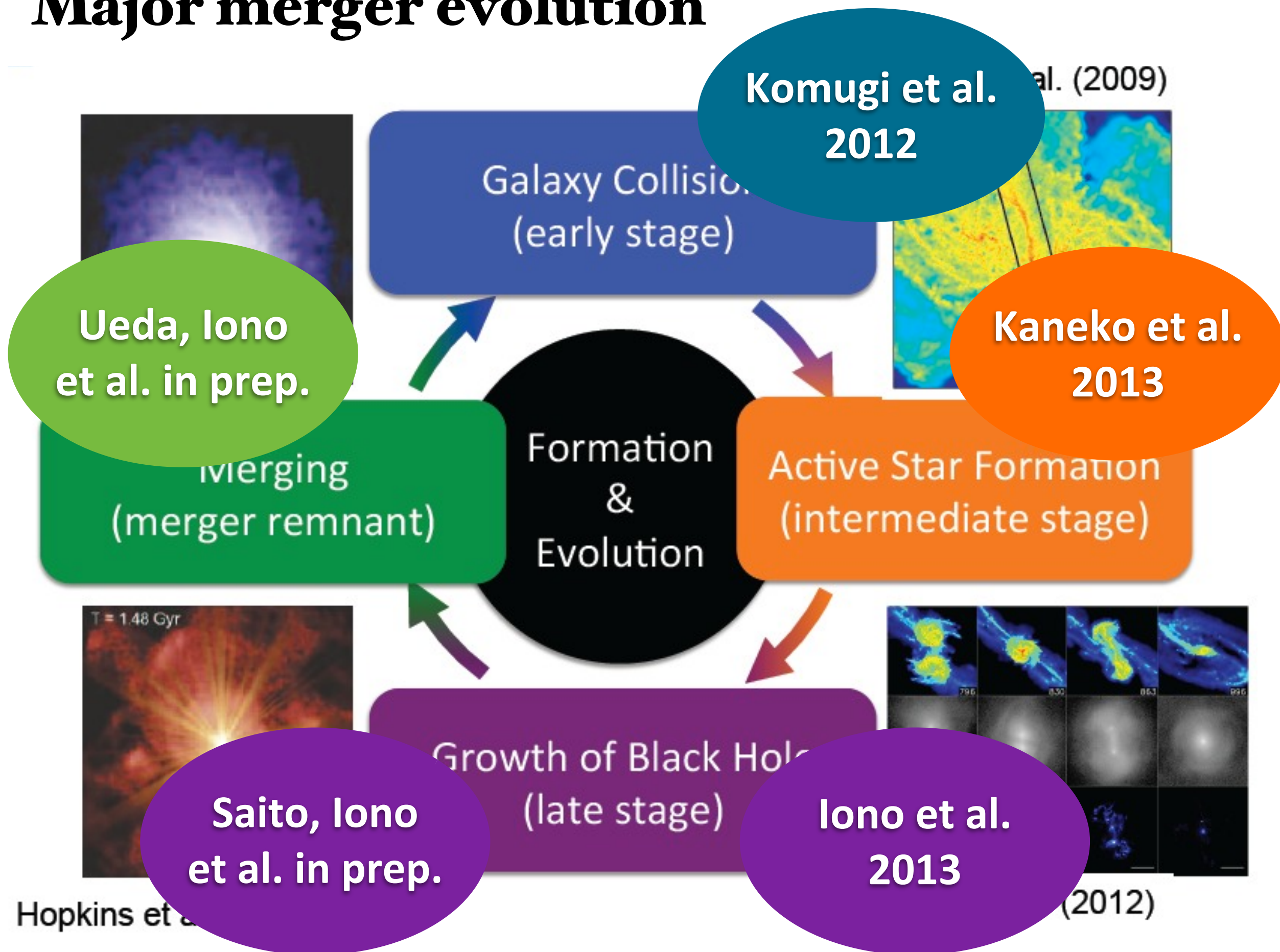


Pair fraction in the Universe



1503 galaxies (70 micron selected) in the $0.1 < z < 3.5$ universe
using COSMOS (Kartaltepe et al. 2010)

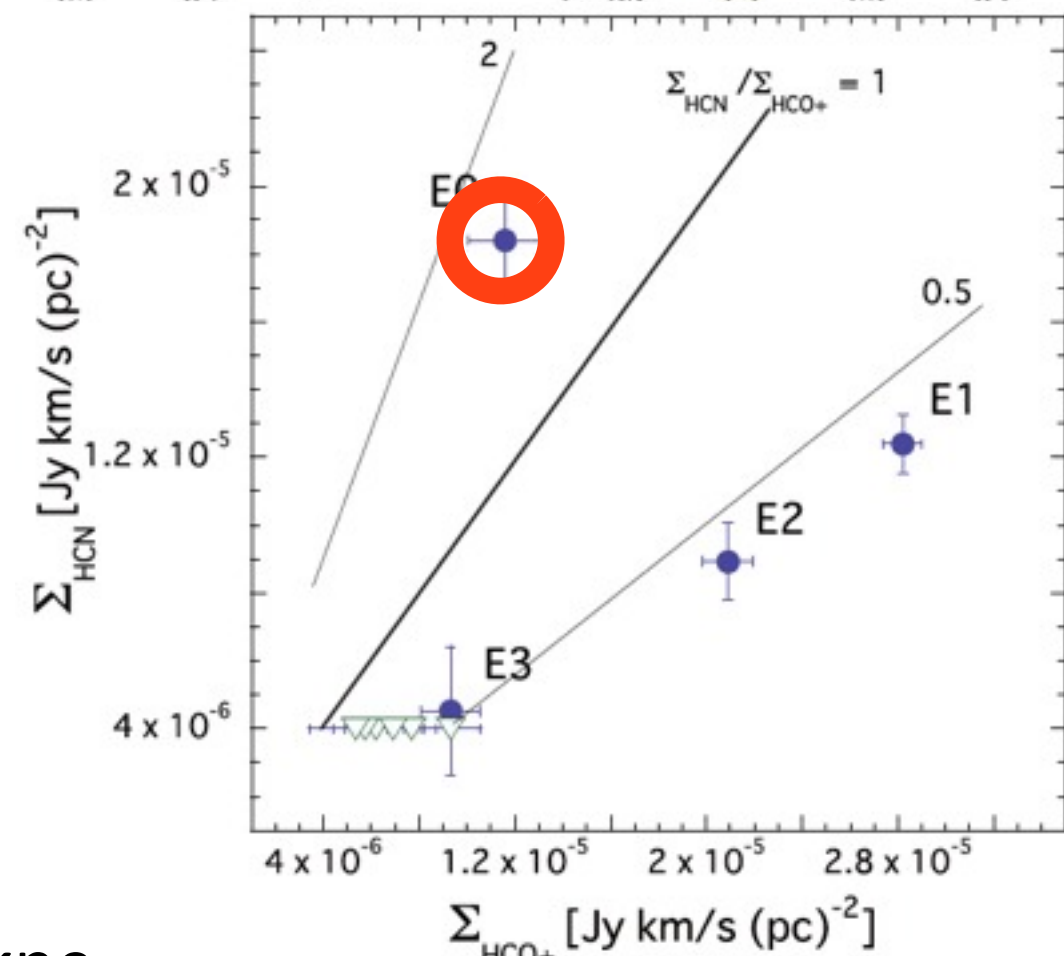
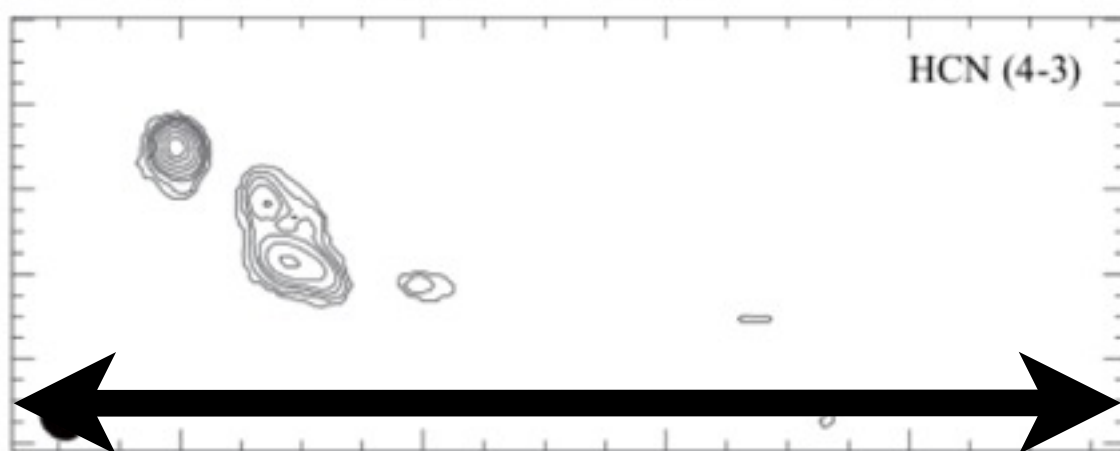
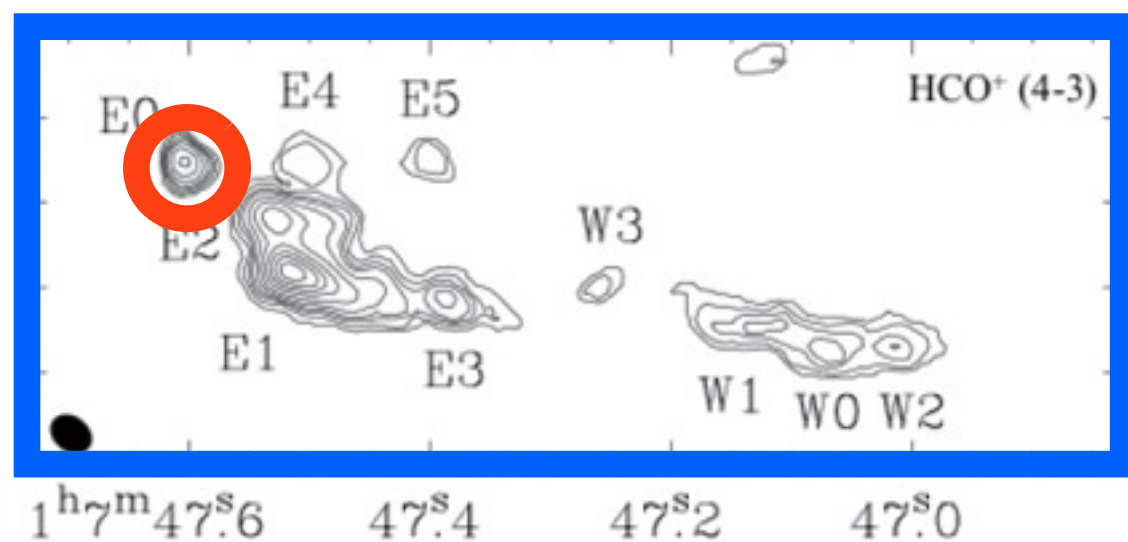
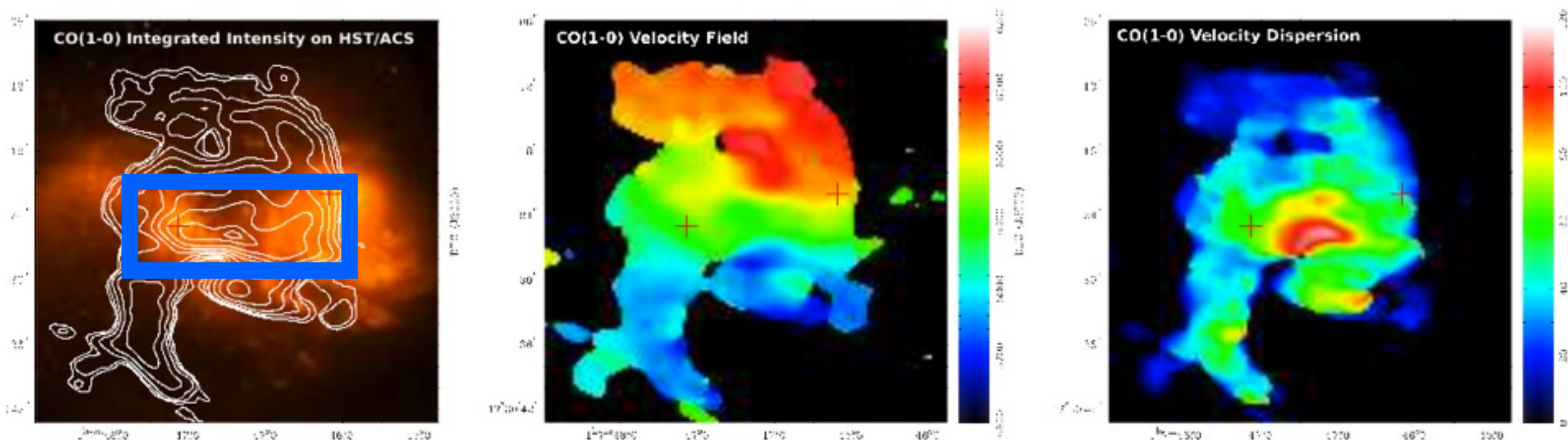
Major merger evolution



The late-stage LIRG VV114

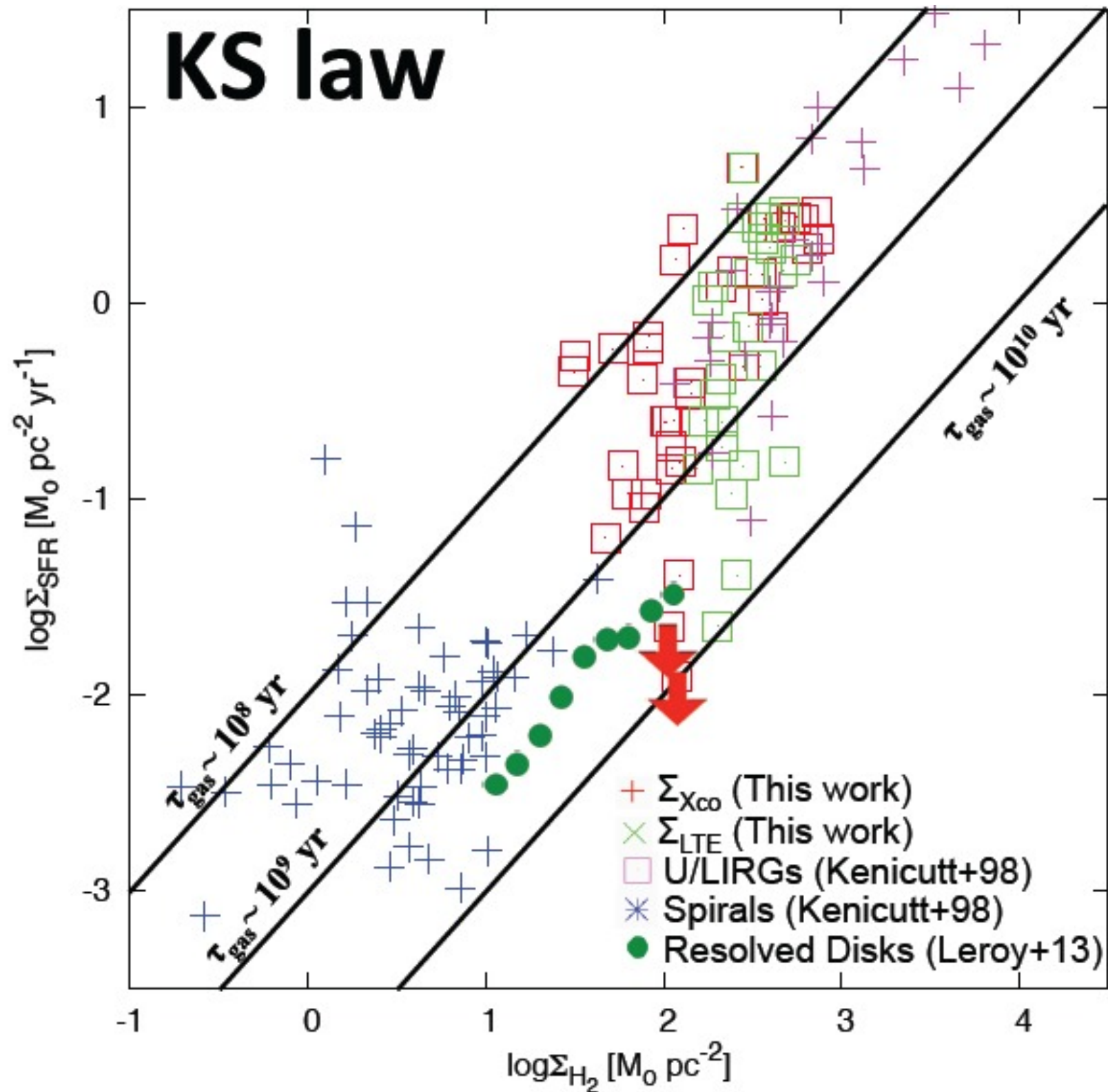
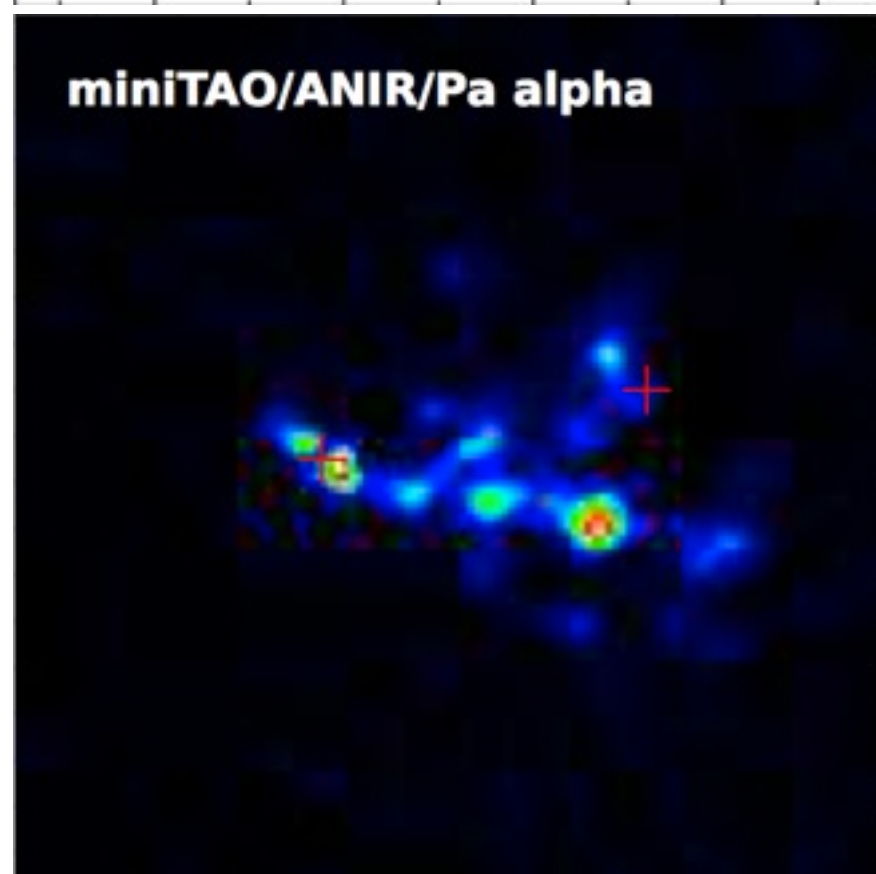
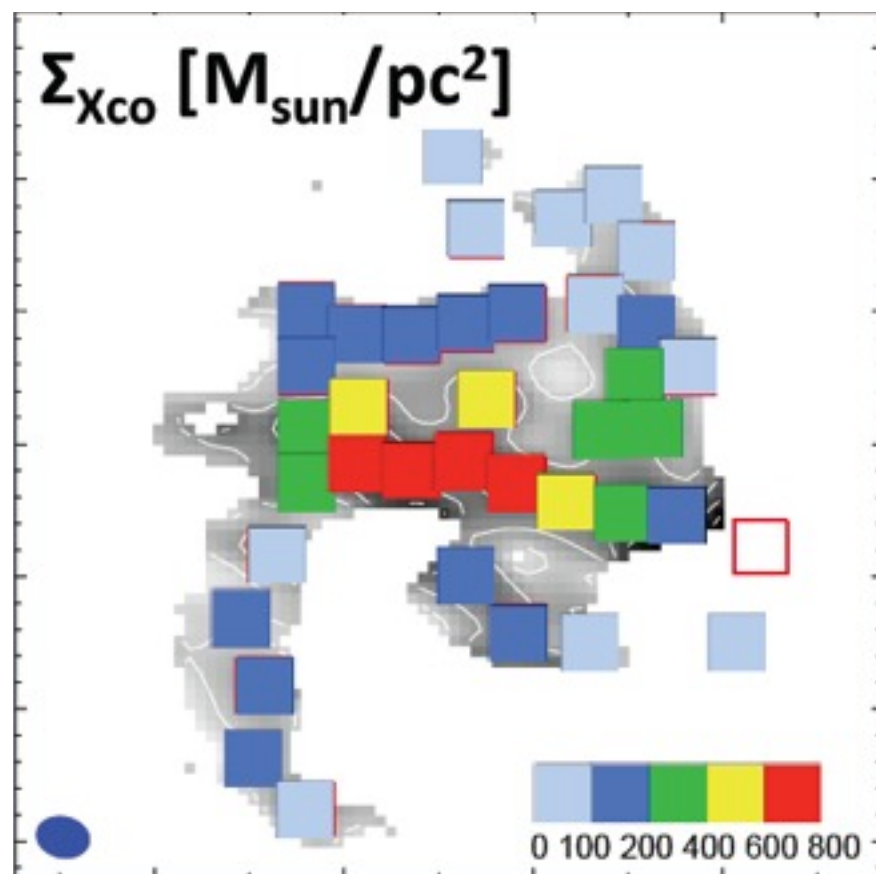


VV114: CO(1-0) & HCN(4-3)/HCO⁺(4-3)



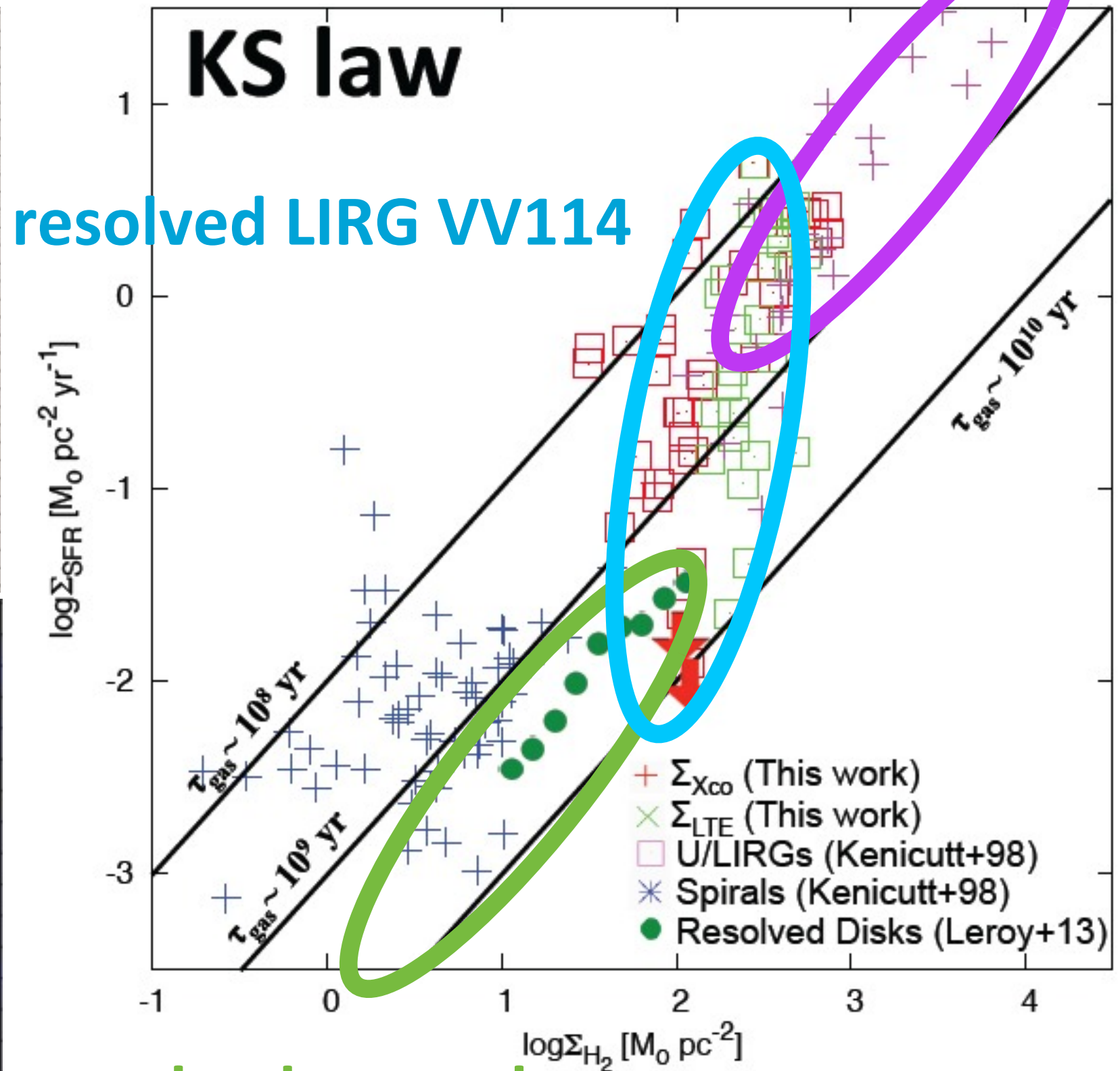
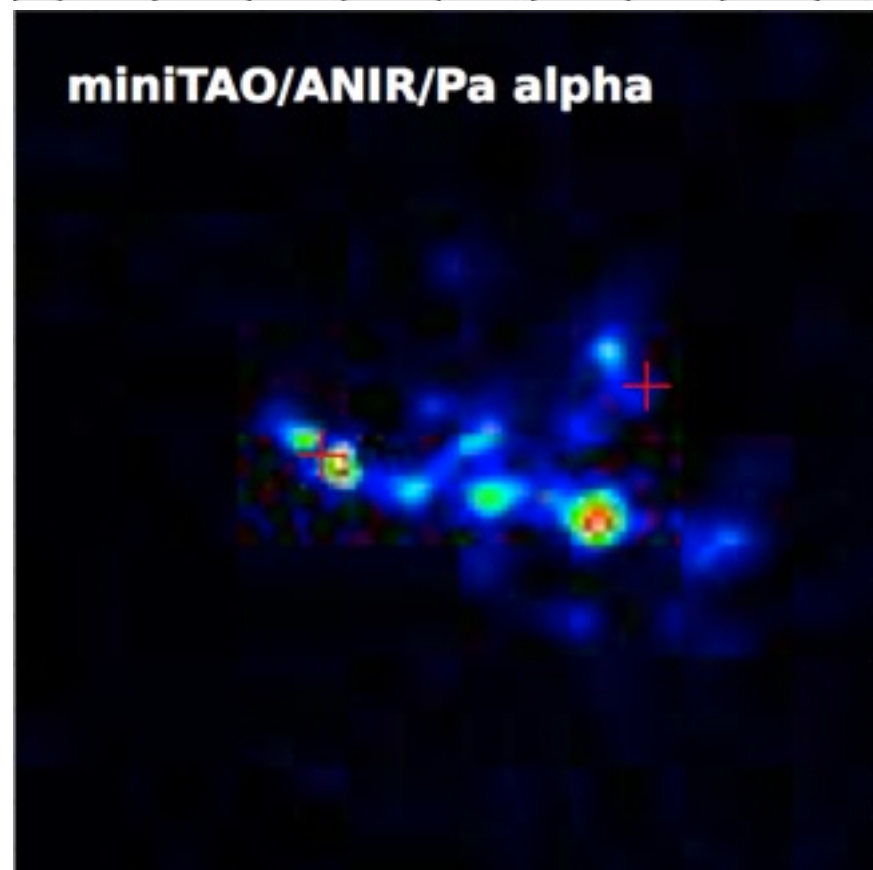
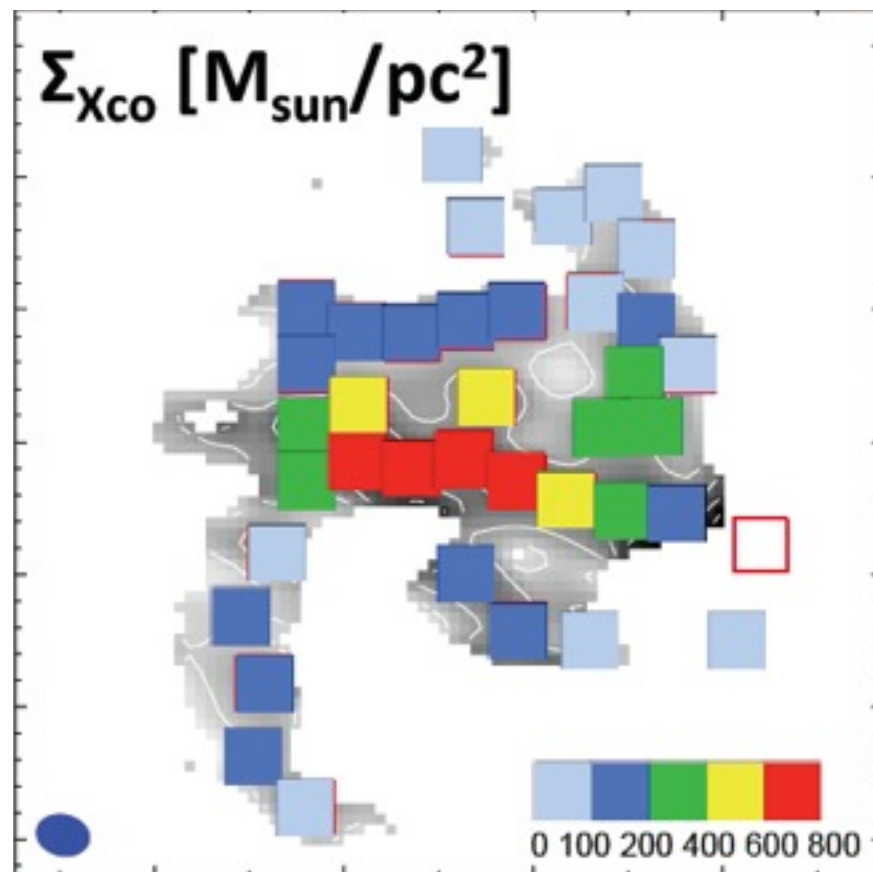
Iono et al. 2013

VV114: KS law (800 pc box)



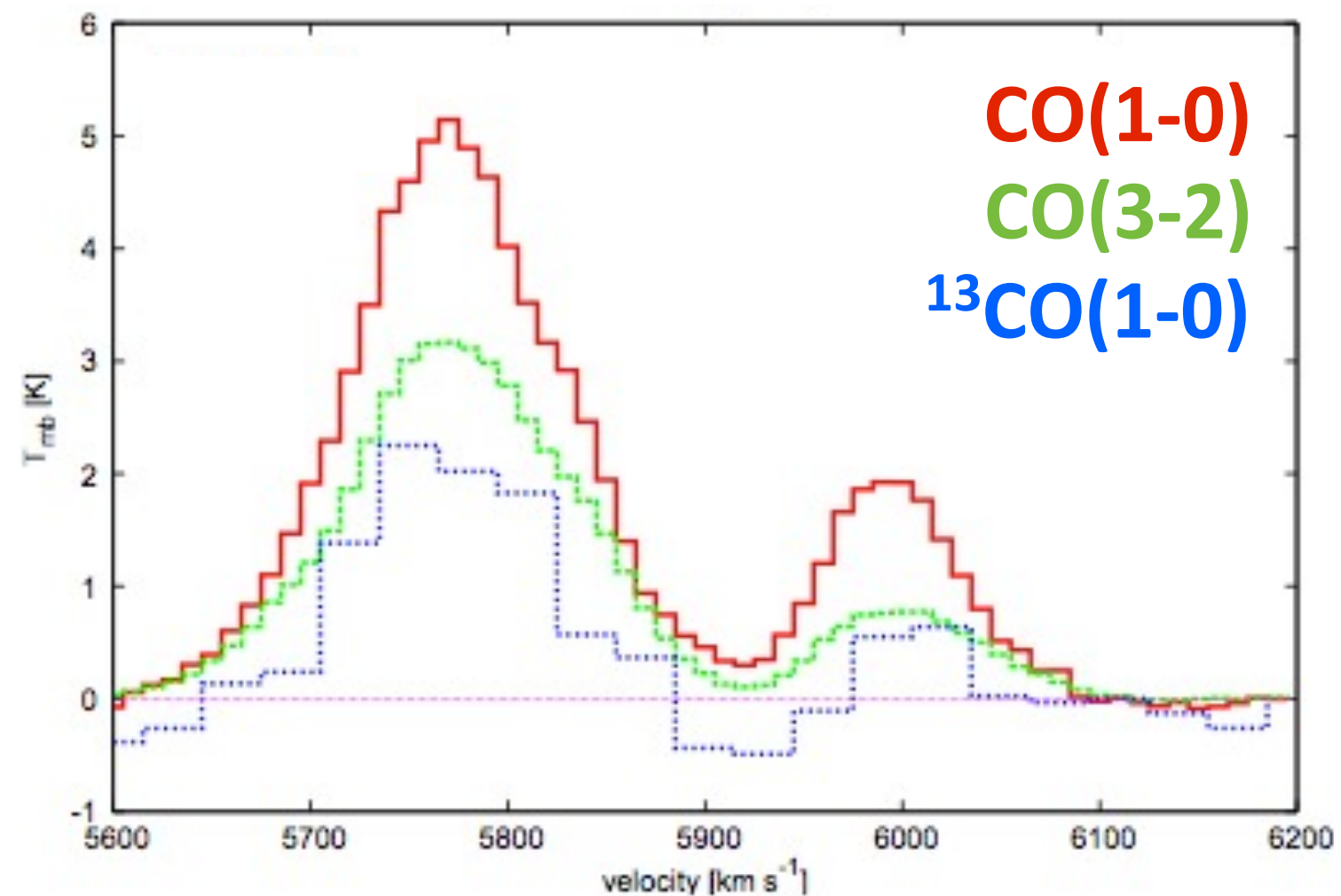
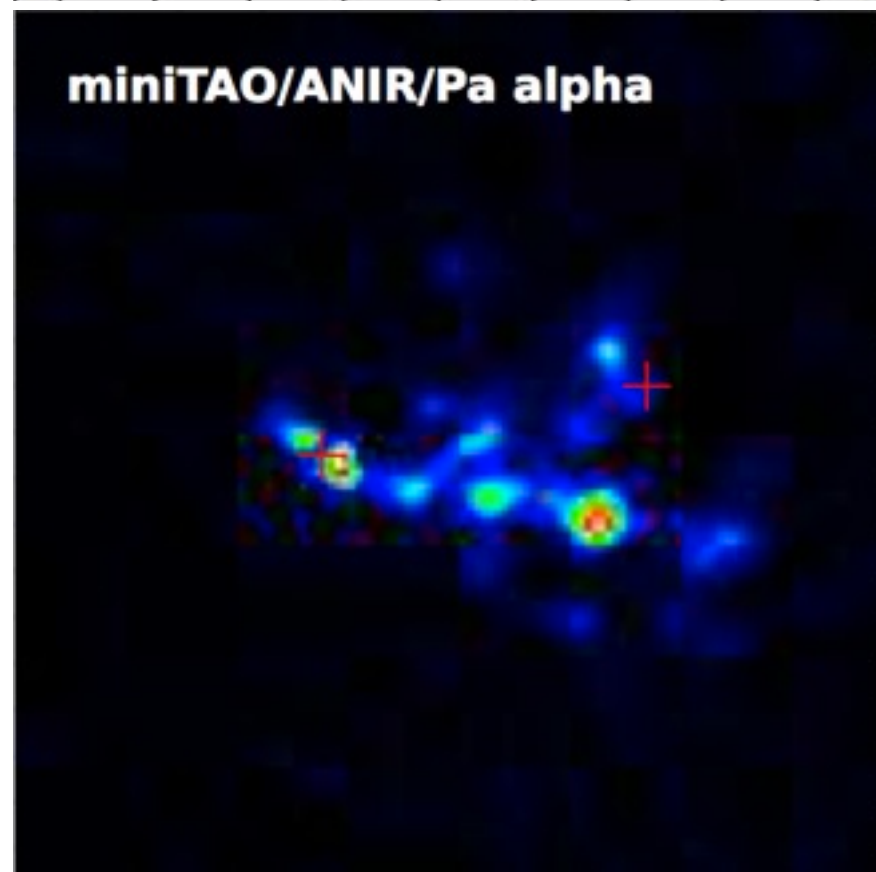
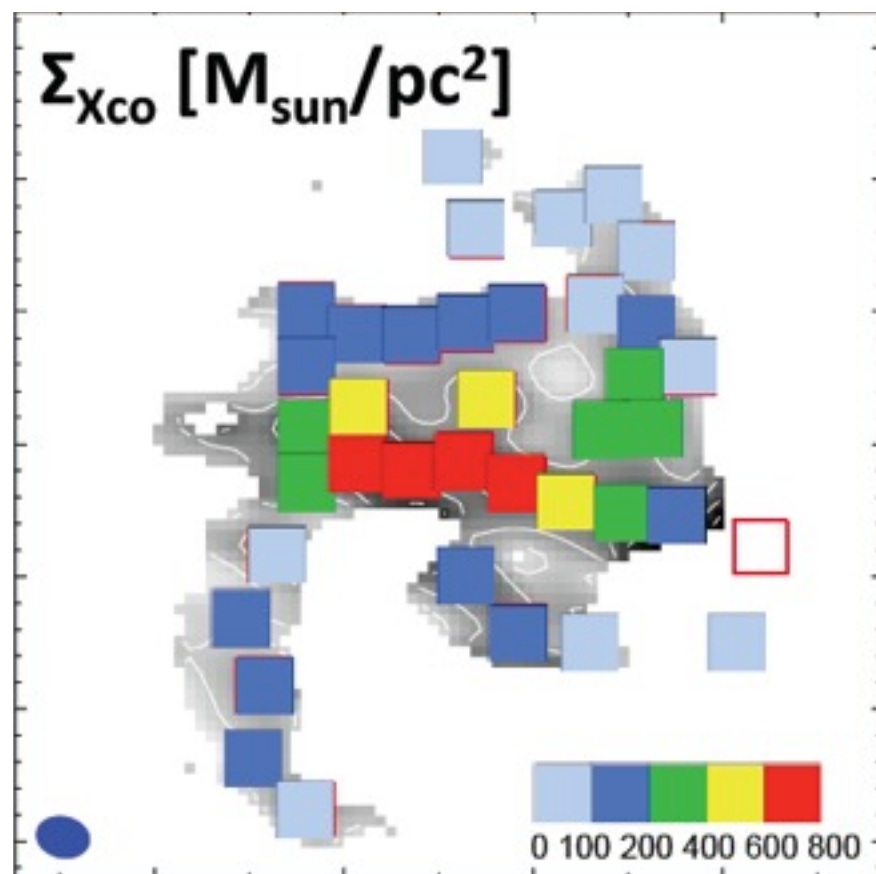
Saito, Iono et al. in prep.

VV114: KS law (2'' resolution)



Saito, Iono et al. in prep.

VV114: KS law (2'' resolution)

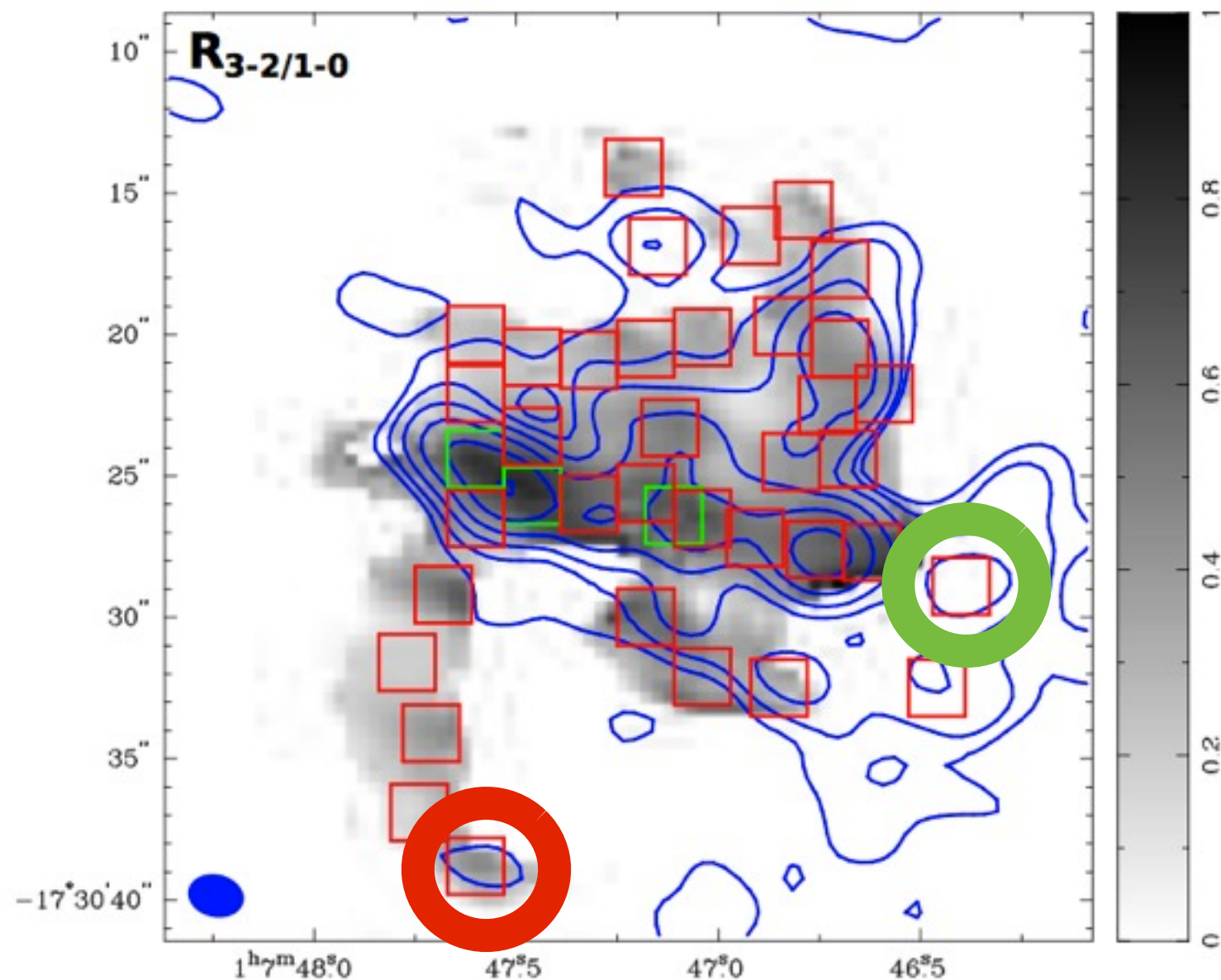
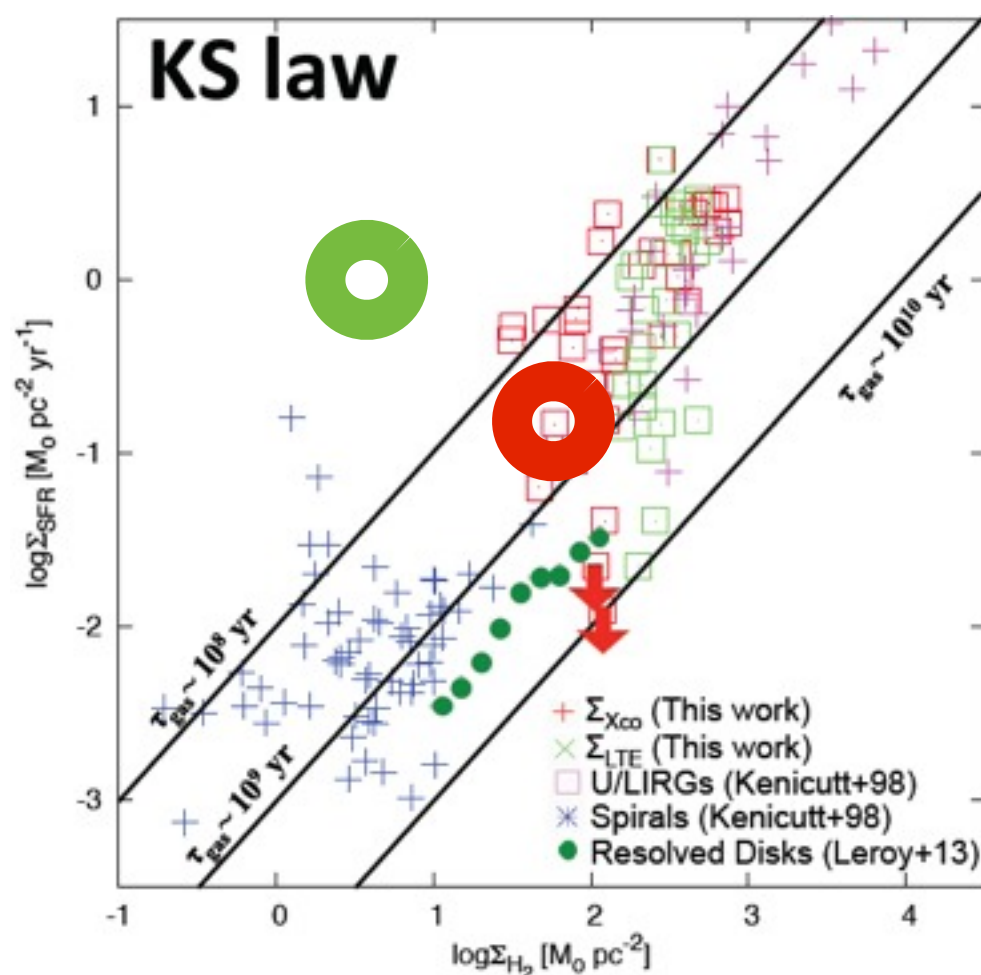


Which peak is associated with SF?

The late-stage mergers show very complicated kinematics.

Saito, Iono et al. in prep.

VV114: Specific Pa alpha peaks



-  Tidal dwarf galaxy formation at the edge of tidal tail?
-  molecular gas poor/exhausted SF region?

Saito, Iono et al. in prep.

Cold molecular gas and ionized gas in merger remnants

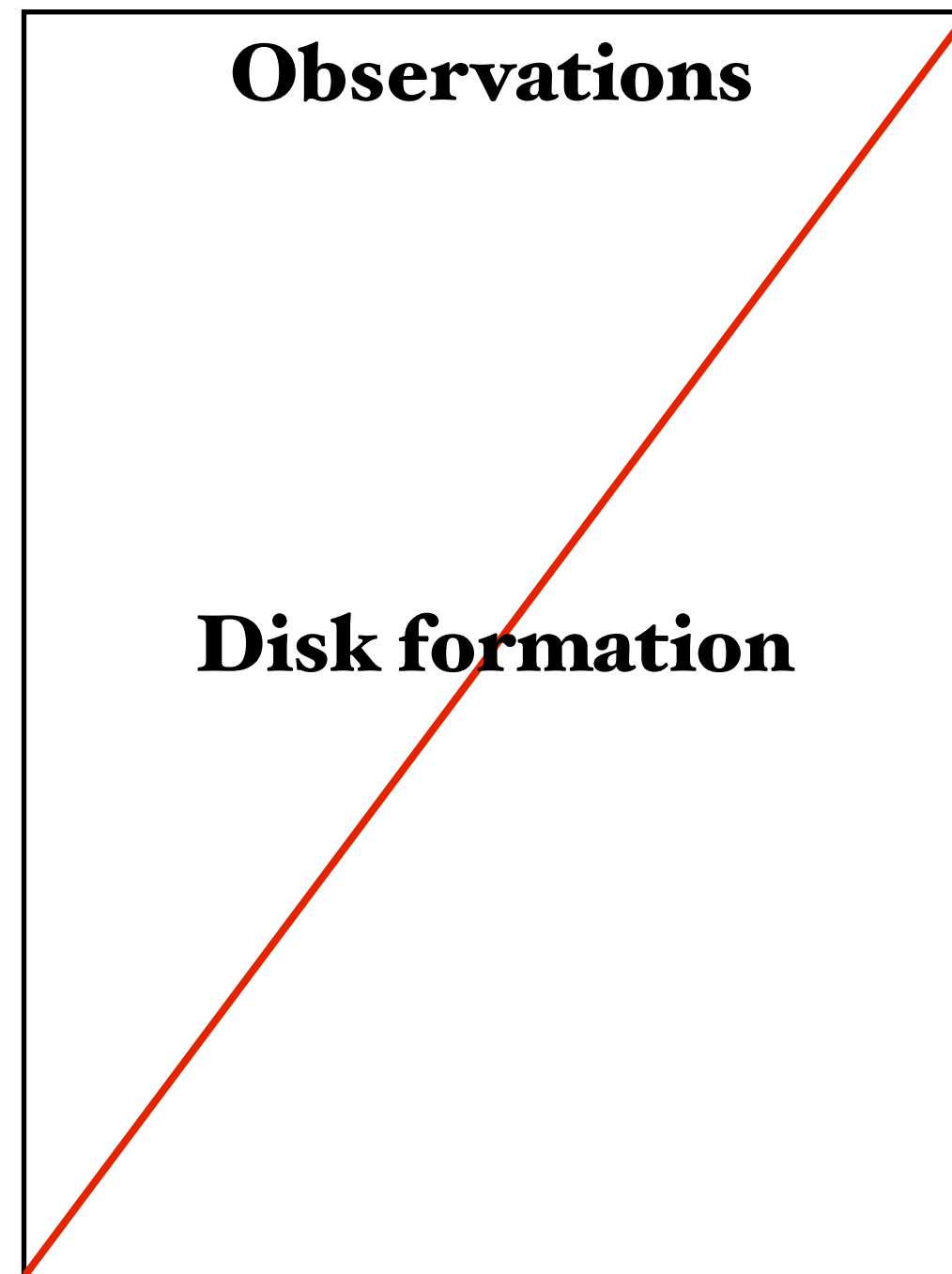
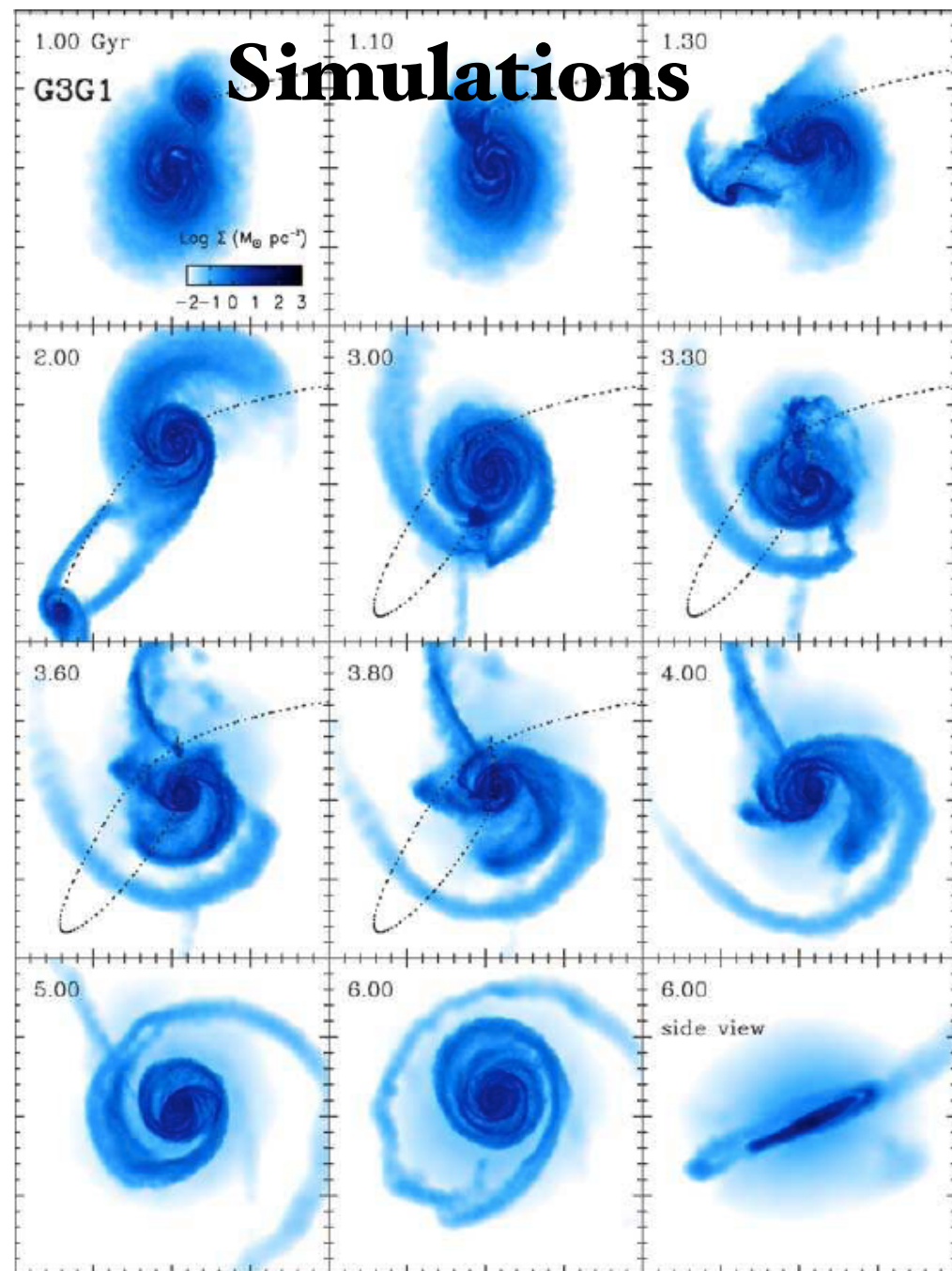


Previous Observations vs. Simulations



(Simulations) Toomre77, Barnes&Hernquist92, Naab&Burkert03
 (Observations) Schweizer&Seitzer92, Schweizer+96

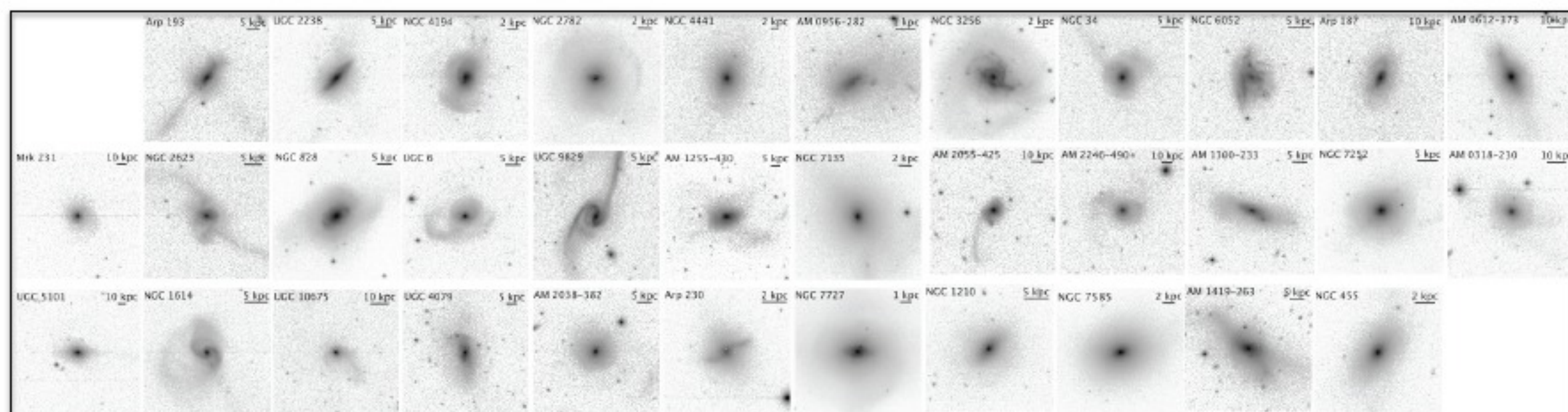
Recent Observations vs. Simulations



(Simulations) Cox+08, Barnes+02, Springel&Hernquist05
(Observations) Rothberg&Joseph04

Merger remnant sample

- 37 galaxies out of Rotheberg & Joseph (2004) catalog
- Rotheberg & Joseph (2004) is a catalog of 51 merger remnants compiled from 4 catalogs of peculiar galaxies (e.g., Arp, VV,..), and then selected based on K-band
 1. Optical morphology (tidal tails, loops)
 2. Single nucleus + No nearby companion



Summary

- I) **The late-stage LIRG VV114 has complicated kinematics.**
 - a. Which peak is associated with star-formation activity?
 - b. Does the TDG have same velocity as the tidal arm?
 - c. Why is there the Pa alpha peak without molecular gas at the western edge?
- 2) **CO of merger remnants show disk-like rotation.**
 - a. Is ionized gas kinematics consistent with the gas kinematics?
 - b. Do stars born in the gas disks?
 - e.g. counterrotating galaxies NGC5898, 7097 (Bettoni+84, Caldwell+86)

=> The SWIMS is important to develop our studies with ALMA.